

PHYSICS 140A : STATISTICAL PHYSICS
HW ASSIGNMENT #3 SOLUTIONS

(1) Consider a generalization of the situation in §4.4 of the notes where now three reservoirs are in thermal contact, with any pair of systems able to exchange energy.

- (a) Assuming interface energies are negligible, what is the total density of states $D(E)$? Your answer should be expressed in terms of the densities of states functions $D_{1,2,3}$ for the three individual systems.
- (b) Find an expression for $P(E_1, E_2)$, which is the joint probability distribution for system 1 to have energy E_1 while system 2 has energy E_2 and the total energy of all three systems is $E_1 + E_2 + E_3 = E$.
- (c) Extremize $P(E_1, E_2)$ with respect to $E_{1,2}$. Show that this requires the temperatures for all three systems must be equal: $T_1 = T_2 = T_3$. Writing $E_j = E_j^* + \delta E_j$, where E_j^* is the extremal solution ($j = 1, 2$), expand $\ln P(E_1^* + \delta E_1, E_2^* + \delta E_2)$ to second order in the variations δE_j . Remember that

$$S = k_B \ln D \quad , \quad \left(\frac{\partial S}{\partial E} \right)_{V,N} = \frac{1}{T} \quad , \quad \left(\frac{\partial^2 S}{\partial E^2} \right)_{V,N} = -\frac{1}{T^2 C_V} .$$

- (d) Assuming a Gaussian form for $P(E_1, E_2)$ as derived in part (c), find the variance of the energy of system 1,

$$\text{Var}(E_1) = \langle (E_1 - E_1^*)^2 \rangle .$$

Solution :

- (a) The total density of states is a convolution:

$$D(E) = \int_{-\infty}^{\infty} dE_1 \int_{-\infty}^{\infty} dE_2 \int_{-\infty}^{\infty} dE_3 D_1(E_1) D_2(E_2) D_3(E_3) \delta(E - E_1 - E_2 - E_3) .$$

- (b) The joint probability density $P(E_1, E_2)$ is given by

$$P(E_1, E_2) = \frac{D_1(E_1) D_2(E_2) D_3(E - E_1 - E_2)}{D(E)} .$$

- (c) We set the derivatives $\partial \ln P / \partial E_{1,2} = 0$, which gives

$$\frac{\partial \ln P}{\partial E_1} = \frac{\partial \ln D_1}{\partial E_1} - \frac{\partial D_3}{\partial E_3} = 0 \quad , \quad \frac{\partial \ln P}{\partial E_2} = \frac{\partial \ln D_2}{\partial E_2} - \frac{\partial D_3}{\partial E_3} = 0 ,$$

where $E_3 = E - E_1 - E_2$ in the argument of $D_3(E_3)$. Thus, we have

$$\frac{\partial \ln D_1}{\partial E_1} = \frac{\partial \ln D_2}{\partial E_2} = \frac{\partial \ln D_3}{\partial E_3} \equiv \frac{1}{T}.$$

Expanding $\ln P(E_1^* + \delta E_1, E_2^* + \delta E_2)$ to second order in the variations δE_j , we find the first order terms cancel, leaving

$$\ln P(E_1^* + \delta E_1, E_2^* + \delta E_2) = \ln P(E_1^*, E_2^*) - \frac{(\delta E_1)^2}{2k_B T^2 C_1} - \frac{(\delta E_2)^2}{2k_B T^2 C_2} - \frac{(\delta E_1 + \delta E_2)^2}{2k_B T^2 C_3} + \dots,$$

where $\partial^2 \ln D_j / \partial E^2 = -1/2k_B T^2 C_j$, with C_j the heat capacity at constant volume and particle number. Thus,

$$P(E_1, E_2) = \frac{\sqrt{\det(\mathcal{C}^{-1})}}{2\pi k_B T^2} \exp\left(-\frac{1}{2k_B T^2} C_{ij}^{-1} \delta E_i \delta E_j\right),$$

where the matrix \mathcal{C}^{-1} is defined as

$$\mathcal{C}^{-1} = \begin{pmatrix} C_1^{-1} + C_3^{-1} & C_3^{-1} \\ C_3^{-1} & C_2^{-1} + C_3^{-1} \end{pmatrix}.$$

One finds

$$\det(\mathcal{C}^{-1}) = C_1^{-1} C_2^{-1} + C_1^{-1} C_3^{-1} + C_2^{-1} C_3^{-1}.$$

The prefactor in the above expression for $P(E_1, E_2)$ has been fixed by the normalization condition $\int dE_1 \int dE_2 P(E_1, E_2) = 1$.

(d) Integrating over E_2 , we obtain $P(E_1)$:

$$P(E_1) = \int_{-\infty}^{\infty} dE_2 P(E_1, E_2) = \frac{1}{\sqrt{2\pi k_B \tilde{C}_1 T^2}} e^{-(\delta E_1)^2 / 2k_B \tilde{C}_1 T^2},$$

where

$$\tilde{C}_1 = \frac{C_2^{-1} + C_3^{-1}}{C_1^{-1} C_2^{-1} + C_1^{-1} C_3^{-1} + C_2^{-1} C_3^{-1}}.$$

Thus,

$$\langle (\delta E_1)^2 \rangle = \int_{-\infty}^{\infty} dE_1 (\delta E_1)^2 = k_B \tilde{C}_1 T^2.$$

(2) Consider a two-dimensional gas of identical classical, noninteracting, massive relativistic particles with dispersion $\varepsilon(\mathbf{p}) = \sqrt{\mathbf{p}^2 c^2 + m^2 c^4}$.

(a) Compute the free energy $F(T, V, N)$.

(b) Find the entropy $S(T, V, N)$.

(c) Find an equation of state relating the fugacity $z = e^{\mu/k_B T}$ to the temperature T and the pressure p .

Solution :

(a) We have $Z = (\zeta A)^N / N!$ where A is the area and

$$\zeta(T) = \int \frac{d^2p}{h^2} e^{-\beta\sqrt{p^2c^2+m^2c^4}} = \frac{2\pi}{(\beta\hbar c)^2} (1 + \beta mc^2) e^{-\beta mc^2} .$$

To obtain this result it is convenient to change variables to $u = \beta\sqrt{p^2c^2+m^2c^4}$, in which case $p dp = u du / \beta^2 c^2$, and the lower limit on u is mc^2 . The free energy is then

$$F = -k_B T \ln Z = N k_B T \ln \left(\frac{2\pi\hbar^2 c^2 N}{(k_B T)^2 A} \right) - N k_B T \ln \left(1 + \frac{mc^2}{k_B T} \right) + N mc^2 .$$

where we are taking the thermodynamic limit with $N \rightarrow \infty$.

(b) We have

$$S = -\frac{\partial F}{\partial T} = -N k_B \ln \left(\frac{2\pi\hbar^2 c^2 N}{(k_B T)^2 A} \right) + N k_B \ln \left(1 + \frac{mc^2}{k_B T} \right) + N k_B \left(\frac{mc^2 + 2k_B T}{mc^2 + k_B T} \right) .$$

(c) The grand partition function is

$$\Xi(T, V, \mu) = e^{-\beta\Omega} = e^{\beta p V} = \sum_{N=0}^{\infty} Z_N(T, V, N) e^{\beta\mu N} .$$

We then find $\Xi = \exp(\zeta A e^{\beta\mu})$, and

$$p = \frac{(k_B T)^3}{2\pi(\hbar c)^2} \left(1 + \frac{mc^2}{k_B T} \right) e^{(\mu - mc^2)/k_B T} .$$

Note that

$$n = \frac{\partial(\beta p)}{\partial\mu} = \frac{p}{k_B T} \implies p = n k_B T .$$

(3) A three-level system has energy levels $\varepsilon_0 = 0$, $\varepsilon_1 = \Delta$, and $\varepsilon_2 = 4\Delta$. Find the free energy $F(T)$, the entropy $S(T)$ and the heat capacity $C(T)$.

Solution :

We have

$$Z = \text{Tr} e^{-\beta H} = 1 + e^{-\beta\Delta} + e^{-4\beta\Delta} .$$

The free energy is

$$F = -k_B T \ln Z = -k_B T \ln(1 + e^{-\Delta/k_B T} + e^{-4\Delta/k_B T}) .$$

To find the entropy S , we differentiate with respect to temperature:

$$S = -\left. \frac{\partial F}{\partial T} \right|_{V,N} = k_B \ln(1 + e^{-\Delta/k_B T} + e^{-4\Delta/k_B T}) + \frac{\Delta}{T} \cdot \frac{e^{-\Delta/k_B T} + 4e^{-4\Delta/k_B T}}{1 + e^{-\Delta/k_B T} + e^{-4\Delta/k_B T}} .$$

Now differentiate with respect to T one last time to find

$$C_{V,N} = k_B \left(\frac{\Delta}{k_B T} \right)^2 \cdot \frac{e^{-\Delta/k_B T} + 16e^{-4\Delta/k_B T} + 9e^{-5\Delta/k_B T}}{(1 + e^{-\Delta/k_B T} + e^{-4\Delta/k_B T})^2} .$$

(4) Consider a many-body system with Hamiltonian $\hat{H} = \frac{1}{2}\hat{N}(\hat{N} - 1)U$, where \hat{N} is the particle number and $U > 0$ is an interaction energy. Assume the particles are identical and can be described using Maxwell-Boltzmann statistics, as we have discussed. Assuming $\mu = 0$, plot the entropy S and the average particle number N as functions of the scaled temperature $k_B T/U$. (You will need to think about how to impose a numerical cutoff in your calculations.)

Solution :

The grand partition function is

$$\Xi(T, \mu) = e^{-\beta\Omega} = e^{\beta pV} = \sum_{N=0}^{\infty} e^{-N(N-1)\beta U/2} ,$$

where we have taken $\mu = 0$ and we have assumed that each state of definite particle number $|N\rangle$, is nondegenerate. We then have the grand potential

$$\Omega(T, \mu) = -k_B T \ln \Xi = -k_B T \ln \left(\sum_{N=0}^{\infty} e^{-N(N-1)U/2k_B T} \right)$$

The entropy is

$$S = -\frac{\partial \Omega}{\partial T} = k_B \ln \left(\sum_{N=0}^{\infty} e^{-N(N-1)U/2k_B T} \right) + \frac{U}{2T} \cdot \frac{\sum_{N=0}^{\infty} N(N-1) e^{-N(N-1)U/2k_B T}}{\sum_{N=0}^{\infty} e^{-N(N-1)U/2k_B T}} .$$

This must be evaluated numerically. The results are shown in Fig. 1. Note that $\lim_{T \rightarrow 0} S(T) = k_B \ln 2$, which indicates a doubly degenerate ground state. This is because both $|N=0\rangle$ and $|N=1\rangle$ have energy $E_0 = E_1 = 0$.

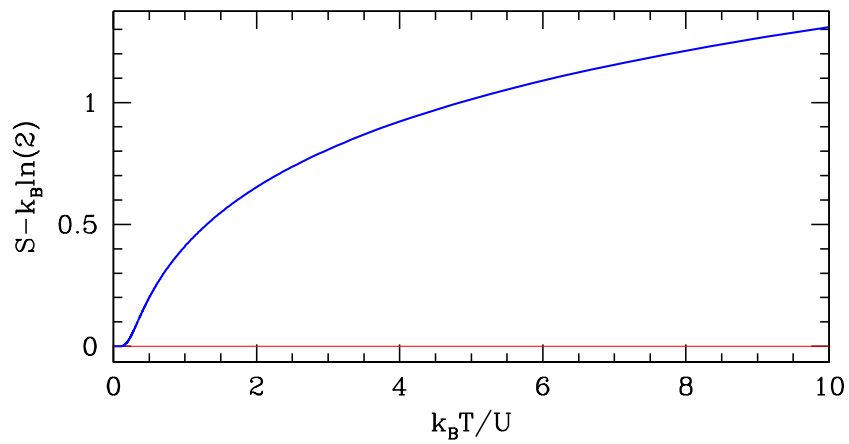


Figure 1: Entropy as a function of dimensionless temperature for problem #4. Note that $S(T = 0) = \ln 2$ because the states $|N = 0\rangle$ and $|N = 1\rangle$ are degenerate.