

Nonlinear Oscillators

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I.) Nonlinear Oscillations - An Introduction

a) Overview / Roadmap

Can formulate sequence of problems:

- 1.) Simple harmonic oscillator - linear
- external force $\left\{ \begin{array}{l} \text{deterministic} \\ \text{random} \end{array} \right.$

$$\ddot{x} + \alpha \dot{x} + \omega_0^2 x = f_{\text{ext}}$$

\uparrow damping (friction const.) $\rightarrow k/m$

if $f_{\text{ext}} = f e^{i\delta t} \Rightarrow \begin{cases} x = A e^{i\delta t} \\ A = F/m(\omega_0^2 - \gamma^2 + 2i\alpha\delta) \end{cases}$

- 2.) Variation: SHO (Linear), with Parametric Variation
i.e. $\omega_0^2 = \omega_0^2(t)$

$\omega_0^2(t)$ $\left\{ \begin{array}{l} \text{slow variation } \frac{1}{\omega_0^2} \frac{d}{dt} \omega_0^2 \ll \omega_0 \Rightarrow \text{adiabatic theory} \\ \text{fast variation } \frac{1}{\omega_0^2} \frac{d}{dt} \omega_0^2 > \omega_0 \Rightarrow \text{parametric instability} \\ \quad \quad \quad \sim 2\omega_0 \end{array} \right.$

For parametric instability, prototype system is:

$$\begin{cases} \ddot{x} + \omega^2(t) x = 0 \\ \omega^2(t) = \omega_0^2 (1 + h \cos \gamma(t)) \\ \gamma \sim 2\omega_0 + \epsilon, \quad \epsilon \ll \omega_0 \end{cases} \begin{cases} \omega^2 = \omega^2(t) \\ \Rightarrow \partial L / \partial t \neq 0 \\ \Rightarrow \text{energy not const., here} \\ \text{Mathieu's Egn.} \\ \text{instability for } \gamma \approx 2\omega_0 \text{ resonance} \end{cases}$$

then true nonlinear problems,

3) Nonlinear Oscillator - Conservative

i.e. $\ddot{x} + \omega_0^2 x + \epsilon x^3 = 0$ Duffing's Equation

↑
anharmonic
nonlinear } term

Here, conservative system, with bounded phase space trajectories

i.e. $\frac{d}{dt} \dot{x}^2 + F(x) = E$ \rightarrow quasi-periodic orbits bounded

↓
 $\omega_0^2 x^2/2 + \epsilon x^4/4$

\rightarrow secularities in perturbation theory

$$\ddot{x} + \omega_0^2 x + \epsilon x^3 = 0$$

$\omega^2 = \omega_0^2 \langle x^2 \rangle + \epsilon \langle x^4 \rangle$ $x^{(0)} \sim a \sin \omega_0 t$

$\langle x^2 \rangle \sim a^2 \sin^2 \omega_0 t$ $\epsilon x^3 \sim a^3 \sin^3 \omega_0 t \sin \omega_0 t$

$\sim a^3 \sin \omega_0 t$ (on resonance)

\Rightarrow secularities $\sim t \cos \omega_0 t$? ? ?

\Rightarrow nonlinear frequency shift / \rightarrow removes secularities!

\rightarrow simplest example of Reductive P.T.

$$\Rightarrow \omega = \omega_0 + \omega^{(2)}$$

amplitude dep.

i.e. $\omega_0^2 x + \epsilon x^3$

$$\rightarrow (\omega_0^2 + \epsilon \overline{x^2}) x$$

2nd-order correction

$$\sim \frac{3\epsilon a^2}{8\omega_0}$$

$$\sim (\omega_0^2 + (\#)\epsilon a^2) x$$

↳ effective ω -correction!

4.) Forced Nonlinear Oscillator

i.e. $\ddot{x} + \alpha \dot{x} + \omega_0^2 x + \epsilon x^3 = f(t)$

\uparrow NL \uparrow forcing

Forced Duffing's Equation

Here \rightarrow role of nonlinearity/anharmonicity in resonance

\rightarrow i.e. contrast: forced SHO ($\epsilon=0$)

$$f_{ext} = f e^{i\omega t}$$

$$x = B e^{i\omega t}$$

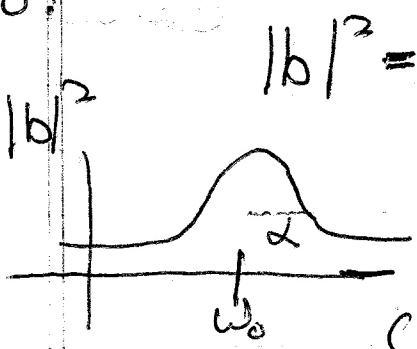
dissipation resolves resonance

$$B = b e^{i\omega t}$$

$$b = f/m \left[(\omega_0^2 - \omega^2)^2 + 4\alpha^2 \omega^2 \right]^{1/2}$$

$$\tan \delta = 2\alpha\omega / (\omega^2 - \omega_0^2)$$

so:



$$|b|^2 = \frac{f^2}{m^2} \left[(\omega_0^2 - \omega^2)^2 + 4\alpha^2 \omega^2 \right] \rightarrow \frac{f^2}{4m^2 \omega_0^2}$$

resonance $\rightarrow \omega_0$
linewidth $\rightarrow \alpha$

$$\left(\frac{1}{(\omega_0 - \omega)^2 + \alpha^2} \right)^*$$

(near resonance)

(symmetric)

but for NL forced SHO;

→ nonlinear frequency shift enters resonance!
here:

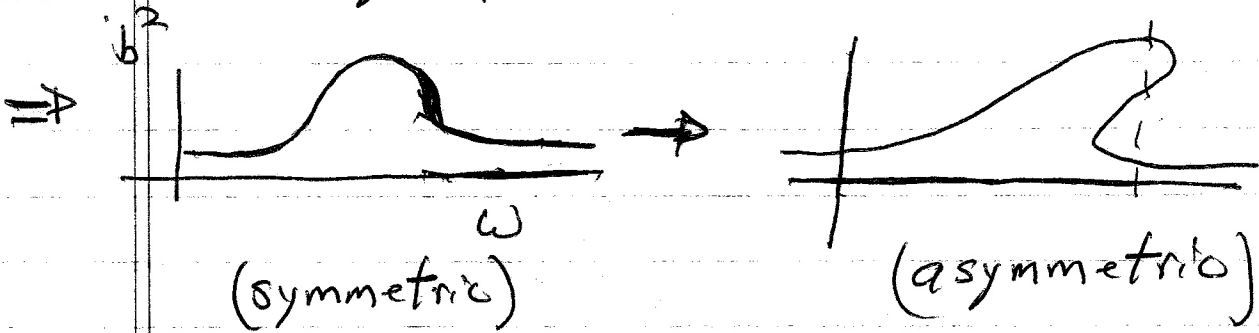
$$\omega - \omega_0 \rightarrow \omega - \omega_0 - Kb^2$$

{ amplitude-dependent
shift

→

$$b^2 = \left(\frac{F^2}{4m^2\omega_0^2} \right) \frac{1}{[(\omega - \omega_0 - Kb^2)^2 + d^2]}$$

→ cubic eqn. for b^2



S-curve, bifurcations,
mode jumping, etc.

6. Nonlinear Oscillator with \pm Dissipation

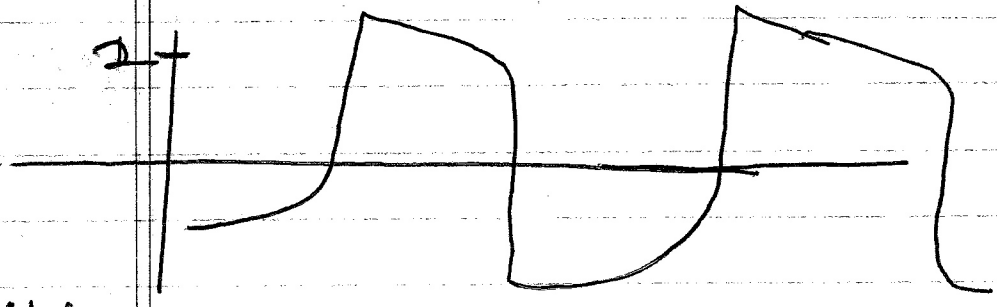
e.g. Van-der-Pol oscillator / Eqn.:

$$\ddot{x} + \alpha(x^2 - 1)\dot{x} + \omega_0^2 x = 0$$

key feature: nonlinear
variable sign > dissipation
(friction)

$$\Rightarrow \begin{cases} x^2 < 1 \rightarrow \alpha(x^2 - 1) < 0 \Rightarrow \begin{cases} \text{negative dissipation} \\ \text{instability} \end{cases} \\ x^2 > 1 \rightarrow \alpha(x^2 - 1) > 0 \Rightarrow \begin{cases} \text{positive dissipation} \\ \rightarrow \text{stability / saturation} \end{cases} \end{cases}$$

\Rightarrow bursty behavior, characteristic of limit cycle (at $x^2 = 1$)



Note:

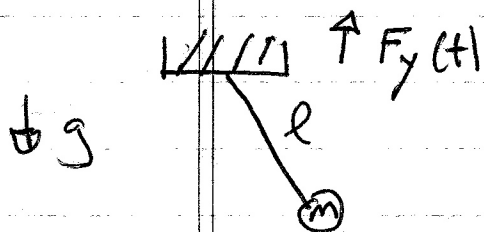
In contrast to previous examples, system is fundamentally dissipative = nonlinearity in friction.

System/Eqn.	Nonlinearity	External
Mathieu (Parametric)	None	$\omega_0^2 = \omega_0^2(t)$ Fast
Duffing (Freq. Shift)	Potential	None
Forced Duffing (Bi-stability/Bifurcation)	Potential	Forcing
Van-der-Pol (Limit Cycle)	Friction	None

Parametric Resonance and Instability

2. Parametric Instability

→ consider pendulum with support acted on by vertical force



↓ + → down

so $g \Rightarrow g - F_y(t)/m$

$$\therefore \ddot{\theta} = \ddot{\theta} + \frac{g}{l} \theta \rightarrow \ddot{\theta} + \left(\frac{g}{l} - \frac{a(t)}{l} \right) \theta = 0$$

let $a(t) = a_0 \cos(\alpha t)$

$$\Rightarrow \ddot{\theta} + \omega_0^2 \theta - \frac{a_0 \cos(\alpha t)}{l} \theta = 0$$

∴ of Mathieu's equation genre, i.e.

$$\ddot{x} + \omega_0^2 (1 + a \cos(\gamma t)) x = 0$$

↓
 $\omega^2 = \omega^2(t)$, hence parametric oscillator

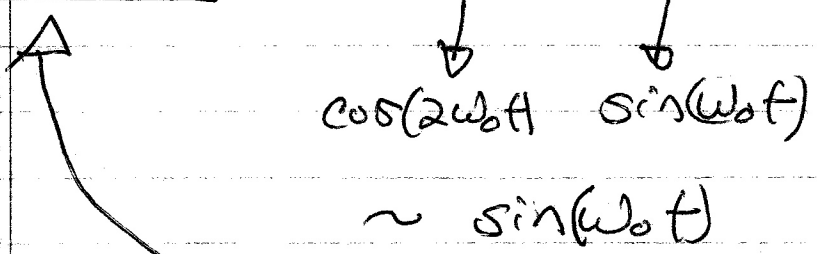
Parametric oscillator $\leftrightarrow \omega^2(t)$ periodic

→ Some observations:

a) informal - consider what might happen?

for instability, observe can produce secularly if $\gamma \sim 2\omega_0$ via beat at fundamental

$$\ddot{x} + \omega_0^2 x + a \cos(\gamma t) \omega_0^2 x = 0$$



resonant drive of fundamental oscillator

\Rightarrow secularly \rightarrow instability (why?)

\therefore Solution of oscillator at ω_0 beats with parameter oscillation \Rightarrow secularly

\therefore parametric resonance at/near $\gamma \sim 2\omega_0$ (twice fundamental)

Note: here $\omega^2 = \omega^2(t) \Rightarrow \partial L / \partial t \neq 0$ energy not conserved

\Rightarrow work done on system (e.g. LGM oscillating pendulum support)

\Rightarrow source of energy for instability

What is relation of this to 3-mode parametric instability calculation (200A)?

b) Formal (Floquet theory) $\left\{ \begin{array}{l} \text{What} \\ \text{Mathematics} \\ \text{Predicts} \end{array} \right.$
(What type solution possible?)

- $\omega(t)$ periodic, with period $T = 2\pi/\gamma$

$$\therefore \begin{cases} \omega(t+T) = \omega(t) \\ \text{eqn. invariant under } t \rightarrow t+T \end{cases}$$

\therefore if $x_1(t), x_2(t)$ are 2 independent solutions of basic eqn.

$\Rightarrow x_1(t), x_2(t)$ must transform to linear combinations of themselves upon $t \rightarrow t+T$ (linear eqn.)

and

can choose x_1, x_2 s/t

$$\begin{aligned} x_1(t+T) &= \mu_1 x_1(t) \\ x_2(t+T) &= \mu_2 x_2(t) \end{aligned}$$

(here "can choose" means can diagonalize transformation matrix)

\rightarrow most general functions having this property are:

$$\begin{aligned} x_1(t) &= \mu_1^{t/T} \pi_1(t) \\ x_2(t) &= \mu_2^{t/T} \pi_2(t) \end{aligned}$$

$$\left\{ \begin{array}{l} \text{where:} \\ \pi_i(t+T) = \pi_i(t) \end{array} \right.$$

second, observe since linear equation
⇒ Wronskian constant

$$\dot{x}_2 x_1 - \dot{x}_1 x_2 = \text{const.}$$

$$\begin{matrix} x_2 (\ddot{x}_1 + \omega^2(t) x_1) = 0 \\ x_1 (\ddot{x}_2 + \omega^2(t) x_2) = 0 \end{matrix} \Rightarrow \frac{d}{dt} (x_2 \dot{x}_1 - \dot{x}_2 x_1) = 0$$

but

$$W(x_1, x_2) = (U_1, U_2)^{-1} W(x_1(t+T), x_2(t+T))$$

c.i.e. consider time translation by T

$$\rightarrow \left\{ \begin{array}{l} U_1, U_2 = 1 \\ W(x_1, x_2) = \begin{pmatrix} U_2^{t/T} & U_1^{t/T} \\ -U_1^{t/T} & U_2^{t/T} \end{pmatrix} \end{array} \right.$$

Can also observe: $\begin{pmatrix} e^{(\ln U_2) t/T} & e^{(\ln U_1) t/T} \\ -e^{(\ln U_1) t/T} & e^{(\ln U_2) t/T} \end{pmatrix}$

1) coeffs in oscillator eq, so
 $x(t)$ an integral $\rightarrow x^*$ a solution

⇒

2) U_1, U_2 same as U_1^*, U_2^*
c.i.e.

$$\begin{matrix} U_1 = U_2^* \\ U_2 = U_1^* \end{matrix} \quad \text{or} \quad \begin{matrix} U_1 = U_1^* \\ U_2 = U_2^* \end{matrix} \quad \left. \vphantom{\begin{matrix} U_1 = U_2^* \\ U_2 = U_1^* \end{matrix}} \right\} \begin{array}{l} \text{both} \\ \text{real} \end{array}$$

if ①, $U_1, U_2 = 1 \Rightarrow U_1 = 1/U_1^* \Rightarrow |U_1|^2 = |U_2|^2 = 1$
② $U_2 = 1/U_1^* \Rightarrow |U_1|^2 = |U_2|^2 = 1$
(trivial)

if ② $\mu_1 \mu_2 = 1$; μ_1, μ_2 real \Rightarrow

$$\Rightarrow x_1(t) = \mu^{t/T} \pi_1(t), \quad x_2(t) = \mu^{-t/T} \pi_2(t)$$

i.e. $\left. \begin{array}{l} \uparrow \text{ increasing} \\ \uparrow \text{ decreasing} \end{array} \right\} \text{ solution} \Rightarrow \left\{ \begin{array}{l} \text{Parametric} \\ \text{Instability} \end{array} \right.$

[N.B. Exponential, not secular, growth]!

\Rightarrow "true" instability is possible

\rightarrow Some Calculation (as basic structure of the solution established).

Consider Mathieu's eqn:

$$\ddot{x} + \omega_0^2 [1 + h \cos(2\omega_0 t + \epsilon)t] x = 0$$

bounds on ϵ for instability?

For solution, SHO \Rightarrow

$$x = a \cos(\omega_0 t) + b \sin(\omega_0 t)$$

so, in spirit of multiple-time-scale P.T.
(i.e. $\omega^2(t)$ enters via $h \ll 1 \Rightarrow$ expect
slow time scale variation of coefficients)

$$x = a(t) \cos[(\omega_0 + \epsilon/2)t] + b(t) \sin[(\omega_0 + \epsilon/2)t]$$

\downarrow
coeffs become slowly varying

Plugging of in:

$$\ddot{x} = (a(t) \cos[(\omega_0 + \epsilon/2)t]) + o.T. \quad \text{other term}$$

$$= -(\omega_0 + \epsilon/2)^2 a(t) \cos[\] - 2(\omega_0 + \epsilon/2) \dot{a}(t) \sin[\] + \ddot{a} \cos[\] + o.T.$$

neglect \ddot{a} , \ddot{b} as h.o. in slowness (recall amplitude eqn. deriv.)

\Rightarrow ω_0^2 term, only

$$-(\omega_0 + \epsilon/2)^2 a(t) \cos[\] - 2\dot{a}(t) (\omega_0 + \epsilon/2) \sin[\]$$

$$-(\omega_0 + \epsilon/2)^2 b(t) \sin[\] + 2\dot{b}(t) (\omega_0 + \epsilon/2) \cos[\]$$

$$+ \omega_0^2 [a(t) \cos[\] + b(t) \sin[\]]$$

$$+ \omega_0^2 h \cos(2\omega_0 t) [a(t) \cos[\] + b(t) \sin[\]]$$

$$= 0$$

Now; - neglect $O(\epsilon^2)$ terms \Rightarrow only $\omega_0 \epsilon$ term survives.

- observe $\cos[(\omega_0 + \epsilon/2)t] \cos[2\omega_0 t]$

$$= \frac{1}{2} \cos[3(\omega_0 + \epsilon/2)t] + \frac{1}{2} \cos[(\omega_0 + \epsilon/2)t]$$

Resonant
Contribution is
interesting one here.

{ fast oscillation } \rightarrow Resonant with Fundamental
(i.e. expect h.o. in h)

$$\Rightarrow$$

$$-\omega_0 \epsilon (a(t) \cos[\] + b(t) \sin[\])$$

$$- 2 \dot{a} (\omega_0 + \epsilon/2) \sin[\] + 2 \dot{b} (\omega_0 + \epsilon/2) \cos[\]$$

$$+ \frac{\omega_0^2 h}{2} [a(t) \cos[\] - b(t) \sin[\]]$$

$$= 0$$

Regrouping coeffs. $\cos[\]$, $\sin[\]$;

$$\sin[\] (-2\omega_0 \dot{a} - b\omega_0 \epsilon - \omega_0^2 h b/2)$$

$$+ \cos[\] (2\dot{b}\omega_0 - a\epsilon\omega_0 + \frac{1}{2} h\omega_0^2 a) = 0$$

$$\Rightarrow$$

$$(2\omega_0) \dot{a} + (\omega_0 \epsilon) b + (\frac{\omega_0^2 h}{2}) b = 0$$

$$(2\omega_0) \dot{b} - (\omega_0 \epsilon) a + (\frac{\omega_0^2 h}{2}) a = 0$$

$$\Rightarrow$$

$$\begin{cases} \dot{a} + (\epsilon/2) b + (\omega_0 h/4) b = 0 \\ \dot{b} - (\epsilon/2) a + (\omega_0 h/4) a = 0 \end{cases}$$

Basic
System
of
Eqs for
Amplitude
Variation

$a(t) = a_0 e^{st}$
 $b(t) = b_0 e^{st}$
exponentially growing/damping solutions

⇒

$$s a_0 + (\epsilon/2 + \omega_0 h/4) b_0 = 0$$

$$\left(-\frac{\epsilon}{2} + \frac{\omega_0 h}{4}\right) a_0 + s b_0 = 0$$

$$s^2 = \frac{\omega_0^2 h^2}{16} - \frac{\epsilon^2}{4} = \frac{1}{4} \left(\frac{\omega_0^2 h^2}{4} - \epsilon^2 \right)$$

⇒ Parametric instability criterion

Growth rate

Observe:

- instability for:

$$\epsilon^2 = (\gamma - 2\omega_0)^2 < \frac{\omega_0^2 h^2}{4}$$

$\omega_0 \rightarrow$ Fundamental
 $\gamma \rightarrow$ parametric variation freq.

amplitude of variation
 $h^2 > 4(\gamma - \omega_0)^2 / \omega_0^2$

↑
 i.e. sufficiently close to resonance
 ⇒ growth.

for $(\gamma - 2\omega_0)^2 > \omega_0^2 h^2 / 4 \rightarrow$ oscillation

- amplitude of $\omega_0^2(t)$ variation sets proximity threshold

integer

more generally, can show when $n\gamma = 2\omega_0$
 \Rightarrow parametric resonance. of course, higher
 $n \Rightarrow$ resonance region $\sim h^n$

- with friction, find threshold for instability:

c.e. $(\gamma - 2\omega_0)^2 < \left[\left(\frac{1}{2} h \omega_0 \right)^2 - 4\alpha^2 \right]$

\uparrow
friction coeffs

c.e. P.I. growth must be damped!
 Friction raises required h .

- Pumping on swing

\rightarrow "pumping" \rightarrow change of I

$$\ddot{\theta} + \frac{mgl}{I(t)} \theta = 0$$

$$I(t) = I_0 + \epsilon I_1(t)$$

$$\ddot{\theta} + \frac{g}{l} \theta + \frac{\epsilon g}{l} \frac{\Delta I(t)}{I} \theta = 0$$

$$+ \alpha \dot{\theta}$$

need pump twice per cycle

Nonlinear Oscillators - Conservative

→ Here, concerned with $\left\{ \begin{array}{l} \text{nonlinear} \\ \text{conservative} \end{array} \right\}$ oscillator systems, usually small perturbations about/away from the SHO

Prototype: Duffing's Equation

$$\ddot{x} + \omega_0^2 x + \epsilon x^3 = 0$$

SHO, ω_0

↓
NL, anharmonic term

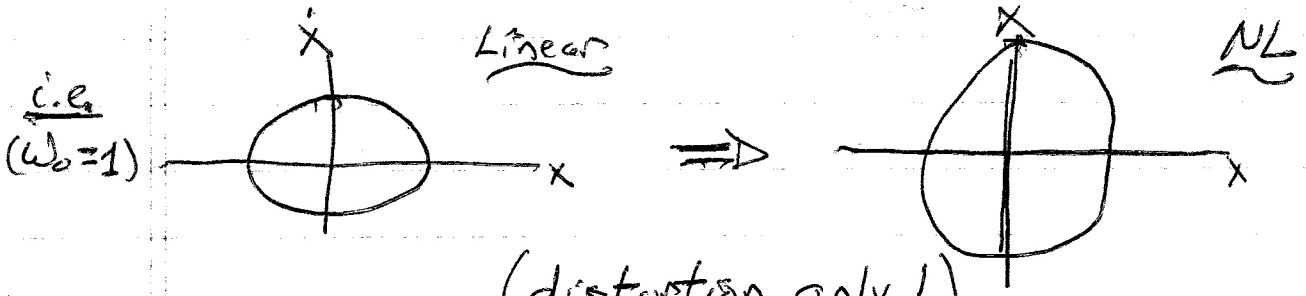
Observe:

$$H = \frac{1}{2} \dot{x}^2 + \frac{\omega_0^2 x^2}{2} + \frac{\epsilon x^4}{4}$$

$$\text{so } V(x) = \frac{1}{2} \omega_0^2 x^2 + \frac{\epsilon x^4}{4} \quad (V(x)' > 0)$$

- natural question to ask re: $\omega = \omega(\epsilon)$?
i.e. evolution of periodic/quasi-periodic orbits upon perturbation.

Now note: $V(x)$ bounded \Rightarrow phase space contours (from below) trajectory "contained"



(distortion only!)
so orbit must be bounded!

- A clue;

Observe: $\omega^2 = \frac{1}{2} \omega_0^2 \langle x^2 \rangle / \langle x^2 \rangle / 2$
 for SHO, $\langle \dots \rangle = \frac{1}{T} \int_0^T \dots dt$

\Rightarrow might expect, taking $T = 2\pi/\omega_0$

$$\omega^2 = \left(\frac{1}{2} \omega_0^2 \langle x^2 \rangle + \frac{\epsilon}{4} \langle x^4 \rangle \right) / \frac{\langle x^2 \rangle}{2}$$

(solve with SHO 'trial' fctn.)

Now; $\langle x^4 \rangle = \langle a^4 (\cos \omega_0 t)^4 \rangle$
 $= a^4 (3/4)$

$$\boxed{\omega^2 = \omega_0^2 + \frac{3\epsilon}{4} a^2}$$

\rightarrow amplitude dependent frequency! (NL frequency shift)

i.e. $\omega = \omega_0 \rightarrow$

$\Delta \rightarrow$ (nearly) correct result.

$\omega = \omega(\omega_0, \epsilon, a)$ \rightarrow frequency becomes amplitude dependent.

Systematics - Computational Procedure

- expand in ϵ ! ~~!!~~ (surprise !!)

$$\ddot{x} + \omega_0^2 x + \epsilon x^3 = 0 \quad ; \quad x(0) = 0$$

$$x = x^{(0)} + \epsilon x^{(1)} + \dots$$

$$\therefore O(\epsilon^0) : \quad \ddot{x}^{(0)} + \omega_0^2 x^{(0)} = 0$$

$$x^{(0)} = a \sin(\omega_0 t)$$

$$O(\epsilon^1) : \quad \ddot{x}^{(1)} + \omega_0^2 x^{(1)} = -\epsilon (x^{(0)})^3$$

but $(x^{(0)})^3 = a^3 \sin^3 \omega_0 t$

$$= \frac{a^3}{4} (3 \sin \omega_0 t + \sin 3\omega_0 t)$$

$$\Rightarrow \ddot{x}^{(1)} + \omega_0^2 x^{(1)} = -\frac{\epsilon a^3}{4} (3 \sin \omega_0 t + \sin 3\omega_0 t)$$

\downarrow
 resonant drive
 \Rightarrow secular!

yields $x^{(1)} = \text{homog} + C t \cos \omega_0 t$

\downarrow
 is this physical?

$$\Rightarrow -\omega_0^2 c t \cos \omega_0 t - 2C \omega_0 \sin \omega_0 t + \omega_0^2 c t \cos \omega_0 t = -\frac{\epsilon a^3}{4} (3 \sin \omega_0 t)$$

$$C = \frac{3\epsilon}{8\omega_0} ; \quad \epsilon \sim (\text{freq})^2, \text{ dimensionally}$$

$$\therefore x = a \sin \omega_0 t + \frac{3}{8} \frac{\epsilon}{\omega_0} t \cos \omega_0 t + \dots$$

secularity \rightarrow $|x|$ diverges linearly in time

Unphysical \Rightarrow recall closed phase space trajectories!

What's going on?

Aside: A trivial example $\left| \begin{matrix} P \\ \circ \\ \circ \\ \circ \end{matrix} \right| \left| \begin{matrix} \\ \circ \\ \circ \\ \circ \end{matrix} \right|$

$$\ddot{x} + (\pm \epsilon)^2 x = 0 ; \quad x(0) = 0 \\ \dot{x}(0) = 1$$

if expand in ϵ ;

$$\ddot{x} + x + 2\epsilon x + \epsilon^2 x = 0 \quad (\text{exactly solvable})$$

$$\epsilon^{(0)} ; \quad \ddot{x}^{(0)} + x^{(0)} = 0 ; \quad x^{(0)} = \sin t$$

$$x''(t) + x'(t) = -2x(t)$$

$$\Rightarrow x^{(h)} = c t \left[\cos t \right] + \text{homog.}$$

$$2c \sin t - c t \cos t + c t \cos t = -2 \sin t$$

$$c = 1$$

• secular! ? ? ?

This is clearly idiotic, since we all know

$$x(t) = \sin[(1+\epsilon)t] \quad \text{trivially solves the problem!}$$

Moral of this story:

- ② - to avoid secular, must allow frequency shift; i.e. here $\omega = 1 \rightarrow \omega = 1 + \epsilon$
gives all a warm, fuzzy.....
- ① - secular results from breakdown of naive expansion in ϵ at long times, observe:

i.e. $x(t) = \sin[(1+\epsilon)t]$,

Taylor expansion in $\epsilon \Rightarrow \approx \sin t + \epsilon t \cos t$

\hookrightarrow secularity \rightarrow artifact of expansion

→ The Fix:

- admit nonlinear frequency shift

(i.e. method of Poincaré-Linstedt)

Top of Reductive P.T. ice-berg ----

- trick is to:

a) expand x, ω on equal footing

$$x = x^{(0)} + \epsilon x^{(1)} + \dots$$

$$\omega = \omega^{(0)} + \epsilon \omega^{(1)} + \dots$$

and use/choose $\omega^{(1)}$, etc. to cancel secularities

i.e. " solvability condition " \leftrightarrow secularities removed.

i.e. $\frac{d^2 x}{dt^2} + x + \epsilon x^3 = 0$

usual long time behavior

now! $t = S (1 + \epsilon \omega_1 + \epsilon^2 \omega_2 + \dots)$

time param.

and $x = x_0 + \epsilon x_1 + \epsilon^2 x_2 + \dots$

Improved Details

$$\ddot{X} + \omega_0^2 X + \epsilon X^3 = 0$$

$$\text{now } \left\{ \begin{array}{l} X = X^{(1)} + X^{(2)} + X^{(3)} \\ \quad \quad \quad \downarrow \\ \quad \quad \quad \epsilon \\ \quad \quad \quad \epsilon(\epsilon) \quad \rightarrow O(\epsilon^2) \\ \\ \omega = \omega_0 + \omega^{(1)} + \omega^{(2)} + \dots \\ \quad \quad \quad \rightarrow O(\epsilon) \text{ (shift)} \end{array} \right.$$

$$\text{and } X_0 = a \cos \omega t$$

\rightarrow note ω (not ω_0)

Now; re-write:

$$\frac{\omega_0^2}{\omega^2} \ddot{X} + \omega_0^2 X = -\epsilon X^3 - \left(1 - \frac{\omega_0^2}{\omega^2}\right) \ddot{X}$$

extra terms assure LHS = 0, in lowest order.

\Rightarrow

$$\begin{aligned} & \frac{\omega_0^2}{\omega^2} \left(-\omega^3 X^{(1)} + \ddot{X}_2 \right) + \omega_0^2 \left(X^{(1)} + X_2 \right) \\ &= -\epsilon \left(X^{(1)} + X^{(2)} \right)^3 - \left(\frac{\omega_0^2}{\omega^2} - 1 \right) \left(-\omega^3 X^{(1)} + \ddot{X}_2 \right) \end{aligned}$$

$$\begin{aligned}
 & -\omega_0^2 \cancel{X^{(4)}} + \ddot{X}_2 + \omega_0^2 \cancel{X^{(4)}} + \omega_0^2 X_2 \\
 & = -\epsilon X^{(4)3} - (-\omega^2 + \omega_0^2) X^{(4)} \\
 & \quad - \left(1 - \frac{\omega_0^2}{\omega^2} \right) \ddot{X}_2
 \end{aligned}$$

⇒

$$\begin{aligned}
 \ddot{X}_2 + \omega_0^2 \ddot{X}_2 & = -\epsilon X^{(4)3} + (\omega^2 - \omega_0^2) X^{(4)} \\
 & \quad \underbrace{\quad}_{0(\epsilon)} \quad \underbrace{\quad}_{0(\epsilon)} \\
 & \quad \underbrace{\quad}_{\text{h.o.}} \\
 & \quad - \left(1 - \frac{\omega_0^2}{\omega^2} \right) \ddot{X}_2
 \end{aligned}$$

$$\omega^2 = \omega_0^2 + 2\omega_0\omega, \epsilon \quad ; \quad \omega^2 - \omega_0^2 = 2\omega_0\omega, \epsilon$$

$$\begin{aligned}
 X^{(4)3} & = a^3 \cos^3 \omega t = a^3 \left[\frac{1}{2} + \frac{\cos 2\omega t}{2} \right] \cos \omega t \\
 & = a^3 \left[\frac{\cos \omega t}{2} + \frac{1}{2} \left(\frac{1}{2} \right) (\cos 3\omega t + \cos \omega t) \right] \\
 & = a^3 \left[\frac{3}{4} \cos \omega t + \frac{1}{4} \cos 3\omega t \right]
 \end{aligned}$$

$$\ddot{x}_2 + \omega_0^2 \dot{x}_2 = -\epsilon \left[a^3 \left(\frac{3}{4} \cos \omega t + \frac{1}{4} \cos 3\omega t \right) \right] + 2\omega_0 \omega_1 \epsilon a \cos \omega t$$

Now to dodge the secularity, need cancel all resonant terms on RHS

⇒

$$\ddot{x}_2 + \omega_0^2 \dot{x}_2 = -\epsilon \left[\cos \omega t \right] \left[\frac{3a^3 - 2\omega_0 \omega_1 a}{4} \right] - \epsilon \frac{a^3}{4} \cos 3\omega t$$

$$\omega_1 = \frac{3}{8} a^2 / \omega_0$$

and

$$\omega = \omega_0 + \epsilon \left(\frac{3}{8} a^2 / \omega_0 \right) + \dots$$

NL Frequency shift!

and crank \Rightarrow

$$x^{(2)} = \frac{-1}{2} \left(\frac{G a^3}{16 \omega_0^2} \right) \cos 3\omega t.$$

Point:

- need get frequencies correct to avoid unphysical resonances, secularities...
- frequency correction \Rightarrow NL frequency shift

N.B. : compare:

- exact: $\omega_1 = 3/8 a^2 / \omega_0$

- rough: $\omega_1 = 3/4 a^2 / \omega_0$

Moral: Use frequency shift to eliminate secularly causing term on RHS

so:

$$\frac{d^3}{ds^2} (x_0 + \epsilon x_1 + \epsilon^2 x_2 + \dots) + \left[(x_0 + \epsilon x_1 + \epsilon^2 x_2 + \dots) + \epsilon (x_0 + \epsilon x_1 + \epsilon^2 x_2 + \dots)^3 \right] (1 + \epsilon \omega_1 + \epsilon^2 \omega_2 + \dots)^2 = 0$$

$$O(\epsilon^0): \quad \frac{d^2 x_0}{ds^2} + x_0 = 0$$

$$O(\epsilon^1): \quad \frac{d^2 x_1}{ds^2} + x_1 = -x_0^3 + 2\omega_1 x_0 = 0$$

$$O(\epsilon^2): \quad \frac{d^2 x_2}{ds^2} + x_2 = 3x_0^2 x_1 - 2\omega_1 (x_1 + x_0^2) - (\omega_1^2 + 2\omega_2) x_0$$

etc.

Now,

$$O(\epsilon^0): \quad \frac{d^2 x_0}{ds^2} + x_0 = 0$$

→ phase

$$x_0 = a \cos(s + \phi)$$

$\hat{u}(E^0)$:

$$\frac{d^2 X_1}{ds^2} + X_1 = -X_0^3 - 2\omega_1 X_0$$

$$= -a^3 \cos^3(s+\phi) - 2\omega_1 a \cos(s+\phi)$$

etc

Aut

$$\cos^3(s+\phi) = \cos(s+\phi) \left[\frac{1}{2} + \frac{1}{2} \cos[2(s+\phi)] \right]$$

$$= \frac{1}{2} \cos(s+\phi) + \frac{1}{2} \cos(s+\phi) \cos[2(s+\phi)]$$

$$= \frac{\cos(s+\phi)}{2} + \frac{1}{4} \cos[3(s+\phi)] + \frac{1}{4} \cos[s+\phi]$$

$$= \frac{3}{4} \cos(s+\phi) + \frac{1}{4} \cos(3(s+\phi))$$

\Rightarrow

①

②

$$\frac{d^2 X_1}{ds^2} + X_1 = -a^3 \left(\frac{3}{4} \cos(s+\phi) + \frac{1}{4} \cos(3(s+\phi)) \right)$$

③

$$-2\omega_1 a \overset{\cos}{\uparrow} [s+\phi]$$

so

①, ③ $\sim \cos(st + \phi)$

resonates with RHS \Leftrightarrow
will drive secularity.

② $\sim \cos[3(st + \phi)]$

non-secular drive \rightarrow harmless

$$\frac{d^2 X_1}{ds^2} + X_1 = \left[-\frac{3}{4} a^2 - 2\omega_1 \right] a \cos(st + \phi) + \frac{1}{4} \cos(3(st + \phi))$$

so $\left\{ \omega_1 = -\frac{3}{8} a^2 \right.$ removes secularity !!

$t = s \left(1 - \frac{3}{8} \epsilon a^2 + \dots \right)$

$\omega = \omega_0 \left[1 + \frac{3}{8} \epsilon a^2 + \dots \right]$

NL
 frequency
 shift \downarrow

i.e.
 $s = t / \left(1 - \frac{3}{8} \epsilon a^2 \right) \approx t \left(1 + \frac{3}{8} \epsilon a^2 \right)$

and $X_1 = \frac{1}{3} a^3 \cos[3(st + \phi)]$

→ Frequency shift + recovers guess estimate 24

→ example 2 - Nonlinear Klein Gordon Eqn.

KG eqn: (Scalar Field)

$$\frac{1}{c_0^2} \frac{\partial^2 \phi}{\partial t^2} - \frac{\partial^2 \phi}{\partial x^2} + m^2 \phi = 0$$

what physical system does this describe?

→ dispersion relation:

$$\omega^2 = c_0^2 k^2 + m^2$$

(Calc plasma wave)

$$\mathcal{L} = \frac{\dot{\phi}^2}{2c_0^2} - \frac{1}{2} \left[\frac{\partial \phi}{\partial x} \right]^2 - \frac{m^2 \phi^2}{2}$$

for nonlinearity:

$$U = \frac{1}{2} \left(\frac{\partial \phi}{\partial x} \right)^2 + \frac{m^2 \phi^2}{2} + \frac{\alpha \phi^4}{4}$$

⇒

$$\frac{\partial^2 \phi}{\partial t^2} - c_0^2 \frac{\partial^2 \phi}{\partial x^2} + m^2 \phi = -\alpha \phi^3$$
physics?

Mean solution $\left\{ \begin{array}{l} \omega^2 = c_0^2 k^2 + m^2 \\ \phi = \phi(x-ct) = \phi(\theta) \end{array} \right.$

→ look for wave train solutions

sneaky → converts to ODE problem.

so, for NL problem

$$(c^2 - c_0^2) \phi'' + m^2 \phi = -\alpha \phi^3$$

∴ expect nonlinearity will produce nonlinear phase velocity shift!

$$\Rightarrow c = c_0^{(a)} + \alpha^2 c_2 + \dots \quad \left. \begin{array}{l} \\ \\ \end{array} \right\} \begin{array}{l} \text{by correspondence} \\ \text{with Duffing} \\ (a \leftrightarrow c) \end{array}$$

$$\phi = a \phi_1 + \alpha^3 \phi_3 + \dots$$

n.b. here $\phi = \phi(\theta) \leftrightarrow$ wave train solution
∴ only parameter is c

$$\Rightarrow (c_0^{(a)2} - c_0^2) [a \phi_1'' + \alpha^3 \phi_3''] + m^2 [a \phi_1 + \alpha^3 \phi_3] = -\alpha [a \phi_1 + \alpha^3 \phi_3]^3 = 0$$

$$(c_0^{(a)2} - c_0^2) \phi_1'' + m^2 \phi_1 = 0 \quad O(a)$$

$$(c_0^{(a)2} - c_0^2) \phi_3'' + m^2 \phi_3 = -2 c_0^{(a)} c_2 \phi_1''$$

$$\phi_3 = O(a^3)$$

now, $\partial(\phi)$:

$$\phi_1 = \phi_0 \cos k\theta \quad (\text{contained in } \phi)$$

$$C^{(1)2} = \omega^2 + \frac{m^2}{k^2}$$

$\partial(C^{(1)})$:

$$\left((C^{(1)})^2 - \omega^2 \right) \phi_0'' + m^2 \phi_3$$

$$= -2 C^{(1)} C_2 (\cos k\theta)''$$

$$= -\cancel{\alpha} (\cos k\theta)''$$

$$= 2 C^{(1)} C_2 k^2 \cos k\theta$$

$$- \cancel{\alpha} \cos k\theta \left[\frac{1}{2} + \frac{1}{2} \cos 2\theta k \right]$$

$$= 2 C^{(1)} C_2 k^2 \cos k\theta - \frac{\cancel{\alpha}}{2} \cos k\theta$$

$$+ \frac{\cancel{\alpha}}{2} \cos 3k\theta - \frac{\cancel{\alpha}}{2} \cos k\theta$$

$$= 2 C^{(1)} C_2 k^2 \cos k\theta - \frac{3 \cancel{\alpha}}{4} \cos k\theta$$

$$+ \frac{\cancel{\alpha}}{4} \cos 3k\theta$$

so, to kill secularity:

$$2c^{(0)} c_2 k^2 = \frac{3 \alpha}{4}$$

$$c_2 = \frac{3 \alpha}{8 c^{(0)} k^2}$$

$$c = c^{(0)} + a^2 c_2$$

$$= \left(c_0^2 + \frac{m^2}{k^2} \right)^{1/2} \left[1 + \frac{3 \alpha a^2}{8 k^2 c^{(0)2}} \right]$$

$$c = \left(c_0^2 + \frac{m^2}{k^2} \right)^{1/2} \left[1 + \frac{3 a^2 \alpha}{8 (c_0^2 k^2 + m^2)} \right]$$

nonlinear speed change

shift

Key Point: \rightarrow By expanding ω in ϵ , method of Poincaré and Lindstedt introduces additional degrees of freedom, so one can remove secularities order-by-order in P.T.

\rightarrow essence of NL oscillator is NL frequency shift, i.e.

AS of asymptotics

$$\omega = \omega_0 + \frac{3\epsilon a^2}{8\omega_0}$$

Forced Anharmonic Oscillator - Mode Jumping

- here, consider next step in development

\Rightarrow forced Duffing's eqn.

$$\ddot{x} + \underbrace{2\lambda \dot{x}}_{\text{friction}} + \omega_0^2 x + \underbrace{\alpha x^2}_{\alpha=0, \text{initially}} + \beta x^3 = \frac{F \cos \omega t}{m}$$

issue: combination $\left\{ \begin{array}{l} \text{resonance} \\ \text{forcing} \\ \text{NL} \end{array} \right.$

Recall, for linear forced SHO; (see 31)

$$a^2 = F^2 / 4m^2 \omega_0^2 (\epsilon^2 + \lambda^2) \quad \text{is } \begin{cases} \text{amplitude} \\ \text{equation} \end{cases}$$

$$\epsilon = \omega - \omega_{res}; \quad \omega_{res} = \omega_0 \quad (\text{trivial})$$

Then for NL system: $\omega_{res} = \omega_0 + K a^2$
 (near primary, linear resonance) $K = \frac{3}{8} \frac{G}{\omega_0} \Rightarrow$ NL shift

$$a^2 = F^2 / 4m^2 \omega_0^2 ((\omega - \omega_0 - K a^2)^2 + \lambda^2)$$

$$\left\{ \begin{array}{l} a^2 [(\epsilon - K a^2)^2 + \lambda^2] = F^2 / 4m^2 \omega_0^2 \\ a(\epsilon) \text{ relation} \end{array} \right.$$

\Rightarrow cubic equation for a^2 !! (3 roots \rightarrow which?)

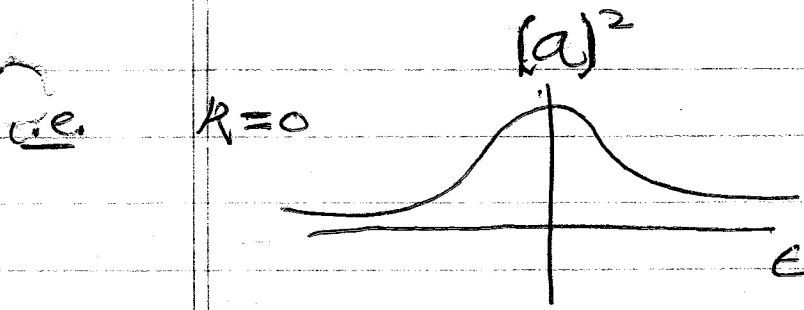
Observe:

- addition of NL \Rightarrow NL ω -shift \Rightarrow
non-trivial amplitude equation

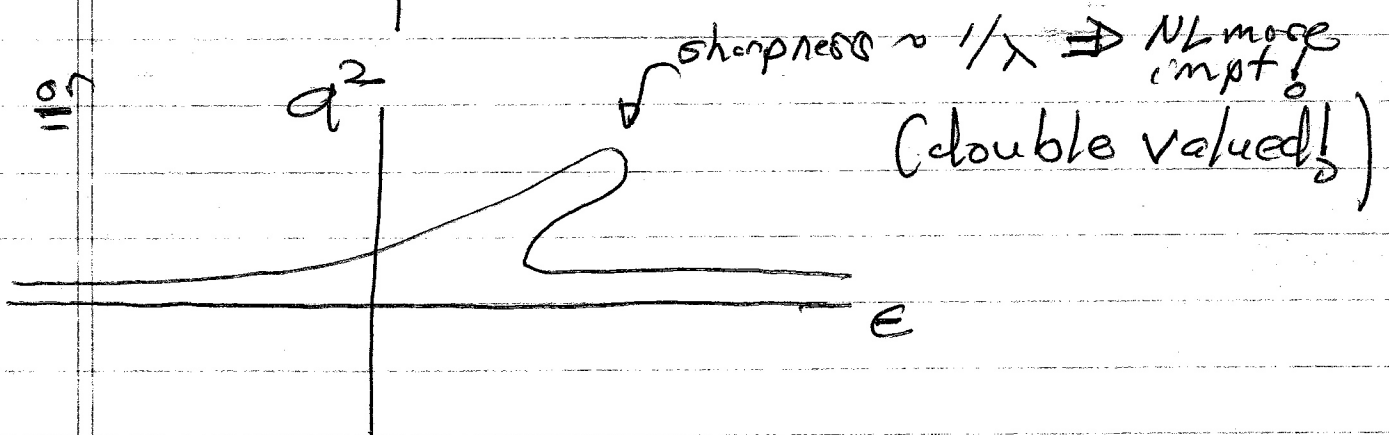
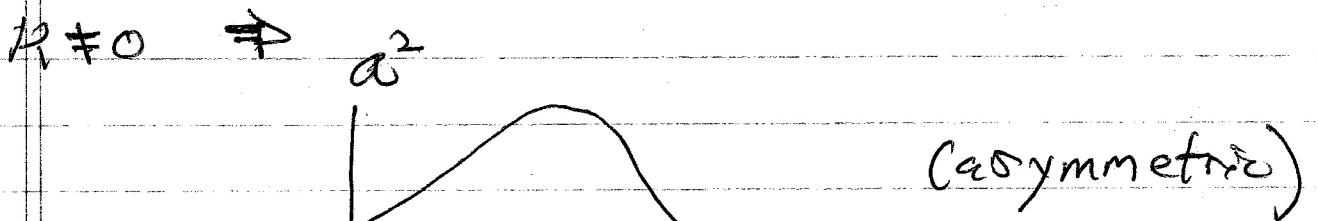
- for $\omega = \omega_{res}$ ($\epsilon = K a^2$), $a^2 = F^2 / 4m^2 \omega_0^2 \lambda^2$
 peak unchanged

but

asymmetry induced in resonance curve!



i.e. Lorentzian (symmetric)



Can the resonance curve be double-valued?
 - yes! \Rightarrow when $da^2/d\epsilon \rightarrow \infty$.

Now, amplitude equation: $\dot{a}(\epsilon)$

$$a^2 [(\epsilon - ka^2)^2 + \lambda^2] = f^2 / 4m^2\omega_0^2$$

$$\Rightarrow a^2 [\epsilon^2 - 2k\epsilon a^2 + (ka^2)^2 + \lambda^2] = f^2 / 4m^2\omega_0^2$$

Can re-write amplitude equation as:

$$F(a^2, \epsilon) = a^2 [\epsilon^2 - 2\epsilon(Ka^2) + (Ka^2)^2 + \lambda^2] - \frac{F^2}{4m^2\omega_0^2} = 0$$

So, for $\frac{da^2}{d\epsilon}$ on curve (defn.)

$$dF = 0 = (\partial F / \partial a^2) da^2 + (\partial F / \partial \epsilon) d\epsilon$$

$$\Rightarrow \frac{da^2}{d\epsilon} = - \frac{(\partial F / \partial \epsilon)}{(\partial F / \partial a^2)}$$

∴ $\partial F / \partial a^2 = 0$ for $\left\{ \begin{array}{l} \text{double valuedness} \\ \text{2 roots} \end{array} \right. \rightarrow$ though $\partial F / \partial a^2 = 0$ is coalescence pt. of two

$$F = K^2(a^2)^3 - 2\epsilon K(a^2)^2 + a^2(\epsilon^2 + \lambda^2) - \text{const.}$$

$$\partial F / \partial a^2 = 3K^2(a^2)^2 - 4\epsilon K a^2 + (\epsilon^2 + \lambda^2) = 0$$

$$\text{if } x = Ka^2$$

$$\partial F / \partial a^2 = 0 = 3x^2 - 4\epsilon x + (\epsilon^2 + \lambda^2)$$

$$\Rightarrow x = \frac{4\epsilon \pm \sqrt{16\epsilon^2 - 12(\epsilon^2 + \lambda^2)}}{6}$$

$$= \frac{2}{3}\epsilon \pm \frac{1}{3}(\epsilon^2 - 3\lambda^2)^{1/2}$$

$$\underline{\text{so}} \quad ka^2 = \frac{2}{3} \epsilon \pm \frac{1}{3} (\epsilon^2 - 3\lambda^2)^{1/2}$$

∴ → double valued-ness at inflection pt.
 { 2-root coalescence



i.e. when $\epsilon^2 = 3\lambda^2$, $\Rightarrow ka^2 = 2\epsilon/3$

in terms external force magnitude ($a(\epsilon)$ relation)

$$F^2 = 4m^2 \omega_0^2 a^2 \left[\epsilon^2 - 2\epsilon ka^2 + (ka^2)^2 + \lambda^2 \right]$$

$$\begin{cases} ka^2 = \frac{2}{3} \epsilon = \frac{2\sqrt{3}}{3} \lambda \\ \epsilon = \sqrt{3} \lambda \end{cases}$$

and plugging into $F^2 \Rightarrow$

$$F_{\text{crit}}^2 = 32m^2 \omega_0^2 \lambda^3 / 3\sqrt{3} |K|$$

{ i.e. $F > F_{\text{crit}}$
 required
 for inflection

→ for peak value of amplitude

$$\partial F / \partial \epsilon = 0 \quad (\Leftrightarrow da^2 / d\epsilon = 0)$$

$$\Rightarrow 2\epsilon - 2ka^2 = 0 \Rightarrow \epsilon = ka^2$$

$$\Rightarrow \omega - \omega_0 = ka^2, \text{ usual!}$$

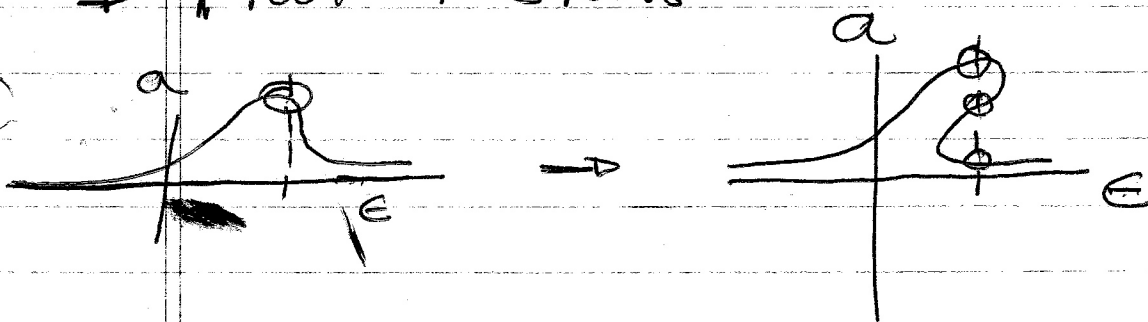
i.e. peak is at resonance (here $\omega = \omega_0 + Rq^2$), as usual.

→ What's Going On?

- for $f^3 > f_{crit}^2 = 32m^2\omega_0^2 \lambda^3 \sqrt{3} |K|$

⇒ inflection occurs → bifurcation

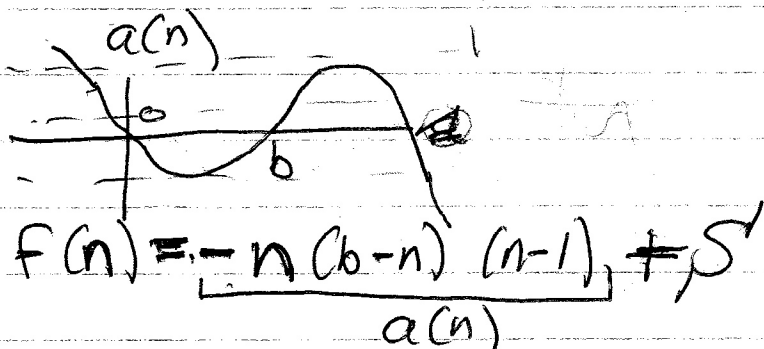
↓
 ⇒ 1 root → 3 roots



- of 3 roots ⇒ 2 stable
 1 unstable

i.e. demonstration

$$\frac{\partial n}{\partial t} = F(n)$$



$$\frac{\partial n}{\partial t} = 0 \Rightarrow S = +n(b-n)(n-1)$$

↓
 control ⇒ 3 or 1 roots, depending on S'

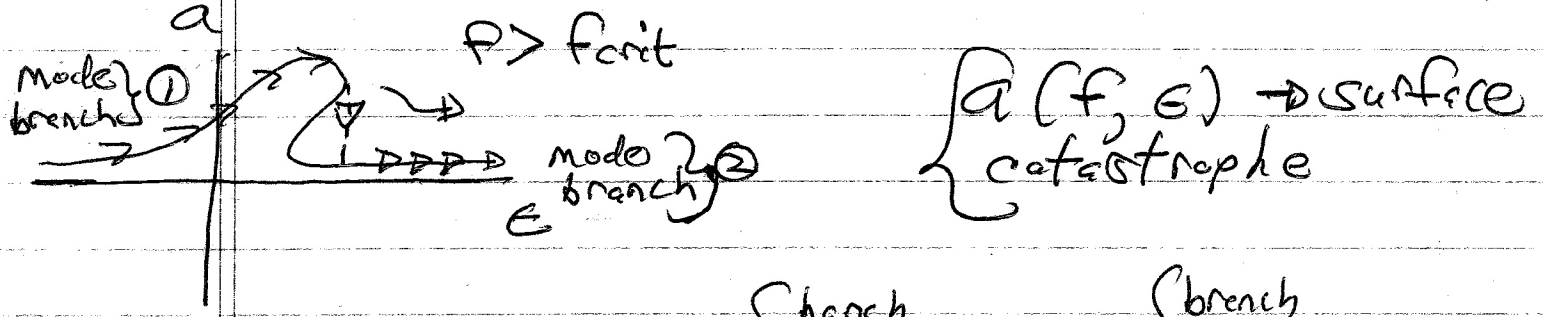
but $n = n_{sol} + \delta n$

$\frac{\partial n}{\partial t} = dn f'(n_{sol})$

$f' > 0 \rightarrow$ instability
 $f' < 0 \rightarrow$ stability

- i.e. $f' > 0 \rightarrow$ unstable root (1)
- $f' < 0 \rightarrow$ stable roots (2)

- 'bifurcation' occurs \rightarrow jump between 2 stable branches (bifurcation \leftrightarrow inflection criterion)

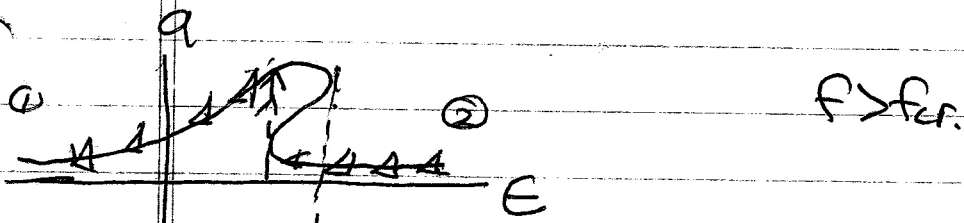


$E > E_{crit}$
 $f > f_{crit}$ \Rightarrow jump from {branch mode 1} to {branch mode 2}

- system exhibits "hysteresis"

i.e.

consider reversal of evolution from (2) \rightarrow (1), i.e.

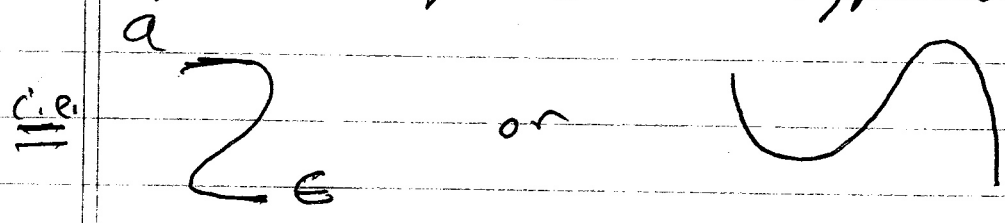


E_{crit} for forward transition
 E_{crit} for back transition

c.e.

$E_{fwd} > E_{back}$
 system tends to 'hang' in mode 2

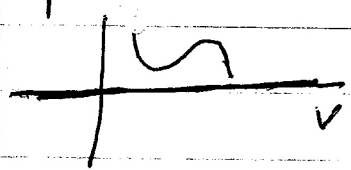
→ driven Duffing oscillator is classic example of "S-curve" type bifurcation



S-curve ⇒ bi-stable system with unstable root in between, yielding transitions or mode-jumping

⇒ akin phase transition
 ①, ② ↔ 2 phases

S curve ↔ p-v curve



mode jumping ↔ first order transition

Now,

- have examined impact of NL on resonance phenomena, at primary/linear resonance

- but, is this the whole story?

⇒ NL-induced resonance phenomena? → { new resonance physics }

Now, re-insert αx^2 term!

$$\ddot{x} + 2\lambda \dot{x} + \omega_0^2 x + \alpha x^2 + \beta x^3 = \frac{f_{ext}}{m}$$

↑
quadratic
NL

hence

$$\omega_{res} = \omega_0 + K a^2$$

$$K = \left(\frac{3\beta}{8\omega_0} - \frac{5\alpha^2}{12\omega_0^3} \right)$$

↓
shift contribution, due quadratic → derive

why $O(\alpha^2)$?
→ $x^{(2)} x^{(1)}$ beat

Why α^2 ?

- observe $-\alpha (x^{(1)})^2 \rightarrow \frac{1}{2} \cos(2\omega_0 t)$ > non-resonant drives to $O(a^2)$

⇒ lowest secular contribution from αx^2 nonlinearity is in $x^{(1)}$ equation
here power a , not a^2