

# MHD Stability and Energy Principle

## a) Basic Ideas

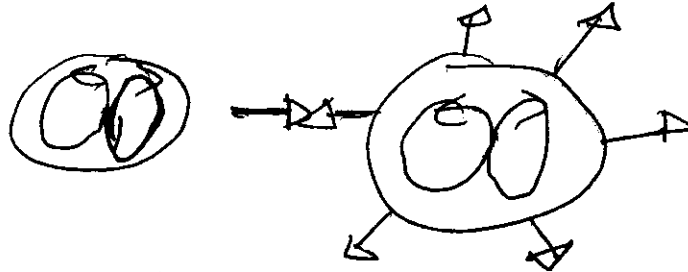
- up till now, concerned with localized instabilities, local theory, etc

c.i.e.  $\vec{\phi} \sim \vec{\phi}_H e^{i(k \cdot x - \omega t)}$ , etc.

but

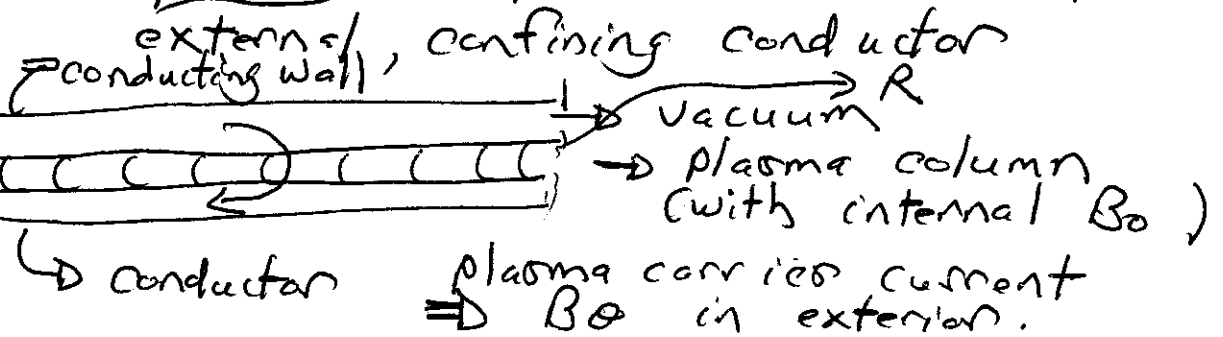
- to understand implications of stability for confinement, need consider effects of boundary conditions, etc.

c.i.e. recall: virial theorem  $\Rightarrow$  plasmoid cannot confine itself



∴ natural to consider simple confined plasma, the linear pinch - plasma column with

Linear pinch



N.B. Note effect of B.C.'s:

- plasma - vacuum interface to electrically, thermally isolate plasma, and allow small displacements
- conductor (i.e. wall) introduces electrical "no-slip" b.c.'s  
 i.e.  $E_{\text{tan}}|_{\text{Wall}} = 0$  ,  $B_n|_{\text{Wall}} = 0$   
 $\Rightarrow$  restrict perturbation.
- plasma - vacuum interface akin to air - water interface in waves.

idealizations:

- vacuum gap
- $B_{\theta z}$  in plasma (i.e.  $B_{\theta z}$  pervades gap, as well)
- $B_{\theta}$  external only ("skin current")  
 $\Rightarrow$  in reality,  $B_{\theta} = B_{\theta}(r)$  as  $J_z(r)$  distributed

still, linear pitch has been very useful idealized venue for stability studies...

$\Rightarrow$  Equilibrium

Now, for equilibrium:

$$\begin{aligned} \rho \frac{d\underline{v}}{dt} &= -\underline{\nabla} p + \underline{j} \times \underline{B} \\ &= -\underline{\nabla} \left( p + \frac{B^2}{8\pi} \right) + \frac{\underline{B} \cdot \underline{\nabla} \underline{B}}{4\pi} \end{aligned}$$

note:  $B_{\theta z}$  irrelevant to equilibrium!  $\rightarrow$  ignore.

then, noting  $J_z = \frac{1}{4\pi} \frac{1}{r} \frac{\partial}{\partial r} (r B_{\theta})$

$\Rightarrow$  have eqbm. condition:

$$\frac{dp}{dr} = -\frac{1}{4\pi} B_0 \frac{1}{r} \frac{\partial}{\partial r} (r B_0)$$

$r^2 \times$  above  $\Rightarrow$

$$\int_0^R dr r^2 \frac{dp}{dr} = -\frac{1}{4\pi} \int_0^R r^2 dr B_0 \frac{\partial}{\partial r} (r B_0)$$

assuming  $p(R) = 0$  (limit of confined pressure)

$\Rightarrow$

$$+2 \int_0^R dr r p(r) = +\frac{1}{8\pi} (r B_0(r))^2 \Big|_0^R$$

$$\Rightarrow \boxed{2 \int_0^R dr r p(r) = \frac{1}{8\pi} (R B_0(R))^2}$$

(isothermal)

Take  $T_i = T_e = T$   $n_e = n_i$   
 $\therefore p = 2n^2 k_B T$

$$N = \frac{\# \text{ particles}}{\text{length}} = \int_0^R dr 2\pi r n$$

$\downarrow$   
line density

and 
$$I = \frac{1}{2} R B_0(R) = \int dr \, 2\pi r J_z(r)$$

$\therefore 2Nk_b T = \frac{1}{2} I^2$

i.e. more current  $\Rightarrow$  more pressure held!

$\Rightarrow$

$$I = 2(Nk_b T)^{1/2}$$

Bennett  
criterion

$\rightarrow$  Stability

Now, convenient to think in terms of displacements, for linear theory:

i.e. perturbation  $\Rightarrow \underline{x} \rightarrow \underline{x} + \underline{\xi}(\underline{x}, t)$   
 $\downarrow$   
displacement

so 
$$\underline{\partial} \underline{v} = \frac{\partial}{\partial t} \underline{\xi}$$

$\downarrow$   
perturbed  
velocity

then have:

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \underline{v}) = 0$$

$$\Rightarrow \textcircled{1} \quad \partial_t \rho + \nabla \cdot (\rho \underline{v}) = 0$$

$$\rho \rho^{-\gamma} = \text{const.}$$

$$\Rightarrow \frac{d}{dt} \rho \rho^{-\gamma} = 0, \quad \frac{d}{dt} = \frac{\partial}{\partial t} + \underline{v} \cdot \nabla$$

$$\rho^{-\gamma} \frac{\partial \rho}{\partial t} + (-\gamma) \rho^{-\gamma-1} \rho \frac{\partial \rho}{\partial t} + \left[ \left( \frac{\partial \underline{v}}{\partial t} \cdot \nabla (\rho \rho^{-\gamma}) \right) \right] = 0$$

$$\Rightarrow \partial_t \rho + \underline{v} \cdot \nabla \rho = \gamma \rho (\partial_t \rho + \underline{v} \cdot \nabla \rho)$$

and using ①

$$\textcircled{2} \quad \partial_t \rho = -\gamma \rho \nabla \cdot \underline{v} - \underline{v} \cdot \nabla \rho$$

$$\Rightarrow \frac{\partial}{\partial t} \partial_t \underline{B} = \nabla \times \underline{\tilde{v}} \times \underline{B}$$

$$= \nabla \times \frac{\partial \underline{v}}{\partial t} \times \underline{B}$$

$$\therefore \textcircled{3} \quad \partial_t \underline{B} = \nabla \times (\underline{v} \times \underline{B})$$

and similarly:

$$\textcircled{4} \quad \underline{\omega} = \frac{1}{4\pi} \underline{\nabla} \times [\underline{\nabla} \times (\underline{\varepsilon} \times \underline{B})]$$

so

$$\rho_0 \frac{\partial \underline{V}}{\partial t} = -\underline{\nabla} \tilde{\rho} + \underline{\sigma} \times \underline{B}$$

$$\Rightarrow \rho_0 \frac{\partial \underline{E}}{\partial t} = \underline{F}(\underline{E}) \quad \text{where}$$

$$\underline{F} = \underline{\nabla} \left[ \gamma \rho \underline{\nabla} \cdot \underline{\varepsilon} + \underline{\varepsilon} \cdot \underline{\nabla} \rho \right] + \frac{1}{4\pi} \left\{ \left[ \underline{\nabla} \times \underline{\nabla} \times (\underline{\varepsilon} \times \underline{B}) \right] \times \underline{B} + (\underline{\nabla} \times \underline{B}) \times \left[ \underline{\nabla} \times (\underline{\varepsilon} \times \underline{B}) \right] \right\}$$

with boundary conditions restricting  $\omega^2$

i.e.

$$-\omega^2 \rho_0 \underline{E} = \underline{F}(\underline{E}) + \text{B.C.'s}$$

s/t  $\omega^2$  quantized, so  $\omega^2 \rightarrow \omega_n^2$   
 $\underline{E} \rightarrow \underline{E}_n$   
 eigenfunctions, etc.

N.B. : Will show  $\underline{F(\omega)}$  hermitian  
 $\Rightarrow \omega^2$  real

in MHD have stable oscillations  
 $(\Rightarrow \omega^2 > 0)$  or purely growing modes  $(\omega^2 < 0)$ .

Now

$\rightarrow$  For boundary conditions: [plasma-vacuum interface]

refer internal (to plasma) perturbations  
 with  $i$

refer external with  $e$

So, for external fields: (vacuum region)

$$\underline{\delta E}^e = -\frac{1}{c} \frac{\partial}{\partial t} \underline{\delta A}^e \quad (\text{induced only})$$

$$\underline{\delta B}^e = \nabla \times (\underline{\delta A}^e).$$

and require tangential electric field  
 continuous at interface.

i.e. if  $\hat{n} \equiv$  unit normal to plasma,

then :



$$\hat{n} \times \left[ \underline{\underline{\partial E}}^i + \frac{1}{c} \underline{\underline{\partial V}} \times \underline{\underline{B}}^i \right]$$

$$= \hat{n} \times \left[ \underline{\underline{\partial E}}^e + \frac{1}{c} \underline{\underline{\partial V}} \times \underline{\underline{B}}^e \right]$$

but ideal MHD  $\Rightarrow \underline{\underline{\partial E}}^i + \frac{\underline{\underline{\partial V}}}{c} \times \underline{\underline{B}}^i = 0$

$$\Rightarrow \hat{n} \times \left[ -\frac{1}{c} \frac{\partial}{\partial t} \underline{\underline{\partial A}}^{(e)} + \frac{1}{c} \frac{\partial \underline{\underline{\epsilon}}}{\partial t} \times \underline{\underline{B}}^e \right] = 0$$

$$\hat{n} \times \frac{\partial}{\partial t} \underline{\underline{\partial A}}^{(e)} = - \left( \hat{n} \cdot \frac{\partial \underline{\underline{\epsilon}}}{\partial t} \right) \underline{\underline{B}}^e - (\hat{n} \cdot \underline{\underline{B}}^e) \frac{\partial \underline{\underline{\epsilon}}}{\partial t}$$

but here (and anything sensible),

$$\hat{n} \cdot \underline{\underline{B}}_0 = 0$$

$$\Rightarrow \boxed{\hat{n} \times \underline{\underline{\partial A}}^{(e)} = -(\hat{n} \cdot \underline{\underline{\epsilon}}) \underline{\underline{B}}^e} \quad \left[ \begin{array}{l} \text{electromag.} \\ \text{B.C. at} \\ \text{interface} \end{array} \right]$$

Now, at conducting wall

$$\hat{n} \times \underline{\underline{\partial A}}^{(e)} = 0 \quad (E_{\text{tan}} = 0)$$

$$\hat{n} \cdot (\nabla \times \underline{\underline{\partial A}}^{(e)}) = 0 \quad (B_n = 0)$$

and, in vacuum:

$$\left\{ \frac{1}{c^2} \frac{\partial^2}{\partial t^2} \underline{A}^{(e)} + \nabla \times (\nabla \times \underline{A}^{(e)}) = 0 \right.$$

Finally, must have pressure balance at interface:

$$\rho'_{tot}(\underline{x} + \underline{\varepsilon}) = \rho^e(\underline{x} + \underline{\varepsilon})$$

$$\rho_{tot} \equiv \rho + \frac{B^2}{8T}$$

so, to first order:

$$(1 + \underline{\varepsilon} \cdot \underline{\nabla}) \left[ \rho + \frac{B'^2}{8T} + \delta\rho + \frac{1}{4T} \underline{B}' \cdot \delta \underline{B}' \right]$$

$$= (1 + \underline{\varepsilon} \cdot \underline{\nabla}) \left[ \frac{(B^e)^2}{8T} + \frac{1}{4T} \underline{B}^e \cdot \delta \underline{B}^e \right]$$

and, using (2) for  $\delta\rho \Rightarrow$

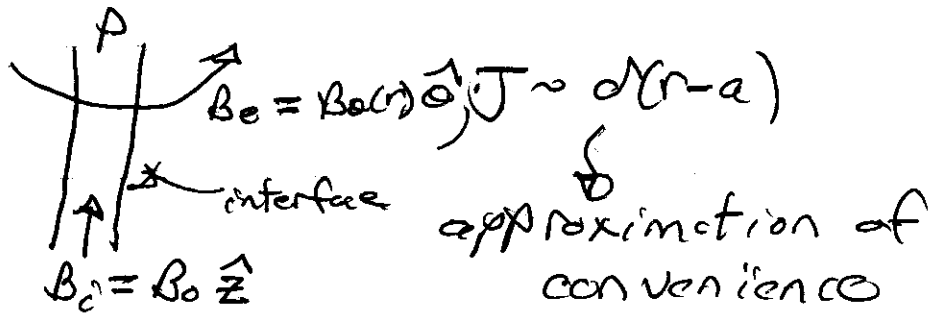
$$\left\{ -\gamma \rho \underline{\nabla} \cdot \underline{\varepsilon} + \frac{1}{4T} \underline{B}' \cdot [\delta \underline{B}' + \underline{\varepsilon} \cdot \underline{\nabla} \underline{B}'] \right.$$

$$\left. = \frac{1}{4T} \underline{B}^e \cdot [\delta \underline{B}^e + \underline{\varepsilon} \cdot \underline{\nabla} (B^e)] \right.$$

# Sausage and Kink Modes

more MHD stability  $\rightarrow$  'Links and kinks'

Consider surface current sheet pinch:



- As  $dp/dr = -\frac{1}{4\pi} \frac{B_0}{r} \frac{d}{dr} (r B_0)$  pressure to far field elev. to confinement

profile is flat inside ( $\nabla (B_0^2 / 4\pi) = 0$ ) (ideal MHD)

- recall derived:  $\rightarrow$  generalized displacement

$$\rho_0 \frac{d^2 \underline{\epsilon}}{dt^2} = F(\underline{\epsilon})$$

$\hat{F} \rightarrow$  force (hermitian operator)

$$= \nabla \cdot [\gamma \rho \nabla \cdot \underline{\epsilon} + \underline{\epsilon} \cdot \nabla \rho]$$

$$+ \frac{1}{4\pi} \left\{ [\nabla \times \nabla \times (\underline{\epsilon} \times \underline{B})] \times \underline{B} + (\nabla \times \underline{B}) \times [\nabla \times (\underline{\epsilon} \times \underline{B})] \right\}$$

and  $\hat{n} \times \underline{dA}^e = -(\hat{n} \cdot \underline{\epsilon}) \underline{B}^e$  at interface

$$\hat{n} \times \underline{dA}^e = \hat{n} \cdot (\nabla \times \underline{dA}^e) = 0$$

at wall

$\hat{n} =$  unit normal to plasma

relates external e.m. to displacement

no displ., MHD

in gap:  $\frac{1}{c^2} \frac{d^2 A^e}{dt^2} + \nabla \times (\nabla \times dA^e) = 0$

and, at interface:

$$-\gamma \rho \nabla \cdot \underline{\underline{\epsilon}} + \frac{1}{4\pi} \underline{\underline{B}}^i \cdot [\underline{\underline{d}} \underline{\underline{B}}^i + (\underline{\underline{\epsilon}} \cdot \underline{\underline{\nabla}}) \underline{\underline{B}}^i]$$

$$= \frac{1}{4\pi} \underline{\underline{B}}^e \cdot [\underline{\underline{d}} \underline{\underline{B}}^e + (\underline{\underline{\epsilon}} \cdot \underline{\underline{\nabla}}) \underline{\underline{B}}^e]$$

perturbed pressure balance

Now, for surface current pinch:

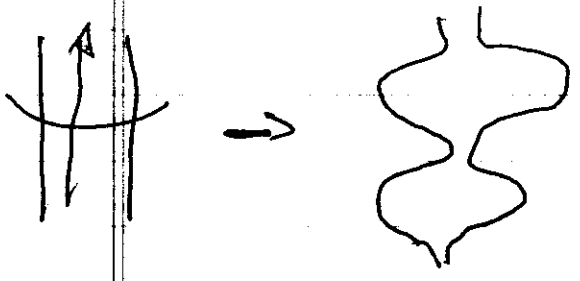
$$(\nabla \times \underline{\underline{B}})_{in} = 0 \quad (\nabla A)_{in} = 0 \Rightarrow \text{simplified}$$

no internal current, pressure gradient

so

$$\underline{\underline{F}} = \nabla [\gamma \rho \nabla \cdot \underline{\underline{\epsilon}}] + \frac{1}{4\pi} [\nabla \times \nabla \times (\underline{\underline{\epsilon}} \times \underline{\underline{B}})] \times \underline{\underline{B}}$$

Now, consider sausage (Link's):



$$\underline{\underline{\epsilon}} = (\epsilon_z, \epsilon_r, 0)$$

$\epsilon_\theta = 0$  (azimuthally symmetric perturbation)  
 ( $\epsilon_\theta \rightarrow$  torsion  $\rightarrow$  torsional Alfvén parity)

Thus - necking amplifies  $B_\theta$ ,  $\rho$ -inforcing necking.  $\Rightarrow$  destabilizing

out - necking compresses, bends  $B_z$   $\rightarrow$  }  
stabilizing.

$\Rightarrow$  suggests role of  $B_z$  in pinches/tokamak is for stability, and expect  $B_z$  min for stability.

$$\text{Now } \underline{\underline{\epsilon}} = \begin{Bmatrix} \epsilon_r(r) \\ \epsilon_z(r) \end{Bmatrix} e^{i(m\theta + kz)} e^{-i\omega t}$$

for surface current pinch  $\Rightarrow \underline{J}_{in} = 0, \underline{V}_{in} = 0$

$$-\omega^2 \rho \underline{\underline{\epsilon}} = \gamma \rho \underline{\nabla} (\underline{\nabla} \cdot \underline{\underline{\epsilon}}) + \frac{1}{4\pi} \left[ \underline{\nabla} \times \underline{\nabla} \times (\underline{\underline{\epsilon}} \times \underline{B}^i) \right] \times \underline{B}^i$$

taking  $m=0 \Rightarrow$  (azimuthally symmetric perturbation)

$$-\omega^2 \rho \epsilon_z = ck \gamma \rho \underline{\nabla} \cdot \underline{\underline{\epsilon}}; \quad \text{for } \epsilon_z$$

$$-\omega^2 \rho \epsilon_r = \gamma \rho \frac{d}{dr} (\underline{\nabla} \cdot \underline{\underline{\epsilon}}) + \frac{1}{4\pi} B_z \left\{ \frac{d}{dr} \frac{B_z}{r} \frac{d}{dr} (r \epsilon_r) - k^2 B_z \epsilon_r \right\}$$

then, can re-write:

for  $\epsilon_r$ .

$$\epsilon_z: \quad (C_s^2 k^2 - \omega^2) \epsilon_z = C_s^2 \frac{ik}{r} \frac{d}{dr} (r \epsilon_r)$$

and, for  $\epsilon_r$ ;

$$(V_A^2 k^2 - \omega^2) \epsilon_r = (c_s^2 + V_A^2) \frac{d}{dr} \left[ \frac{1}{r} \frac{d}{dr} (r \epsilon_r) \right] + ikc_s^2 \frac{d\epsilon_z}{dr}$$

Now, eliminating  $\epsilon_z$ :

$$\frac{d^2}{dr^2} \epsilon_r + \frac{1}{r} \frac{d\epsilon_r}{dr} - k^2 \epsilon_r = 0$$

$$k^2 = \frac{(c_s^2 k^2 - \omega^2)(V_A^2 k^2 - \omega^2)}{c_s^2 V_A^2 k^2 - (c_s^2 + V_A^2) \omega^2}$$

$$= k^2 \left[ \frac{1 + (\omega/k)^4}{c_s^2 V_A^2 - (c_s^2 + V_A^2) (\omega/k)^2} \right]$$

Now,  $\epsilon_r$  eqn.  $\Rightarrow$  Bessel eqn.

Clearly  $\epsilon_z|_{r=0} \rightarrow 0$  ; so

$$\epsilon_r = \sum_{l=0}^{\infty} I_0(kr)$$

Can use equations of motion to relate  $\epsilon_z$  to  $\epsilon_r$ !

$$\hat{\Sigma}_z = -\frac{c^2 k}{4} \left[ \frac{C_s^2 (V_A^2 k^2 - \omega^2) - V_A^2 \omega^2}{C_s^2 (V_A^2 k^2 - \omega^2)} \right] I_0'(kr)$$

Now, applying B.C.'s:

perturbed

- for pressure balance at boundary:

$$-\gamma \rho \nabla \cdot \hat{\underline{\underline{\epsilon}}} + \frac{1}{4\pi} \underline{\underline{B}}^i \cdot \underline{\underline{\nabla}} \underline{\underline{A}}^i = \frac{1}{4\pi} \underline{\underline{B}}^e \cdot \left[ \underline{\underline{\nabla}} \underline{\underline{B}}^e + \hat{\underline{\underline{\epsilon}}} \cdot \underline{\underline{\nabla}} \underline{\underline{B}}^e \right]$$

but  $\nabla \times \underline{\underline{\epsilon}} \times \underline{\underline{B}} = \underline{\underline{\nabla}} \underline{\underline{B}} \Rightarrow$

$$\underline{\underline{\nabla}} \underline{\underline{B}}_z = \frac{B_0}{r} \frac{d}{dr} (r \underline{\underline{\epsilon}}_r)$$

\*\*\*

$$\left\{ -\rho \nabla \cdot \hat{\underline{\underline{\epsilon}}} - \frac{B_0^2}{4\pi r} \frac{d}{dr} (r \hat{\underline{\underline{\epsilon}}}_r) = \frac{1}{4\pi} \underline{\underline{B}}_z^e \cdot \left( \nabla \times \underline{\underline{\nabla}} \underline{\underline{A}}^e + \hat{\underline{\underline{\epsilon}}}_r \frac{d}{dr} \underline{\underline{B}}^e \right) \right\}$$

- Now, for tangential e-field continuity  $\Rightarrow$

$$\hat{\underline{\underline{n}}} \times \underline{\underline{\nabla}} \underline{\underline{A}}^e = -(\underline{\underline{n}} \cdot \underline{\underline{\epsilon}}) \underline{\underline{B}}_\phi \quad (\text{forces term } \underline{\underline{\nabla}} \underline{\underline{A}}^e)$$

$$\Rightarrow \underline{\underline{\nabla}} \underline{\underline{A}}^e = \begin{pmatrix} \underline{\underline{\epsilon}}_r B_\phi \\ \underline{\underline{z}} \\ \delta A_r^e \\ \phi \end{pmatrix} \quad \begin{matrix} \downarrow \\ \text{form for} \\ \underline{\underline{A}}^e \end{matrix}$$

Now, in the gap:

slow  $\rightarrow$  neglect displ. current in MHO

$$\nabla \times \nabla \times \underline{A}^e - \frac{1}{c^2} \frac{\partial^2 \underline{A}^e}{\partial t^2} = 0$$

$$\Rightarrow \frac{1}{r} \frac{\partial}{\partial r} \left[ r \frac{\partial}{\partial z} \underline{A}_r^e - \frac{\partial}{\partial r} (B_\theta \epsilon_r) \right] = 0$$

$$- \frac{\partial}{\partial z} \left[ \frac{\partial}{\partial z} (\underline{A}_r^e) - \frac{\partial}{\partial r} (B_\theta \epsilon_r) \right] = 0$$

Now, as  $B_\theta = B_\theta(r)$  and  $\epsilon_r = \epsilon_r(r)$ , only:

$$\Rightarrow \frac{\partial}{\partial z} \underline{A}_r^e = f(r)$$

$$\frac{\partial}{\partial z} \underline{A}_r^e - \frac{\partial}{\partial r} (B_\theta \epsilon_r) = g(z)$$

now, as both  $\uparrow \uparrow$  are functions of  $r$ ,  
only,  $g(z)$  must be constant.

Take const  $\rightarrow 0$ , i.e. as only quantified level <sup>absolute</sup>  
external field.

Then, two eqns. above  $\Rightarrow$



$$dB^e \equiv \underline{\nabla} \times dA^e = 0 \quad \rightarrow \quad \text{i.e. no perturbation in external field}$$

Then, perturbed pressure <sup>balance</sup> boundary condition (\*\*)  
 $\Rightarrow$

$$\left\{ -\gamma P \underline{\nabla} \cdot \hat{\underline{E}} - \frac{B_0^2}{4\pi r} \frac{d}{dr} (r \hat{\underline{E}}_r) = -\frac{1}{4\pi r} B_0^2 \hat{\underline{E}}_r \right\} \text{ at } r=R$$

at  $r=R$ . Here used  $B_\phi = I/2\pi r$ .

Using  $\underline{\epsilon}_r, \underline{\epsilon}_z$  eqn. (eqn. motion)  $\Rightarrow$

$$\left[ \frac{\gamma^2 \omega^2 P}{i k C_s^2} - \frac{B_0^2 (C_s^2 k^2 - \omega^2)}{4\pi i k C_s^2} \right] \underline{\epsilon}_z = -\frac{B_0^2}{4\pi r} \underline{\epsilon}_r$$

at  $r=R$

Finally, plugging in for  $\underline{\epsilon}_z(R), \underline{\epsilon}_r(R)$  etc.

$$\Rightarrow \left\{ \omega^2 = k^2 V_A^2 - \frac{B_{ext}^2}{B_0^2} V_A^2 \frac{k}{R} \frac{I_0'(kR)}{I_0(kR)} \right\}$$

$$k^2 = k^2 \left[ 1 + \frac{(\omega/k)^4}{C_s^2 V_A^2 - (C_s^2 + V_A^2) (\omega/k)^2} \right] \quad (V_A \text{ with } B_0)$$

- nominally transcendental equation  $\Leftrightarrow$  numerics
- but  $\omega^2$  takes on only real values

(Hermiticity of  $\underline{F}$ )  $\Rightarrow$

$$B_0^2 \leq \frac{B_{\text{ext}}^2}{(kR)^2} \frac{(kR) I_0'(kR)}{I_0(kR)}$$

condition  
for  
instability

$\Rightarrow$  The message: Toroidal field necessary  
for pinch stability to  $m=0$  'sausage'  
mode.

$\Rightarrow$  physically, toroidal field introduces  
bending penalty: ( $\sim k^2 V_A^2$ ) for sausage  
perturbations.

$\Rightarrow$  N.B. Note similarity:

$$\omega^2 = k^2 c_s^2 - 4\pi G \rho_0 \quad \text{Jeans}$$

(negative compression)

$$\omega^2 = k^2 V_A^2 - \frac{B_{\text{ext}}^2}{B_0^2} \frac{V_A^2}{R} \frac{k}{\bar{I}_0} \frac{I_0'}{\bar{I}_0} \quad \text{Sausage}$$

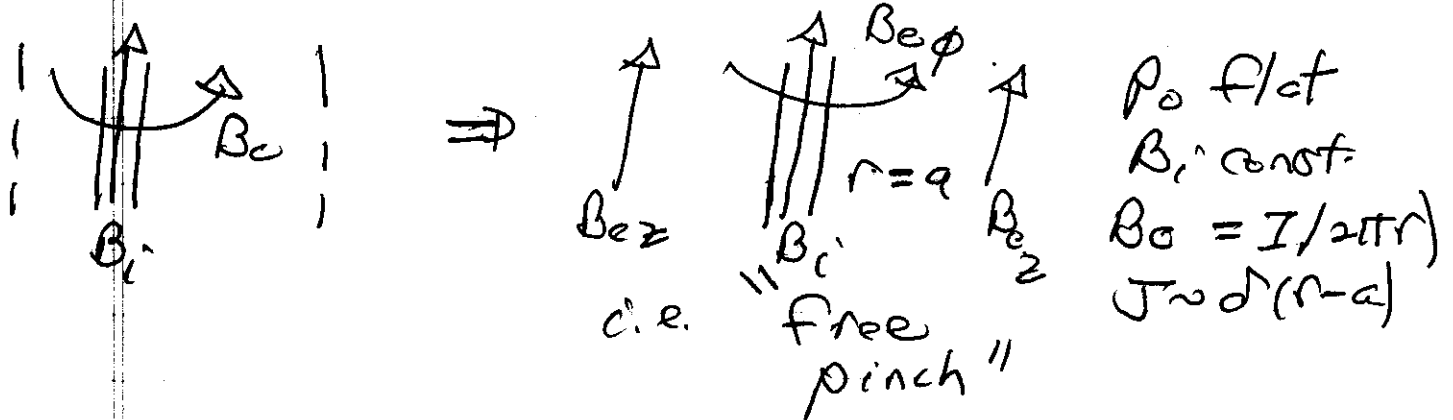
(negative tension)

# Links and kinks, cont'd.

170.

→ Surface Layer Pinch, Again  
(Simplified)

Now → remove outer wall to  $\infty$ .



- cylindrical symmetry

$$\underline{\hat{\underline{\underline{\epsilon}}}} = \underline{\underline{\underline{\epsilon}}}(r) e^{i(m\theta + kz)}$$

- for  $k \neq 0$ , can always arrange  $\underline{\underline{\underline{\epsilon}}}_z$   
so  $\underline{\underline{\underline{\nabla}}} \cdot \underline{\underline{\underline{\epsilon}}} = 0$

⇒ no loss of generality to take incompressible.

- for plasma dynamics:

$$\begin{aligned} \underline{\underline{\underline{\delta B}}} &= \underline{\underline{\underline{\nabla}}} \times (\underline{\underline{\underline{\epsilon}}} \times \underline{\underline{\underline{B}}}_0) \quad B_i \text{ const.} \\ &= \underline{\underline{\underline{B}}}_0 \cdot \underline{\underline{\underline{\nabla}}} \underline{\underline{\underline{\epsilon}}} - \underline{\underline{\underline{\epsilon}}} \cdot \underline{\underline{\underline{\nabla}}} \underline{\underline{\underline{B}}}_0 - \underline{\underline{\underline{B}}}_0 \nabla \cdot \underline{\underline{\underline{\epsilon}}} \end{aligned}$$

so

$$\rho_0 \frac{\partial \tilde{\rho}}{\partial t} = -\nabla \cdot \left( \tilde{\rho} + \frac{\underline{B}_0 \cdot \underline{\tilde{B}}}{4\pi} \right) + \frac{\underline{B}_0 \cdot \nabla \tilde{\rho}}{4\pi} + \frac{\underline{\tilde{B}} \cdot \nabla \underline{B}_0}{4\pi} \sim \underline{B}_0 \cdot \nabla \underline{\tilde{\epsilon}}$$

$$= -\nabla \cdot (\tilde{\rho}^*) + \frac{(\underline{B}_0 \cdot \nabla)(\underline{B}_0 \cdot \nabla) \underline{\tilde{\epsilon}}}{4\pi}$$

$\Rightarrow$

$$\left( -\omega^2 \rho_0 + k \frac{B_0}{4\pi} \right) \underline{\tilde{\epsilon}} = -\nabla \rho^*$$

now,  $\underline{\nabla} \cdot \underline{\tilde{\epsilon}} = 0 \Rightarrow \nabla^2 \tilde{\rho}^* = 0$

$\therefore \left( \nabla_r^2 - \frac{m^2}{r^2} - k^2 \right) \tilde{\rho}^* = 0$  |  $\rho$  well behaved as  $r \rightarrow 0$

$\Rightarrow \tilde{\rho}^*(r) = \tilde{\rho}^*(a) \frac{Im(kr)}{Im(ka)} \left[ \tilde{\rho}^* = \sum_m [a J_m(kr) + b Y_m(kr)] \right]$

(to match B.C.'s)

Thus, for displacement on boundary;

$$\left(-\omega^2 \rho_0 + \frac{k^2 B_i^2}{4\pi}\right) \tilde{E}_r(a) = -k \tilde{\rho}^*(a) \frac{I_m'(ka)}{I_m(ka)}$$

$$\Rightarrow \tilde{E}_r(a) = \frac{(4\pi k \tilde{\rho}^*(a))}{(4\pi \rho_0 \omega^2 - B_i^2 k^2)} \frac{I_m'(ka)}{I_m(ka)}$$

Exterior

$$\begin{aligned} \nabla \times \underline{\tilde{B}} &= 0 & \underline{B} &= \nabla \tilde{\psi} \\ \nabla \cdot \underline{\tilde{B}} &= 0 & \nabla^2 \tilde{\psi} &= 0 \end{aligned}$$

$$\Rightarrow \left(\nabla_r^2 - \frac{1}{r^2} \nabla^2 - k^2\right) \tilde{\psi} = 0$$

$$\Rightarrow \tilde{\psi} = C \frac{k_m(kr)}{k_m(ka)}, \quad \begin{cases} \int_{a0} k_m(kr) \\ \text{well behaved} \\ \text{at } r \rightarrow \infty. \end{cases}$$

→ B.C.'s

Observe have 3 constants  $\left\{ \begin{array}{l} C \\ \tilde{p}(a) \\ \tilde{\epsilon}_r(a) \end{array} \right\}$

Now, ① first: perturbed pressure balance (↑ as before)

$$\tilde{p} + \frac{\underline{B}_i \cdot \tilde{B}_i}{4\pi} \Big|_a = \frac{\underline{B}_e \cdot \tilde{B}_e}{4\pi} + \frac{\tilde{\epsilon}_r}{8\pi} \left( \frac{\partial B_{0e}^2}{\partial n} - \frac{\partial B_{0i}^2}{\partial n} \right)$$

$$= \frac{\underline{B}_e \cdot \nabla \tilde{\psi}}{4\pi} + \frac{\tilde{\epsilon}_r(a)}{8\pi} \left( \frac{\partial B_{0e}^2}{\partial r} \right. \\ \left. - \frac{\partial B_{0i}^2}{\partial r} \right) \Big|_a$$

$$B_{0e} = I/2\pi r \Rightarrow \frac{\partial}{\partial r} (B_{0e}^2) \Big|_a = -\frac{2B_{0e}^2}{a}$$

$$\Rightarrow \tilde{p}^*(a) = \frac{c}{4\pi} \left( \frac{m}{a} B_0 + k B_{ze} \right) C \\ - \frac{B_0^2}{4\pi a} \tilde{\epsilon}_r(a)$$

$$\textcircled{2} \quad \left( \underline{\hat{E}} + \frac{\underline{\hat{V}} \times \underline{B}_{00}}{c} \right) \Big|_a = 0$$

$$\nabla \times \Rightarrow \left( -\frac{1}{c} \frac{\partial \underline{\hat{B}}}{\partial t} + \nabla \times \frac{\underline{\hat{V}} \times \underline{B}_{00}}{c} \right) \Big|_a = 0$$

$$\Rightarrow \hat{n} \cdot \underline{\hat{B}}_0 \Big|_a = \hat{n} \cdot (\nabla \times \underline{\hat{E}} \times \underline{B}_{00}) \Big|_a$$

$$= \hat{n} \cdot (\underline{B}_{00} \cdot \nabla \underline{\hat{E}}) \Big|_a$$

n.b.  
 $(\hat{n} \cdot (\underline{\hat{E}} \cdot \nabla \underline{B})) = 0$

$$\hat{n} \cdot \nabla \psi \Big|_a = (\underline{B}_{00} \cdot \nabla) \underline{\hat{E}}_r \Big|_a$$

$$\Rightarrow \left( i (k B_{0z} + \frac{m\beta}{a}) \underline{\hat{E}}_r(a) = c k \frac{k'_m(k a)}{k_m(k a)} \right)$$

$$\textcircled{3} \quad \underline{\hat{E}}(a) = \left( \frac{4\pi k \rho^*(a)}{4\pi \rho_0 \omega^2 - B_i^2 k^2} \right) \frac{I_m'(ka)}{I_m(ka)}$$

Now ①, ②, ③  $\Rightarrow$  3 cases.

Solvability  $\Rightarrow$  dispersion relation:

$$4\pi\rho_0\omega^2 = \overset{\textcircled{1}}{B_{0z,c}^2} k^2 - \left( \kappa B_{0z0} + \frac{m\beta}{a} \right)^2 \frac{I_m'(ka) k_m(ka)}{I_m(ka) k_m'(ka)}$$

$$- \overset{\textcircled{3}}{B_0^2} k \frac{I_m'(ka)}{I_m(ka)}$$

①  $\rightarrow$  internal field line bending

②  $\rightarrow$  noting  $k_m'/k_m < 0$  , ②  $> 0$

$\Rightarrow$  external field line bending

$\Rightarrow$  ②  $\sim \left. (\underline{\kappa} \cdot \underline{B}_{e0})^2 \right|_{r=a}$

so will vanish when  $(\underline{\kappa} \cdot \underline{B}_e)^2 = 0$

i.e. perturbation aligns with field.



Now note: if  $\underline{k} \cdot \underline{B}_0 = 0$ ,  $\underline{k} \cdot \underline{B}_0 \neq 0$   
 usually  $\Rightarrow$  can't dodge  
 both bending terms, due  
 shear.

③  $\rightarrow$  negative definite!  $\Rightarrow$  origin of instability!

N.B. observe term ③ from  $\frac{\partial}{\partial r} (B_0^2)$ ,

and that  $\frac{\partial}{\partial r} B_0^2 < 0 \Rightarrow$

field decreases with increasing  
 distance from boundary!

Consider 2 cases

1.)  $B_{0z} = 0$

$$4\pi \rho \omega^2 = B_{0z}^2 k^2 + \left(\frac{m B_0}{a}\right)^2 \frac{I_m'}{I_m} \left| \frac{k_m}{k_m'} \right|$$

$$- B_0^2 \frac{k}{a} \frac{I_m'(ka)}{I_m(ka)}$$