

Physics 214 UCSD/225a UCSB

Lecture 9

- Outlook for remainder of quarter
- Halzen & Martin Chapter 3
- Start of Halzen & Martin Chapter 4

Outlook for remaining Quarter

- From now on I will follow H&M more closely.
- We'll basically cover chapters 3,4,5,6
 - A lot of this should be a review of things you have seen already either in advanced QM or intro QFT.
 - Accordingly, I'll be brief at times, and expect you to read up on it as needed !!!
- Then skip chapter 7.
- Then parts of 8,9,10, and 11, where I am not yet sure as to the order I'll do them in.
- Some of this we won't get to until next quarter.

(non-)relativistic Schroedinger Eq.

- Nonrelativistic

$$E = p^2/(2m)$$

- Relativistic

$$E^2 = p^2 + m^2$$

$$\left(i \frac{\partial}{\partial t} + \frac{\nabla^2}{2m} \right) \psi = 0$$

$$-\frac{\partial^2}{\partial t^2} \phi = \left(-\nabla^2 + m^2 \right) \phi$$

In both cases we replace:

$$E \rightarrow i \frac{\partial}{\partial t}$$

$$p \rightarrow -i \nabla$$

Covariant Notation

$$A^\mu = (A^0, \mathbf{A}) ; A_\mu = (A^0, -\mathbf{A})$$

$$A^\mu B_\mu = A^0 B^0 - \mathbf{A} \mathbf{B}$$

The derivative 4-vector is given by:

$$\partial^\mu = \left(\frac{\partial}{\partial t}, -\nabla \right)$$

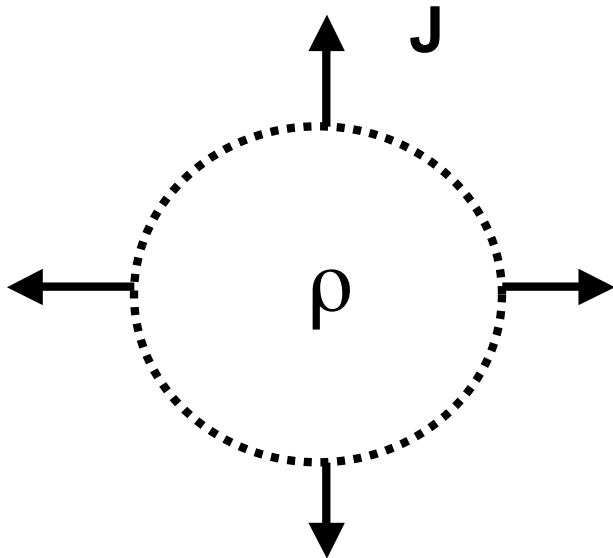
$$\partial_\mu = \left(\frac{\partial}{\partial t}, \nabla \right)$$

With: $\square^2 = \partial_\mu \partial^\mu$

We then get the Klein-Gordon Equation as: $(\square^2 + m^2) \phi = 0$

Continuity Equation

- For scattering, we need to understand the probability density flux \mathbf{J} , as well as the probability density ρ .
- Conservation of probability leads to:



$$-\frac{\partial}{\partial t} \int_V \rho dV = \int_S \vec{j} \cdot \hat{n} ds = \int_V \nabla J dV$$
$$\Rightarrow \nabla J + \frac{\partial \rho}{\partial t} = 0$$

(non-)relativistic Continuity Eq.

- Nonrelativistic

$$\rho = |\psi|^2$$

$$J = \frac{-i}{2m} (\psi^* \nabla \psi - \psi \nabla \psi^*)$$

- Relativistic

$$\rho = i \left(\phi^* \frac{\partial \phi}{\partial t} - \phi \frac{\partial \phi^*}{\partial t} \right)$$

$$J = -i (\phi^* \nabla \phi - \phi \nabla \phi^*)$$

In both cases we have plane wave solutions as:

$$\phi(t, \vec{x}) = N e^{-ip_\mu x^\mu}$$

Covariant Notation

$$j^\mu = (\rho, \vec{j}) \quad \text{Transforms like a 4-vector}$$

$$\partial^\mu j_\mu = 0 \quad \text{Covariant continuity equation}$$

For the plane wave solutions we find:

$$\left. \begin{aligned} \rho &= 2E |N|^2 \\ \vec{j} &= 2\vec{p} |N|^2 \end{aligned} \right\} J^\mu = 2p^\mu |N|^2$$

Why $\rho \propto E$?

$\rho d^3x = \text{constant}$ under lorentz transformations

However, d^3x gets lorentz contracted.

Therefore, ρ must transform time-like, i.e. dilate.

$$d^3x \rightarrow d^3x \cdot \sqrt{1-v^2}$$

$$\rho \rightarrow \rho / \sqrt{1-v^2}$$

Energy Eigenvalues of K.G. Eq.

$$(\square^2 + m^2) \phi = 0$$

Or

$$E^2 = p^2 + m^2$$

↓

$$E = \pm \sqrt{p^2 + m^2}$$

Positive and negative energy solutions !

Feynman-Stueckelberg Interpretation

- Positive energy particle moving forwards in time.
 - Negative energy antiparticles moving backwards in time.
- ⇒ Absorption of positron with $-E$ is the same as emission of electron with $+E$.
- ⇒ In both cases charge of system increases while energy decreases.

Encourage you to read up on this in chapters 3.4 & 3.5 of H&M.

- Will get back to discussing negative energy solutions after we understand scattering in a potential.
- Will use scattering in a potential to discuss perturbation theory.
 - Assume potential is finite in space.
 - Incoming and outgoing states are free-particle solutions “far enough” away from potential.
 - Assume V is a small perturbation throughout such that free particle, i.e. plane wave starting point is a meaningful approximation.

Nonrelativistic Perturbation Theory

- Assume we know the complete set of eigenstates of the free-particle Schroedinger Equation:

$$H\phi_n = E_n \phi_n$$

$$\int_{Vol} \phi_n^* \phi_m d^3x = \delta_{nm}$$

- Now solve Schroedinger Eq. in the presence of a small perturbation $V(\mathbf{x},t)$:

$$(H + V)\psi = i\frac{\partial\psi}{\partial t}$$

Any solution can be expressed as:

$$\psi = \sum_n a_n(t) \phi_n(x) e^{-iE_n t}$$

Plug this into Sch.Eq. and you get:

$$i \sum_n \frac{da_n(t)}{dt} \phi_n(x) e^{-iE_n t} = V(x, t) \sum_n a_n(t) \phi_n(x) e^{-iE_n t}$$

$$\frac{da_f(t)}{dt} = -i \sum_n a_n(t) e^{i(E_f - E_n)t} \int \phi_f^* V \phi_n d^3 x$$

Assume V is small and “seen” for only a finite amount of time.

- At times long ago, the system is in eigenstate i of the free hamiltonian because it's far away from V .
- At times far in the future, the system is in eigenstate f of the free hamiltonian because it's far away from V .
- After integration over time, we thus get:

$$\frac{da_f}{dt} = -ie^{-i(E_f - E_i)} \int d^3x \phi_f^* V \phi_i \quad \Leftarrow \text{starting point: } i \rightarrow f$$

$$V_{fi} \equiv \int d^3x \phi_f^* V(x) \phi_i \quad \Leftarrow \text{Assume } V \text{ is time independent}$$

$$T_{fi} \equiv a_f = -iV_{fi} \int dt e^{-i(E_f - E_i)} = \boxed{-2\pi i V_{fi} \delta(E_f - E_i)}$$

Result of time integration.

Meaning of $T_{fi} = -2\pi i V_{fi} \delta(E_f - E_i)$

- δ -function guarantees energy conservation.
- => Uncertainty principle guarantees that T_{fi} is meaningful only as $t \rightarrow$ infinity.
- We thus define a more meaningful quantity W , the **“transition amplitude per unit time”** by dividing with t , and then letting $t \rightarrow$ infinity.

$$W = \lim_{t \rightarrow \infty} \frac{|T_{fi}|^2}{t} = 2\pi |V_{fi}|^2 \delta(E_f - E_i)$$

Aside

$$W = \lim_{t \rightarrow \infty} \frac{|T_{fi}|^2}{t} = \lim_{t \rightarrow \infty} 2\pi \frac{|V_{fi}|^2}{t} \delta(E_f - E_i) \int dt e^{-i(E_f - E_i)t}$$

The second integral is basically t , and thus cancels with the $1/t$, making the limit building trivial.

$$\begin{aligned} T_{fi} T_{fi}^* &= (-iV_{fi} \int dt e^{-i(E_f - E_i)t}) (-iV_{fi} \int dt e^{-i(E_f - E_i)t})^* = \\ |T_{fi}|^2 &= 2\pi |V_{fi}|^2 \delta(E_f - E_i) \int dt e^{-i(E_f - E_i)t} \end{aligned}$$

Physically meaningful quantities

- The transition probability per unit time, W , becomes physically meaningful once you integrate over a set of initial and final states.
- Though typically, we start with a specific initial and a set of final states:

$$W_{fi} = 2\pi \int dE_f \rho(E_f) |V_{fi}|^2 \delta(E_f - E_i)$$

$$W_{fi} = 2\pi |V_{fi}|^2 \rho(E_i) \quad \leftarrow \text{Fermi's Golden Rule}$$

Fermi's Golden Rule

- We find Fermi's Golden Rule as the leading order in perturbation theory.
- This begs the question, what's the next order, and how do we get it?
- In our lowest order approximation, we scattered from an initial state i to a final state f .
- The obvious improvement is to allow for double scattering from i to any n to f , and sum over all n .

Second Order

$$a_n = -i \int dt e^{i(E_n - E_i)t} V_{ni}$$

$$\frac{da_f(t)}{dt} = LO - i \sum_{n \neq i} a_n(t) e^{i(E_f - E_n)t} V_{fn}$$

$$W_{fi} = 2\pi \left| V_{fi} + \sum_{n \neq i} V_{fn} \frac{1}{E_i - E_n + i\varepsilon} V_{ni} \right|^2 \rho(E_i)$$

What have we learned?

- For each interaction vertex we get a vertex factor V_{fi} .
- For the propagation via an intermediate state we gain a “propagator” factor $1/(E_i - E_n)$.
- The intermediate state is virtual, and thus does not require energy conservation.
- However, energy is conserved between initial and final state.

$$W_{fi} = 2\pi \left| V_{fi} + \sum_{n \neq i} V_{fn} \frac{1}{E_i - E_n + i\epsilon} V_{ni} \right|^2 \rho(E_i)$$

Photon absorption by Particle vs Antiparticle

- Particle scatter in field

$$p_i = (E_1, \vec{p}_1)$$

$$p_f = (E_2, \vec{p}_2)$$

$$T_{fi} \approx \int dt \phi_f^* V \phi_i \approx \int dt e^{iE_2 t} e^{-i\omega t} e^{-iE_1 t}$$
$$\approx \delta(E_2 - (\omega + E_1))$$

- Antipart. scatter in field

$$p_i = (-E_2, -\vec{p}_2)$$

$$p_f = (-E_1, -\vec{p}_1)$$

$$T_{fi} \approx \int dt \phi_f^* V \phi_i \approx \int dt e^{i(-E_1)t} e^{-i\omega t} e^{-i(-E_2)t}$$
$$\approx \delta(E_2 - (\omega + E_1))$$

Particle and antiparticle have the same interaction with EM field.

Pair Creation from this potential

$$p_i = (-E_1, -\vec{p}_1)$$

$$p_f = (E_2, \vec{p}_2)$$

$$T_{fi} \approx \int dt \phi_f^* V \phi_i \approx \int dt e^{iE_2 t} e^{-i\omega t} e^{-i(-E_1)t}$$

$$\approx \delta(E_2 + E_1 - \omega)$$

Energy is conserved as it should be.

This wave function formalism is thus capable of describing particles, antiparticles, and pair production.

“Rules”

- Antiparticles get arrow that is backwards in time.
- “Incoming” and “outgoing” is defined by how the arrows point to the vertex.
- Antiparticles get negative energy assigned.

H&M Chapter 4

- Electrodynamics of Spinless particles
 - We replace p^μ with $p^\mu + eA^\mu$ in classical EM for a particle of charge $-e$ moving in an EM potential A^μ
 - In QM, this translates into:

$$i\partial^\mu \rightarrow i\partial^\mu + eA^\mu$$

- And thus to the modified Klein Gordon Equation:

$$\left(\partial^\mu \partial_\mu + m^2\right)\phi = -V\phi$$

$$V = -ie(\partial^\mu A_\mu + A^\mu \partial_\mu) - e^2 A^2$$

V here is the potential energy of the perturbation.

Take results form Perturbation

$$V = -ie(\partial^\mu A_\mu + A^\mu \partial_\mu) - e^2 A^2$$

$$T_{fi} = -i \int \phi_f^* V(x) \phi_i d^4 x \quad \leftarrow \text{covariant form} \quad \text{Integrate by parts}$$

$$T_{fi} = i \int \phi_f^* ie(A^\mu \partial_\mu + \partial^\mu A_\mu) \phi_i d^4 x + O(e^2)$$

$$T_{fi} = i \int ie(\phi_f^* (\partial_\mu \phi_i) - (\partial_\mu \phi_f^*) \phi_i) A^\mu d^4 x + O(e^2)$$

$$T_{fi} = -i \int J_\mu A^\mu d^4 x + O(e^2)$$

EM current for $i \rightarrow f$ transition.

$$J_\mu = -ie(\phi_f^* (\partial_\mu \phi_i) - (\partial_\mu \phi_f^*) \phi_i)$$

$$J_\mu = -eN_i N_f (p_i + p_f) e^{i(p_f - p_i)x}$$

Using plane wave solutions

Aside on current

Regular current we talked about in the beginning today:

$$\left. \begin{aligned} \rho &= 2E |N|^2 \\ \vec{j} &= 2\vec{p} |N|^2 \end{aligned} \right\} J^\mu = 2p^\mu |N|^2$$

Transition current from i to f:

$$J_\mu = -ie \left(\phi_f^* (\partial_\mu \phi_i) - (\partial_\mu \phi_f^*) \phi_i \right)$$

$$J_\mu = -e N_i N_f (p_i + p_f) e^{i(p_f - p_i)x}$$

The difference is that for regular current $i=f$, and the wave function piece cancels as a result.

Electron Muon Scattering

Overview

- Use what we just did
 - Electron scattering in EM field
- With the field being the one generated by the muon as source.
 - Use covariant form of maxwell's equation in Lorentz Gauge to get V , the perturbation potential.
- Plug it into T_{fi}
- Then head into more general discussion of how to express cross section in terms of invariant amplitude (or “Matrix Element”).

Electron Muon scattering

$$\square^2 A^\mu = J_{(2)}^\mu \quad \text{Maxwell Equation}$$

Note: $\square^2 e^{iqx} = -q^2 e^{iqx}$

$$J_{(2)}^\mu = -eN_B N_D (p_D + p_B)^\mu e^{i(p_D - p_B)x}$$

$$A^\mu = -\frac{1}{q^2} J_{(2)}^\mu \quad \rightarrow = q$$

$$T_{fi} = -i \int J_\mu^{(1)} \frac{-1}{q^2} J_{(2)}^\mu d^4 x \quad \leftarrow \text{Note the symmetry: (1) } \leftrightarrow \text{(2)}$$

$$T_{fi} = -iN_A N_B N_C N_D (2\pi)^4 \delta^{(4)}(p_D + p_C - p_A - p_B) M$$

$$-iM = (ie(p_A + p_C)^\mu) \frac{-ig_{\mu\nu}}{q^2} (ie(p_D + p_B)^\nu)$$

Note the structure: Vertex x propagator x Vertex