

## → Deriving MHD

→ MHD is derived from 2-fluid equations

- first discuss 2 fluid derivation from Boltzmann
- then, discuss reduction to one-fluid MHD (i.e. approximations/limitations - especially in Ohm's Law)

## → deriving fluid equations

Have in general, Boltzmann eqn

$$\frac{\partial f}{\partial t} + \mathbf{v} \cdot \nabla f + \frac{q}{m} \left( \mathbf{E} + \frac{\mathbf{v} \times \mathbf{B}}{c} \right) \cdot \nabla f = c(f) \quad (4)$$

and can assign time scale

collision operator

① ↔  $\omega$  → frequency

② ↔  $v_{th}/L_{||}$

↳ relevant parallel scale

③  $\frac{2}{m} \frac{E}{\Delta V}$        $\Delta V \sim v_{Th} \rightarrow \text{non-resonant}$   
 $\Delta V \sim \Delta v_{Th} \rightarrow \text{resonant}$

$N_L \int$  scattering rate      ( $\rightarrow$  small, usually)

④  $\gamma_{\text{eff}}$  - collision frequency.

For "fluid description", need:

$\rightarrow \gamma_{\text{eff}} > v_{Th} / L_{\parallel}$

i.e. short mean free path limit

or

$\rightarrow \omega > v_{Th} / L_{\parallel}$        $\rightarrow$  all gyrokinetic KSAW,  
 where  $\gamma \rightarrow 0$

i.e. "fluid"

$\leftrightarrow$  blob / fluid element of particles.

$\rightarrow$  what holds blob together?  
 (i.e. prevents dispersal?)

$\Rightarrow$  collisions (i.e. particles collide and scatter prior dispersal)

or

$\Rightarrow$  vibrations in wave.

here, focus on short mean-free path ordering.

For  $C(f) \gg \partial f / \partial t, \underline{v} \cdot \nabla f$ , etc.

1.0.  $C(f) = 0$

$\Rightarrow f = f_{\text{Maxwellian}}$

local collisions drive distribution function to Maxwellian on time scale short compared all else

- n.b. Maxwellian can be shifted, and have gradients.

1st order:

$$\frac{\partial f^{(0)}}{\partial t} + \underline{v} \cdot \nabla f^{(0)} + \frac{q}{m} (\underline{E} + \frac{\underline{v}}{c} \times \underline{B}) \cdot \nabla f^{(0)} = C(f^{(1)})$$

then integrating:

$$\int d^3v \left[ \frac{\partial f^{(0)}}{\partial t} + \nabla \cdot \underline{v} f^{(0)} + \frac{q}{m} \left( \underline{E} + \frac{\underline{v}}{c} \times \underline{B} \right) \cdot \nabla f^{(0)} \right] = \int d^3v C(f^{(1)})$$

*IBP*  
 $\nearrow$   
 $\Delta$  0

Now,  $\int d^3v C(f.) = 0 \rightarrow$  collisions conserve #/a

so, have:

$$\frac{\partial n}{\partial t} + \nabla \cdot (n \underline{v}) = 0$$

c.e. continuity equation

$$\left\{ \begin{array}{l} n \equiv \int d^3v f \\ \underline{v} \equiv \int d^3v \underline{v} f / n \end{array} \right.$$

$\rightarrow$  basic moments.

→ Now first order moment:

$$\int d^3v \underline{v} \left( m \frac{\partial f}{\partial t} + \underline{v} \cdot \nabla f m + \underline{e} (\underline{E} + \underline{v} \times \underline{B}) \cdot \nabla f = 0 \right) \quad (4)$$

$$(1) = m \frac{\partial (\underline{v} f)}{\partial t}$$

$$\underline{v} = \underline{v}(x, t)$$

$$(3) = \int \underline{v} q (\underline{E} + \underline{v} \times \underline{B}) \cdot \frac{\partial f}{\partial \underline{v}}$$

$$= \int \frac{\partial}{\partial \underline{v}} [f \underline{v} (\underline{E} + \underline{v} \times \underline{B})] d^3v - \int f \underline{v} \frac{\partial}{\partial \underline{v}} \cdot (\underline{E} + \underline{v} \times \underline{B})$$

$$- \int f (\underline{E} + \underline{v} \times \underline{B}) \cdot \frac{\partial \underline{v}}{\partial \underline{v}}$$

is

$$= -\underline{e} n (\underline{E} + \underline{v} \times \underline{B})$$

$$(4) = \int d^3v m c(f) \underline{v}$$

$$= \underline{P}_{ij}$$

→ collisional momentum transfer from species i to j

which leaves ②:

$$\begin{aligned}
 \textcircled{2} &= m \int d^3v \underline{v} (\underline{v} \cdot \underline{\nabla}) F \\
 &= m \int d^3v \underline{\nabla} \cdot (F \underline{v} \underline{v}) \\
 &= \underline{\nabla} \cdot \left[ m \int d^3v F \underline{v} \underline{v} \right] = m \underline{\nabla} \cdot (n \overline{\underline{v} \underline{v}})
 \end{aligned}$$

clearly useful to separate  $\underline{v}$  into mean and fluctuating pieces

$$\underline{v} = \underline{V} + \underline{w}$$

$$\begin{aligned}
 \Rightarrow \underline{\nabla} \cdot (n \overline{\underline{v} \underline{v}}) &= \underline{\nabla} \cdot (n \underline{V} \underline{V}) + \underline{\nabla} \cdot (n \overline{\underline{w} \underline{w}}) \\
 &\quad + \underline{\nabla} \cdot n (\underline{V} \overline{\underline{w}} + \overline{\underline{w}} \underline{V})
 \end{aligned}$$

$\swarrow$   
 $\varnothing$ , defn.

$$\underline{\nabla} \cdot (n \underline{V} \underline{V}) = \underline{V} \cdot \underline{\nabla} \cdot (n \underline{V}) + n (\underline{V} \cdot \underline{\nabla}) \underline{V}$$

$$n \overline{\underline{w} \underline{w}} \equiv \underline{\rho}$$

$\downarrow$   
pressure tensor

so, can write for momentum equation

$$m \frac{\partial}{\partial t} (n \underline{V}) + m \underline{V} \cdot \nabla (n \underline{V}) + mn (\underline{V} \cdot \nabla) \underline{V} + \nabla \cdot \underline{P} - qn (\underline{E} + \underline{V} \times \underline{B}) = \underline{R}_j$$

and using continuity :

$$mn \left[ \frac{\partial}{\partial t} \underline{V} + \underline{V} \cdot \nabla \underline{V} \right] = qn (\underline{E} + \underline{V} \times \underline{B}) - \nabla \cdot \underline{P} + \underline{R}_j$$

Now, for form P :

$$\underline{P} = \int d^3V m v_i v_j f$$

in short mean-free-path ordering,

$$\underline{P} \approx \underline{P}_{\text{Maxwellian}}$$

As mean extracted, symmetry  $\Rightarrow$

$$\underline{P} = \int d^3V v_i v_j d_{ij} f$$

is  $\underline{\underline{P}} = \begin{bmatrix} P_1 & 0 & 0 \\ 0 & P_2 & 0 \\ 0 & 0 & P_3 \end{bmatrix}$

pressure tensor diagonal

if isotropic:  $P_1 = P_2 = P_3$

(fast  $\parallel, \perp$  thermal equilibration)

$\Rightarrow \underline{\underline{P}} = p \underline{\underline{I}}$

and pressure reduces to scalar, i.e.

$$m n \left[ \frac{\partial \underline{V}}{\partial t} + \underline{V} \cdot \nabla \underline{V} \right] = q n (\underline{E} + \underline{V} \times \underline{B}) - \nabla p + \underline{A}_{ij}$$

$\rightarrow$  For second order moment  $\rightarrow$  energy  
 (closure  $\leftrightarrow$  energy flux)  $\Rightarrow$  orn. state

2 species  $\Rightarrow p/p_0 = \text{const.}$



→ Single Fluid (→ MHD)

Can define single fluid variables:

$$\rho = n_i M + n_e m \approx n M \quad \rightarrow \text{density}$$

mass velocity:

$$\underline{v} = \frac{1}{\rho} (n_i M \underline{v}_i + n_e m_e \underline{v}_e) \quad \underline{\text{mean velocity}}$$

$$\approx \left[ \frac{M \underline{v}_i + m_e \underline{v}_e}{M + m} \right] \approx \underline{v}_i$$

Current density:

relative velocity

$$\underline{J} = q (n_i \underline{v}_i - n_e \underline{v}_e)$$

$$\approx n q (\underline{v}_i - \underline{v}_e) \quad , \text{ using } qN.$$

Upside:

- Continuity for ions  $\Rightarrow$  single fluid continuity

$$\therefore \boxed{\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \underline{V}) = 0}$$

- adding electron and ion momentum eqns:

$$M n \left( \frac{\partial \underline{V}_i}{\partial t} + \underline{V}_i \cdot \nabla \underline{V}_i \right) = z n \left( \underline{E} + \underline{V}_i \times \underline{B} \right) - \nabla \rho_i + \rho_{i,e}$$

$$m_e n \left( \frac{\partial \underline{V}_e}{\partial t} + \underline{V}_e \cdot \nabla \underline{V}_e \right) = -z n \left( \underline{E} + \underline{V}_e \times \underline{B} \right) - \nabla \rho_e + \rho_{e,i}$$

$\Rightarrow$

$$n \left( \frac{\partial}{\partial t} (M \underline{V}_i + m_e \underline{V}_e) + M (\underline{V}_i \cdot \nabla) \underline{V}_i + m_e (\underline{V}_e \cdot \nabla) \underline{V}_e \right) = z n (\underline{V}_i - \underline{V}_e) \times \underline{B} - \nabla (\rho_i + \rho_e) + \cancel{\rho_{e,i} + \rho_{i,e}}$$

momentum cons.

as:  $m_e \ll M$   
 $\underline{J}$  defn.  
 $\rho = \rho_e + \rho_i$

$$\Rightarrow \rho \left( \frac{\partial \underline{v}}{\partial t} + \underline{v} \cdot \nabla \underline{v} \right) = \underline{j} \times \underline{B} - \nabla \rho + F_{\text{body}}$$

↓  
any additional body force.

Momentum balance.

→ Now, only re-maining non-trivial MHD equation is Ohm's Law.

→ Where the bodies are buried, ...

Consider,  $[ m_e * (\text{ion momentum eqn}) -$   
 $M * (\text{electron momentum eqn.}) ]$

$$\Rightarrow M m_e n \left( \frac{\partial}{\partial t} (\underline{v}_i - \underline{v}_e) + \underline{v}_i \cdot \nabla \underline{v}_i - \underline{v}_e \cdot \nabla \underline{v}_e \right)$$

$$= z n (M + m_e) \underline{E} + z n (m \underline{v}_i + M \underline{v}_e) \times \underline{B}$$

$$- m \nabla \rho_i + M \nabla \rho_e - (M + m) \rho_{e,i}$$

Now, ①  $\underline{P}_{ei}$  = electron-ion momentum transfer

$$= -M n_e \mu \underline{J}$$

②  $M \gg m_e$

③ neglecting advective derivatives

$\Rightarrow$

$$\frac{M m_e n}{Z} \frac{\partial}{\partial t} \left( \frac{\underline{J}}{n} \right) = Z \rho \underline{E} - M n_e \mu \underline{J} + M \nabla \rho_e + Z n (m \underline{v}_i + M \underline{v}_e) \times \underline{B}$$

and can further simplify:

$$m \underline{v}_i + M \underline{v}_e = M \underline{v}_i + m \underline{v}_e - (M - m) (\underline{v}_i - \underline{v}_e) \approx \frac{\rho \underline{v}}{n} - \frac{M}{n_e} \underline{J}$$

Finally, re-arranging  $\Rightarrow$

$$\frac{m_e}{n_e^2} \frac{\partial \underline{J}}{\partial t} = \left( \underline{E} + \underline{v} \times \underline{B} \right) - \frac{M}{n_e} \underline{J} - \frac{Z}{n_e} (\underline{J} \times \underline{B}) + \frac{Z}{n_e} \nabla \rho_e$$

Now, have generalized Ohm's Law:

$$\frac{m_e}{nq^2} \frac{d\vec{J}}{dt} = \left( \vec{E} + \vec{v} \times \vec{B} \right) - \mu \vec{J} - \frac{(\vec{J} \times \vec{B})}{nq} + \frac{\nabla p_e}{nq}$$

② → ideal MHD Ohm's Law

③ → collisional resistivity

resistive  
MHD

bring in ④ : Hall Term

⇒ Hall MHD

bring in ⑤ : Electron thermal force / pressure

⇒ diamagnetic / finite electron  $w_e$   
MHD

i.e. Boltzmann response :  $\vec{E}$  vs  $\frac{\nabla p_e}{nq}$

① : Electron inertia term ( $\sim m_e$ )

⇒ EMHD, electron inertially modified  
MHD.  
( $\omega m_e / nq^2 > 1$ )

For low frequency, strong collisionality, etc.

$$\Rightarrow \underline{\underline{E}} + \underline{v} \times \underline{B} = \mu_0 \underline{J} \quad \left. \vphantom{\underline{J}} \right\} \text{Resistive MHD.}$$

N.B. :- Ohm's Law is most sensitive part of MHD structure  $\rightarrow$  need care.

- high  $\omega \rightarrow$  electron inertia
- tok.  $\mu$ -inst  $\rightarrow$  thermal force term.
- $\lambda \sim c^2 / \omega_{pe}^2 \rightarrow$  Hall term.

→ Reduced MHD

Note: ① full MHD : 3  $\underline{v}$  components  
2  $\underline{B}$  " " ( $\nabla \cdot \underline{B} = 0$ )  
 $\rho$   $\rho$

⇒ 7 components

② if  $\nabla \cdot \underline{v} = 0$  ⇒ 4 components  
( $\rho = \text{const}$ ,  $\rho$  from  $\nabla \cdot \underline{v} = 0$ )

③ strongly magnetized system ⇒ Reduced MHD  
⇒ scalar equations for  $\phi, \psi$  (2 scalar fields)

Now:

- assume strong  $\underline{B}_z$  (strong magnetization → gyrokinetics)

"strong" ⇔  $\rho v^2 \sim \rho \ll B_z^2 / 8\pi$

so motion strongly anisotropic, and small scales generated in  $\perp$  direction only, as strong  $B_z$  inhibits line bending, (energy to perturb strong, high energy density field).

⇒ order :  $B_z \sim v_{\perp} \sim 1$

$B_{\parallel} \sim \alpha_z \sim O(\epsilon)$

Take  $\rho \sim 1$ , as  $\nabla \cdot \underline{v} = 0$  enforced by strong  $B_z$ .

$v_{\perp}^2 \sim \rho \sim B_{\perp}^2$  (i.e. equipartition of energy) (springiness)

$\Rightarrow v_{\perp} \sim \epsilon, \rho \sim \epsilon^2, \partial_t \sim \underline{v}_{\perp} \cdot \nabla_{\perp} \sim \epsilon$

and pressure balance ( $\nabla \cdot \underline{v} = 0$  and incompressibility)

$\delta(B_z^2) \sim 2B_z \delta(B_z) \sim \rho$  (eqbm)

$\Rightarrow \delta B_z \sim \epsilon^2$

" to lowest order  $\Rightarrow B_z = \text{const}$ ,

Now then:

$(\nabla \cdot \underline{B} = 0)$

$$\underline{B} = \hat{z} \times \nabla \psi + B_z \hat{z}$$

$$= \nabla A_{\parallel} \times \hat{z} + B_z \hat{z}$$

$B$  rep. by single scalar potential

$\psi = -A_{\parallel}$

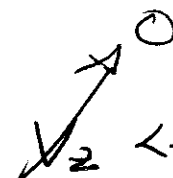
$\nabla \cdot \underline{B} = \partial_z B_z = \epsilon^3 \rightarrow 0$

parallel comp. of vector pot.

Similarly,

$\partial_z \rho \sim o(\epsilon^3)$   
 $\int_{\perp} B_{\perp} \sim \epsilon^3$

$\Rightarrow v_2 \ll v_1$   
 neglect  $v_2$ .





Now, 
$$\underline{E} = -\frac{1}{c} \frac{\partial \underline{A}}{\partial t} - \underline{\nabla} \phi = -\frac{\underline{v} \times \underline{B}}{c}$$

$$\Rightarrow +\frac{1}{c} \frac{\partial \underline{A}}{\partial t} = \frac{\underline{v} \times \underline{B}}{c} - \underline{\nabla} \phi \quad (*)$$

$$B_z = (\underline{\nabla} \times \underline{A}_\perp) \cdot \underline{z}$$

so  $\partial_t A_\perp \sim E^3$  (ala  $\partial_z \rho_z$ )

$\therefore \nabla_\perp \phi \approx \left( \frac{\underline{v} \times \underline{B}}{c} \right)_\perp$ , in  $(*)$

$$\Rightarrow \underline{v}_\perp = \frac{c \underline{z} \times \underline{\nabla} \phi}{B_z}$$

$\perp$  velocity  
 $\rightarrow$  motion  $\perp$  is  
 $\underline{E} \times \underline{B}$ .

Now, taking parallel component of  $(*)$ .  
 (units!)

$$\Rightarrow \frac{\partial \psi}{\partial t} + \underline{v} \cdot \underline{\nabla} \psi = \frac{B_z}{z} \partial_z \phi$$

(vector potential)  
 so have (flux) equation:

$$\frac{\partial \psi}{\partial t} + \underline{v} \cdot \underline{\nabla} \psi = B_z \partial_z \phi$$

$$= B_z \hat{z} + \hat{z} \times \nabla \psi \quad \underline{92}$$

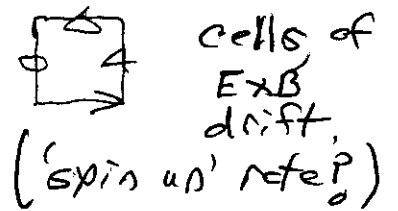
or, alternatively,

$$\frac{\partial \psi}{\partial t} - \underline{B} \cdot \nabla \phi = 0$$

Finally, for  $\phi$ , write:

$\perp$  motion

$$\frac{\partial \underline{v}}{\partial t} + \underline{v} \cdot \nabla \underline{v} = -\frac{\nabla p}{\rho_0} + \frac{\underline{J} \times \underline{B}}{c}$$



$(\nabla \times) \cdot \hat{z} \Rightarrow$  vorticity component ( $\parallel \hat{z}$ ) evolution

$$\begin{aligned} \frac{\partial \omega_z}{\partial t} + \underline{v}_\perp \cdot \nabla \omega_z &= -\cancel{\frac{\nabla \times \nabla p}{\rho_0}} + \hat{z} \cdot \nabla \times \left( \frac{\underline{J} \times \underline{B}}{c} \right) \\ &= \underline{B} \cdot \nabla J_z - \cancel{\underline{J} \cdot \nabla B_z} \quad \Delta B_z \sim \epsilon^3 \\ &\approx \underline{B} \cdot \nabla J_z \end{aligned}$$

$$\frac{\partial \omega_z}{\partial t} + \underline{v} \cdot \nabla \omega_z = \underline{B} \cdot \nabla J_z$$

but:

$$\begin{aligned} \omega_z &= \hat{z} \cdot \nabla \times \underline{v} = \nabla^2 \phi \\ J_z &= \hat{z} \cdot (\nabla \times \underline{B}) \frac{c}{4\pi} = \nabla^2 \psi \end{aligned}$$

so finally have:

$$\frac{\partial}{\partial t} \nabla^2 \phi + \underline{v} \cdot \underline{\nabla} \nabla^2 \phi = \beta_z \frac{\partial}{\partial z} \nabla^2 \psi + \underline{\tilde{B}} \cdot \underline{\nabla} \nabla^2 \psi$$

Finally, have reduced MHD equation:

$$\frac{\partial \psi}{\partial t} + \underline{v} \cdot \underline{\nabla} \psi = \beta_z \frac{\partial \phi}{\partial z} + \eta \nabla^2 \psi$$

$$\frac{\partial}{\partial t} \nabla^2 \phi + \underline{v} \cdot \underline{\nabla} \nabla^2 \phi - \nu \nabla^2 \nabla^2 \phi = \underline{\tilde{B}} \cdot \underline{\nabla} \nabla^2 \psi + \beta_z \frac{\partial}{\partial z} \nabla^2 \psi$$

- note have reduced MHD to 2 scalar evolution equations

- does this look familiar?

consider linearization in straight field,

$$-i\omega \hat{\psi} = ik_z \beta_z \hat{\phi}$$

$$+i\omega k_\perp^2 \hat{\phi} = -ik_z k_\perp^2 \psi$$

but recall for KSAW, with  $k_L^2 \rho_s^2 \rightarrow 0$

$$\hat{\phi} - \frac{\omega}{c} \frac{\hat{A}_{||}}{k_{||}} = 0 \quad - \text{Q.N. } \hat{n}/n_0 \text{ computed}$$

$$\nabla_{||} \hat{J}_{||}/n_0 e l = \partial_s^2 \frac{\partial}{\partial t} \left( \frac{\nabla_{\perp}^2 \hat{\phi}}{T_e} \right) - \text{adding GKE moments}$$

∴ clearly identical!

⇒ Reduced MHD equivalent to one-fluid theory based on gyro-kinetic equations!!

- indicates 2 routes to reduced MHD

Ⓐ Boltzmann Egn. → 2 Fluid Egn. → 1 fluid egn.

→ 'strong field' ordering → RMHD

n.b. 'strong field' ordering at macroscopic level

Ⓑ Boltzmann Egn. → Gyrokinetic Egn. → moment egn.

→ reduced MHD

n.b. 'strong field' ordering at microscopic level.

- for 2D MHD:

$$\frac{\partial \nabla^2 \phi}{\partial t} + \underline{v} \cdot \underline{\nabla} \nabla^2 \phi = \underline{B} \cdot \underline{\nabla} \nabla^2 \psi + \nu \nabla^2 \nabla^2 \phi$$

$$\frac{\partial \psi}{\partial t} + \underline{v} \cdot \underline{\nabla} \psi = \eta \nabla^2 \psi$$

- <sup>①</sup> Conservation Laws, etc. (HW)

$$\frac{d}{dt} E = 0 \quad (\text{to } \eta, \nu), \quad E = \int d^3x \left[ \frac{(\nabla \phi)^2}{2} + \frac{(\nabla \psi)^2}{2} \right]$$

$$\textcircled{2} \quad \mathcal{H} = \underline{A} \cdot \underline{B} \cong B_z \psi$$

↓  
const.

$$\Rightarrow \quad H = \int d^3x B_z \psi, \quad \frac{dH}{dt} = 0, \quad \text{to } o(\eta)$$

Ohm's Law (flux advection) is simple statement  
of helicity conservation: form  $\Gamma \psi$  s/t  $\begin{cases} H \text{ conserved} \\ EM \text{ dissipated} \end{cases}$

$$\textcircled{3} \quad K = \int d^3x \underline{v} \cdot \underline{B} = \int d^3x (\nabla \phi \cdot \nabla \psi)$$

also conserved, to dissipation.