

**PHYSICS 140A : STATISTICAL PHYSICS
HW ASSIGNMENT #4 SOLUTIONS**

(1) Consider a noninteracting classical gas with Hamiltonian

$$\mathcal{H} = \sum_{i=1}^N \varepsilon(\mathbf{p}_i),$$

where $\varepsilon(\mathbf{p})$ is the dispersion relation. Define

$$\xi(T) = h^{-d} \int d^d p e^{-\varepsilon(\mathbf{p})/k_B T}.$$

- (a) Find $F(T, V, N)$.
- (b) Find $G(T, p, N)$.
- (c) Find $\Omega(T, V, \mu)$.
- (d) Show that

$$\beta p \int_0^\infty dV e^{-\beta p V} Z(T, V, N) = e^{-G(T, p, N)/k_B T}.$$

Solution :

(a) We have $Z(T, V, N) = (V\xi)^N/N!$, so

$$F(T, V, N) = -k_B T \ln Z(T, V, N) = -Nk_B T \ln\left(\frac{V}{N}\right) - Nk_B T \ln \xi(T) - Nk_B T.$$

(b) G is obtained from F by Legendre transform: $G = F + pV$, i.e.

$$G(T, p, N) = -Nk_B T \ln\left(\frac{k_B T}{p}\right) - Nk_B T \ln \xi(T).$$

Note that we have used the ideal gas law $pV = Nk_B T$ here.

(c) Ω is obtained from F by Legendre transform: $\Omega = F - \mu N$. Another way to obtain Ω is to use the grand potential $\Xi = \exp(V\xi(T) e^{\mu/k_B T})$, whence

$$\Omega(T, V, \mu) = -Vk_B T \xi(T) e^{\mu/k_B T}.$$

(d) We have

$$Y(T, p, N) = \beta p \int_0^\infty dV e^{-\beta p V} Z(T, V, N) = \frac{\xi^N(T)}{N!} \beta p \int_0^\infty dV V^N e^{-\beta p V} = \left(\frac{k_B T \xi(T)}{p}\right)^N$$

Thus, $G(T, p, N) = -Nk_B T \ln(k_B T \xi/p)$. Note that if we normalize the volume integral differently and define

$$Y(T, p, N) = \int_0^\infty \frac{dV}{V_0} e^{-\beta p V} Z(T, V, N) = \left(\frac{k_B T}{p V_0} \right) \cdot \left(\frac{k_B T \xi(T)}{p} \right)^N,$$

we obtain $G(T, p, N) = -Nk_B T \ln(k_B T \xi/p) - k_B T \ln(k_B T/pV_0)$, which differs from the previous result only by an $\mathcal{O}(N^0)$ term, which is subextensive and hence negligible in the thermodynamic limit.

(2) A three-dimensional gas of magnetic particles in an external magnetic field H is described by the Hamiltonian

$$\mathcal{H} = \sum_i \left[\frac{\mathbf{p}_i^2}{2m} - \mu_0 H \sigma_i \right],$$

where $\sigma_i = \pm 1$ is the spin polarization of particle i and μ_0 is the magnetic moment per particle.

- Working in the ordinary canonical ensemble, derive an expression for the magnetization of system.
- Repeat the calculation for the grand canonical ensemble. Also, find an expression for the Landau free energy $\Omega(T, V, \mu)$.
- Calculate how much heat will be given off by the system when the magnetic field is reduced from H to zero at constant volume, constant temperature, and particle number.

Solution :

(a) The partition function trace is now an integral over all coordinates and momenta with measure $d\mu$ as before, plus a sum over all individual spin polarizations. Thus,

$$\begin{aligned} Z &= \text{Tr} e^{-\mathcal{H}/k_B T} = \frac{1}{N!} \prod_{i=1}^N \sum_{\sigma_i} \int \frac{d^3 x_i d^3 p_i}{h^3} e^{-\mathbf{p}_i^2/2mk_B T} e^{\mu_0 H \sigma_i/k_B T} \\ &= \frac{1}{N!} V^N \lambda_T^{-3N} \left[2 \cosh(\mu_0 H/k_B T) \right]^N, \end{aligned}$$

where $\lambda_T = (2\pi\hbar^2/mk_B T)^{1/2}$ is the thermal wavelength. The Helmholtz free energy is

$$\begin{aligned} F(T, V, H, N) &= -k_B T \ln Z(T, V, H, N) \\ &= -Nk_B T \ln \left(\frac{V}{N \lambda_T^3} \right) - Nk_B T \ln \cosh(\mu_0 H/k_B T) - Nk_B T (1 + \ln 2). \end{aligned}$$

The magnetization is then

$$M(T, V, H, N) = -\frac{\partial F}{\partial H} = N\mu_0 \tanh(\mu_0 H/k_B T) .$$

(b) The grand partition function is

$$\Xi(T, V, H, \mu) = \sum_{N=0}^{\infty} e^{\mu N/k_B T} Z(T, V, N) = \exp \left(V\lambda_T^{-3} \cdot 2 \cosh(\mu_0 H/k_B T) \cdot e^{\mu/k_B T} \right) .$$

Thus,

$$\Omega(T, V, H, \mu) = -k_B T \ln \Xi(T, V, \mu) = -V k_B T \lambda_T^{-3} \cdot 2 \cosh(\mu_0 H/k_B T) \cdot e^{\mu/k_B T} .$$

Then

$$M(T, V, H, \mu) = -\frac{\partial \Omega}{\partial H} = 2\mu_0 \cdot V\lambda_T^{-3} \cdot \sinh(\mu_0 H/k_B T) \cdot e^{\mu/k_B T} .$$

Note that

$$N(T, V, H, \mu) = -\frac{\partial \Omega}{\partial \mu} = V\lambda_T^{-3} \cdot \cosh(\mu_0 H/k_B T) \cdot e^{\mu/k_B T} ,$$

so $M = N\mu_0 \tanh(\mu_0 H/k_B T)$, which agrees with the result from part (a).

(c) Starting with our expression for $F(T, V, N)$ in part (a), we differentiate to find the entropy:

$$S(T, V, H, N) = -\frac{\partial F}{\partial T} = Nk_B \ln \cosh(\mu_0 H/k_B T) - \frac{N\mu_0 H}{T} \tanh(\mu_0 H/k_B T) + S(T, V, 0, N) ,$$

where $S(T, V, 0, N)$ is the entropy at $H = 0$, which we don't need to compute for this problem. The heat absorbed *by* the system is

$$\begin{aligned} Q &= \int dQ = TS(0) - TS(H) = Nk_B T \ln \cosh(\mu_0 H/k_B T) + N\mu_0 H \tanh(\mu_0 H/k_B T) \\ &= Nk_B T \left(x \tanh x - \ln \cosh x \right) , \end{aligned}$$

where $x = \mu_0 H/k_B T$. Defining $f(x) = x \tanh x - \ln \cosh x$, one has $f'(x) = x \operatorname{sech}^2 x$ which is positive for all $x > 0$. Since $f(x)$ is an even function with $f(0) = 0$, we conclude $f(x) > 0$ for $x \neq 0$. Thus, $Q > 0$, which means that the system absorbs heat under this process. *I.e.* the heat released by the system is $(-Q)$.

(3) A classical three-dimensional gas of noninteracting particles has the Hamiltonian

$$\mathcal{H} = \sum_{i=1}^N \left[A |\mathbf{p}_i|^s + B |\mathbf{q}_i|^t \right] ,$$

where s and t are nonnegative real numbers.

- (a) Find the free energy $F(T, V, N)$.
- (b) Find the average energy $E(T, V, N)$.
- (c) Find the grand potential $\Omega(T, V, \mu)$.

Remember the definition of the Gamma function, $\Gamma(z) = \int_{-\infty}^{\infty} du u^{z-1} e^{-u}$.

Solution :

(a) Working in the OCE, the partition function is $Z = \xi_p^N(T) \xi_q^N(T)/N!$, where

$$\xi_p(T) = \frac{1}{h^3} \int d^3p \exp(-A p^s/k_B T)$$

$$\xi_q(T) = \int d^3q \exp(-B q^t/k_B T) .$$

We focus first on the momentum integral, changing variables to $u = A p^s/k_B T$. Then

$$u = \frac{A p^s}{k_B T} \Rightarrow p = \left(\frac{k_B T u}{A}\right)^{1/s}, \quad p^2 dp = \left(\frac{k_B T}{A}\right)^{3/s} \cdot s^{-1} u^{(3/s)-1} du ,$$

and

$$\begin{aligned} \xi_p(T) &= \frac{1}{h^3} \int d^3p \exp(-A p^s/k_B T) = \frac{4\pi}{h^3} \left(\frac{k_B T}{A}\right)^{3/s} \cdot \frac{1}{s} \int_{-\infty}^{\infty} du u^{(3/s)-1} e^{-u} \\ &= \frac{4\pi}{s h^3} \Gamma(3/s) \left(\frac{k_B T}{A}\right)^{3/s}, \end{aligned}$$

where we have used $z \Gamma(z) = \Gamma(z + 1)$. *Mutatis mutandis*,

$$\xi_q(T) = \int d^3q \exp(-B q^t/k_B T) = \frac{4\pi}{t} \Gamma(3/t) \left(\frac{k_B T}{B}\right)^{3/t} .$$

Thus, the free energy is

$$F(T, V, N) = -k_B T \ln Z = -N k_B T \ln \left(\frac{\xi_p(T) \xi_q(T)}{N} \right) - N k_B T .$$

(b) The average energy is

$$E = \frac{\partial}{\partial \beta} (\beta F) = \left(\frac{3}{s} + \frac{3}{t} \right) N k_B T .$$

(c) The grand potential is $\Omega = -k_B T \ln \Xi$, and $\Xi = \exp\left(\xi_p(T) \xi_q(T) e^{\mu/k_B T}\right)$. Thus,

$$\Omega(T, V, N) = -k_B T \xi_p(T) \xi_q(T) e^{\mu/k_B T} .$$

Note that F and Ω are both independent of V , which means that the pressure p vanishes!

(4) A gas of nonrelativistic particles of mass m is held in a container at constant pressure p and temperature T . It is free to exchange energy with the outside world, but the particle number N remains fixed. Compute the variance in the system volume, $\text{Var}(V)$, and the ratio $(\Delta V)_{\text{rms}}/\langle V \rangle$. Use the Gibbs ensemble.

Solution : The Gibbs free energy is

$$G(T, p, N) = -N k_B T \ln \left(\frac{k_B T}{p \lambda_T^3} \right) ,$$

where $\lambda_T = (2\pi\hbar^2/mk_B T)^{1/2}$ is the thermal wavelength. Thus, with

$$Y = e^{-G/k_B T} = \int \frac{dV}{V_0} e^{-\beta p V} Z(T, V, N) ,$$

we have

$$\begin{aligned} \langle V \rangle &= -\frac{1}{\beta} \frac{1}{Y} \frac{\partial Y}{\partial p} = \frac{\partial G}{\partial p} = \frac{N k_B T}{p} \\ \text{Var}(V) &= \langle V^2 \rangle - \langle V \rangle^2 = \frac{1}{\beta^2} \left\{ \frac{1}{Y} \frac{\partial^2 Y}{\partial p^2} - \left(\frac{1}{Y} \frac{\partial Y}{\partial p} \right)^2 \right\} = -k_B T \frac{\partial^2 G}{\partial p^2} = N \left(\frac{k_B T}{p} \right)^2 . \end{aligned}$$

Thus, $(\Delta V)_{\text{RMS}} = \sqrt{\text{Var}(V)}/\langle V \rangle = N^{-1/2}$.