

will
wilson

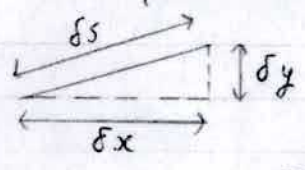
Pset 2 Solutions

Physics 200A Homework 2



The string is continuous, so the Lagrangian density, \mathcal{L} , is of interest

The potential energy is the work done to stretch the string (from equilibrium)



$$\int dV = \int T ds - \int T dx$$

$$\Rightarrow V = \int T \left[1 + \left(\frac{dy}{dx} \right)^2 \right]^{\frac{1}{2}} dx - \int T dx$$

$$\Rightarrow V = \int T \left\{ \left[1 + (d_x y)^2 \right]^{\frac{1}{2}} - 1 \right\} dx$$

So, $\mathcal{L} = \frac{1}{2} \mu (d_t y)^2 - T \left\{ \left[1 + (d_x y)^2 \right]^{\frac{1}{2}} - 1 \right\}$ / $L = \int_0^L dx \mathcal{L}$
 $S = S dt L$

mass per unit length

The E-L equations (for $\mathcal{L} = \mathcal{L}(d_t y, d_x y)$) are given by: $\frac{\delta S}{\delta t} + \frac{\delta S}{\delta x} \left(\frac{\delta S}{\delta y} + \frac{\delta S}{\delta y'} \right) = 0 \dots$

$$\frac{\partial}{\partial t} \frac{\partial \mathcal{L}}{\partial (d_t y)} + \frac{\partial}{\partial x} \frac{\partial \mathcal{L}}{\partial (d_x y)} = 0 \quad \left[\frac{d\mathcal{L}}{dy} = 0 \right]$$

$$\frac{\partial \mathcal{L}}{\partial (d_t y)} = \mu d_t y \quad \frac{\partial \mathcal{L}}{\partial (d_x y)} = -T d_x y \left[1 + (d_x y)^2 \right]^{-\frac{1}{2}}$$

$$\Rightarrow d_t \mu d_t y = d_x \left\{ T d_x y \left[1 + (d_x y)^2 \right]^{-\frac{1}{2}} \right\}$$

If μ, T constant:

$$\mu \frac{d^2 y}{dt^2} = T \frac{d}{dx} \left\{ \frac{d_x y}{[1 + (d_x y)^2]} \right\}$$

This reduces to a wave equation if $|d_x y| \ll 1$ i.e. for a string whose displacement from equilibrium does not change rapidly along the length of the string

$$\mu \frac{d^2 y}{dt^2} = T \frac{d^2 y}{dx^2}$$

b) Taking $\mathcal{L} = \frac{1}{2} \mu (d_t y)^2 - T [1 + (d_x y)^2]^{\frac{1}{2}}$ [The '-T' term has no effect on the e.o.m. \mathcal{L}_0 has been disregarded]

The momentum density conjugate to y :

$$\pi_y = \pi = \frac{d\mathcal{L}}{d(d_t y)} = \mu d_t y$$

\Rightarrow the Hamiltonian density, $\mathcal{H} = \pi \left(\frac{\pi}{\mu} \right) - \frac{1}{2} \mu \left(\frac{\pi}{\mu} \right)^2 + T [1 + (d_x y)^2]^{\frac{1}{2}}$

$$\mathcal{H} = \frac{1}{2} \frac{\pi^2}{\mu} + T [1 + (d_x y)^2]^{\frac{1}{2}}$$

$$\dot{y} = \frac{\partial \mathcal{H}}{\partial \pi} = \frac{\pi}{\mu}$$

$$\dot{\pi} = -\frac{\partial \mathcal{H}}{\partial y} = 0$$

$$c) H = \frac{1}{2} \frac{\pi^2}{\mu} + T \left[1 + \left(\frac{dy}{dx} \right)^2 \right]^{\frac{1}{2}}$$

$$= T \left[1 + \frac{1}{2} \left(\frac{dy}{dx} \right)^2 + \dots \right]$$

$$H = \frac{1}{2} \frac{\pi^2}{\mu} + \frac{1}{2} T \left(\frac{dy}{dx} \right)^2$$

$$\& \text{ also } L = \frac{1}{2} \mu (d_t y)^2 - \frac{1}{2} T (d_x y)^2$$

More generally, taking $H = \pi d_t y - L$

$$\Rightarrow \frac{\delta H}{\delta t} = \frac{\delta \pi}{\delta t} \frac{dy}{dt} + \pi \frac{d^2 y}{dt^2} - \frac{\delta L}{\delta y} \frac{dy}{dt} - \frac{\delta L}{\delta (d_t y)} \frac{d(d_t y)}{dt} - \frac{\delta L}{\delta (d_x y)} \frac{d(d_x y)}{dt}$$

$$= \frac{\delta \pi}{\delta t} \frac{\delta L}{\delta (d_t y)}$$

$$= \frac{\delta L}{\delta y} - \frac{\delta}{\delta x} \frac{\delta L}{\delta (d_x y)} \quad \text{E-L eqns}$$

$$\Rightarrow \frac{\delta H}{\delta t} = \frac{\delta L}{\delta y} \frac{dy}{dt} - \frac{\delta}{\delta x} \frac{\delta L}{\delta (d_x y)} \frac{dy}{dt} - \frac{\delta L}{\delta y} \frac{dy}{dt} - \frac{\delta L}{\delta (d_x y)} \frac{d(d_x y)}{dt}$$

$$= - \frac{\delta}{\delta x} \frac{\delta L}{\delta (d_x y)} \frac{dy}{dx} - \frac{\delta L}{\delta (d_x y)} \frac{d^2 y}{dx dt}$$

$$\Rightarrow \frac{\delta H}{\delta t} = - \frac{\delta}{\delta x} \left\{ \frac{\delta L}{\delta (d_x y)} \frac{dy}{dt} \right\} \quad \text{in 1-D. For 3D, } d_x \rightarrow \nabla$$

C.f. the Poynting theorem:

$$-\frac{du}{dt} = \vec{\nabla} \cdot \vec{S} + \vec{J}_p \cdot \vec{E}$$

\vec{S} ← Poynting vector
 \vec{J}_p ← free current density
 \vec{E} ← electric field

u : energy density

$H \leftrightarrow u$ energy density

$$\frac{\partial \mathcal{L}}{\partial(d_x y)} \frac{dy}{dt} \leftrightarrow S_x \text{ energy flux density}$$

$$\frac{dH}{dt} = -\int \vec{\nabla} \cdot \vec{S} d^3x = -\int \vec{S} \cdot d\vec{S} \quad \text{in general}$$

In this case: $\frac{\partial \mathcal{L}}{\partial(d_x y)} = -T d_x y$

$$\text{so } \vec{S} = -T d_x y d_t y \hat{e}_1$$

← unit vector in the +x direction

PSet 2: Problem 2

Tom Zdyrski

Problem 2: Consider a particle with velocity \vec{v}_1 moving in a region with constant potential U_1 into a region with constant potential U_2 (c.f. fig. 1).

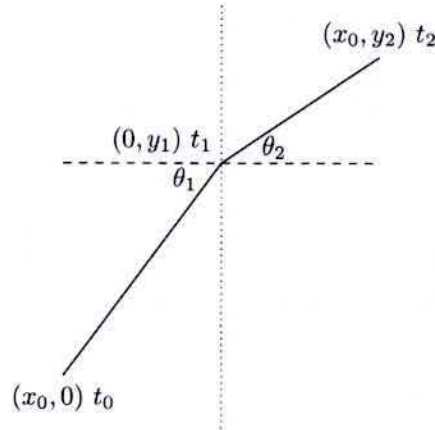


Figure 1: Setup of problem

Part a): v_1 is the initial velocity. How does the direction change? What is v_2 ? For this problem, we fix the end points $(x_0, 0)$ and (x_2, y_2) as well as the end times t_0 and t_2 . We have shown in class that free particles move along geodesics (in flat space, straight lines at constant velocity). Since the potential U is constant on either side of the boundary, the force $\vec{F} = -\vec{\nabla}U$ is also zero on either side of the boundary. Thus, we immediately know it follows a straight path in either region. Labeling the crossing point as $(0, y_1)$ at time t_1 , we get:

$$\begin{aligned} S &= \int_{t_0}^{t_2} L dt = \int_{t_0}^{t_1} \left(\frac{1}{2} m v_1^2 - U_1 \right) dt + \int_{t_1}^{t_2} \left(\frac{1}{2} m v_2^2 - U_2 \right) dt \\ &= \left(\frac{1}{2} m v_1^2 - U_1 \right) \int_{t_0}^{t_1} dt + \left(\frac{1}{2} m v_2^2 - U_2 \right) \int_{t_1}^{t_2} dt \\ &= \left(\frac{1}{2} m v_1^2 - U_1 \right) (t_1 - t_0) + \left(\frac{1}{2} m v_2^2 - U_2 \right) (t_2 - t_1) \\ &= \frac{m(x_0^2 + y_1^2)}{2t_1} - U_1 t_1 + \frac{m(x_0^2 + (y_2 - y_1)^2)}{2(t_2 - t_1)} - U_2 t_2 \end{aligned}$$

Since the endpoints are fixed, $S = S(y_1, t_1)$. Then, the principle of least action gives:

$$dS = 0 \frac{\partial S}{\partial y_1} dy_1 + \frac{\partial S}{\partial t_1} dt_1$$

Since dy_1 and dt_1 are arbitrary, the partials must each be identically zero. Then,

$$\frac{\partial S}{\partial y_1} = m \frac{y_1}{t_1} - m \frac{y_2 - y_1}{t_2 - t_1} \implies m v_1 \sin \theta_1 = m v_2 \sin \theta_2 \implies v_1 \sin \theta_1 = v_2 \sin \theta_2$$

Doing the same for the time-derivative yields,

$$\frac{\partial S}{\partial t_1} = 0 = -\frac{1}{2} m \frac{x_0^2 + y_1^2}{t_1^2} - U_1 + \frac{1}{2} m \frac{x_0^2 + (y_2 - y_1)^2}{(t_2 - t_1)^2} + U_2 \implies \frac{1}{2} m v_1^2 + U_1 = \frac{1}{2} m v_2^2 + U_2 \implies E_1 = E_2$$

Thus, we have

$$\frac{1}{2}mv_2^2 = \frac{1}{2}mv_1^2 + U_1 - U_2 \implies v_2 = v_1 \sqrt{1 + \frac{U_1 - U_2}{\frac{1}{2}mv_1^2}}$$

and

$$\sin \theta_2 = \frac{v_1}{v_2} \sin \theta_1 = \frac{\sin \theta_1}{\sqrt{1 + \frac{U_1 - U_2}{\frac{1}{2}mv_1^2}}}$$

- (a) Part b): Find the ratio of times in the same path for particles with different masses but the same U . Multiplying the Lagrangian by a constant doesn't change the trajectory of the system. Thus, if $m \rightarrow \alpha m$ and $t \rightarrow \beta t$, we know the constant potentials $U_{1,2} \rightarrow U_{1,2}$ is unchanged. Therefore, we require the kinetic term is also unchanged: $1/2mv^2 \rightarrow 1/2m\alpha v^2/\beta^2$. That is, $\beta = \sqrt{\alpha} \implies t \propto \sqrt{m}$ which tells us

$$\frac{t_2}{t_1} = \sqrt{\frac{m_2}{m_1}}$$

- (b) Part c): Find the ratio of times in the same path for particles with the same mass but moving [in] different potentials U_1, U_2 , where $U_2/U_1 = c$, a constant. Likewise, if the two setups differ by changing $U \rightarrow cU$ (i.e. $U_2 = cU_1$), then we must have $t \rightarrow \beta t$ s.t. $1/2mv^2 \rightarrow 1/2mv^2/\beta^2 = 1/2mv^2c$. This shows that $\beta = 1/\sqrt{c} \implies t \propto 1/\sqrt{U}$ showing that

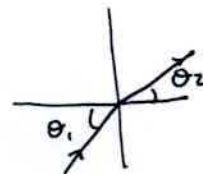
$$\frac{t_2}{t_1} = \sqrt{\frac{U_1}{U_2}}$$

- (c) Part d): What problem in optics does this problem resemble? This problem resembles Snell's Law of refraction. Like Snell's law, we see that the particle paths bend when they cross the boundary. However, a key difference exists: In Snell's law, $\sin \theta \propto v$ whereas here $\sin \theta \propto 1/v$. This represents the fact that light waves refract towards areas with slower wave speed whereas particles refract away from areas with slower particle speed.

* Alternative solution to (a):

the potential energy is independent of the coordinates whose axes are parallel to the plane separating the half-spaces

$$\begin{aligned} \Rightarrow & \bullet mv_1 \sin \theta_1 = mv_2 \sin \theta_2 \\ & \bullet \frac{1}{2}mv_1^2 + U_1 = \frac{1}{2}mv_2^2 + U_2 \end{aligned}$$



\Rightarrow solution.

$$3. a \quad L = -mc^2 \left(1 - \frac{v^2}{c^2}\right)^{1/2} - V(\vec{r})$$

$$\vec{p} = \frac{\partial L}{\partial \dot{\vec{r}}} = \frac{m\dot{\vec{r}}}{\left(1 - \frac{v^2}{c^2}\right)^{1/2}}$$

$$\frac{\partial L}{\partial \vec{r}} = -\frac{\partial V(\vec{r})}{\partial \vec{r}} = -\nabla V(\vec{r})$$

The eq. of motion is then

$$\left[\frac{d\vec{p}}{dt} = -\nabla V(\vec{r}) \right]$$

Since $\left[\vec{p} = \gamma m \vec{v} \right]$ is the correct relativistic momentum, this Lagrangian reproduces relativistic mechanics. Explicitly,

$$\frac{d\vec{p}}{dt} = \frac{m\ddot{\vec{r}}}{\left(1 - \frac{v^2}{c^2}\right)^{1/2}} + \frac{m\dot{\vec{r}}}{\left(1 - \frac{v^2}{c^2}\right)^{3/2}} \frac{\dot{\vec{r}} \cdot \dot{\vec{r}}}{c^2}$$

$$= \gamma m \ddot{\vec{r}} + \gamma^3 m \dot{\vec{r}} \left(\frac{\dot{\vec{r}} \cdot \dot{\vec{r}}}{c^2} \right) = -\nabla V(\vec{r})$$

b. As found above,

$$\vec{p} = \gamma m \dot{\vec{r}}$$

$$\text{so } H = \vec{p} \cdot \dot{\vec{r}} - L = \gamma m v^2 + \frac{mc^2}{\gamma} + V(\vec{r})$$

$$= \frac{m(v^2 + c^2(1 - \frac{v^2}{c^2}))}{\left(1 - \frac{v^2}{c^2}\right)^{1/2}} + V(\vec{r}) = \underbrace{\gamma mc^2 + V(\vec{r})}$$

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due to printing error

$$\begin{aligned} \text{Now, } p^2 c^2 + m^2 c^4 &= \frac{m^2 c^2 v^2}{1 - v^2/c^2} + \frac{m^2 c^4 (1 - v^2/c^2)}{(1 - v^2/c^2)} \\ &= \frac{m^2 c^4}{(1 - v^2/c^2)} \quad \text{so indeed} \end{aligned}$$

$$H = \gamma m c^2 + V(\vec{r}) = \left(p^2 c^2 + m^2 c^4 \right)^{1/2} + V(\vec{r})$$

This is a const. of the motion since $\frac{\partial H}{\partial t} = 0$

c. If $V = V(r)$, then

$$\dot{\vec{p}} = -\nabla V = -\frac{\partial}{\partial r} V(r) \hat{r}$$

$$\begin{aligned} \text{and } \frac{d}{dt} (\vec{r} \times \vec{p}) &= \dot{\vec{r}} \times \vec{p} + \vec{r} \times \dot{\vec{p}} \\ &= \gamma m (\dot{\vec{r}} \times \dot{\vec{r}}) + \vec{r} \times \left(-V'(r) \hat{r} \right) \\ &= 0 \end{aligned}$$

Hence $\vec{r} \times \vec{p}$ is a const of the motion.

Since angular momentum is conserved,
the particle moves in a plane.

Orient \hat{z} normal to this plane, so that $p_\theta = 0$

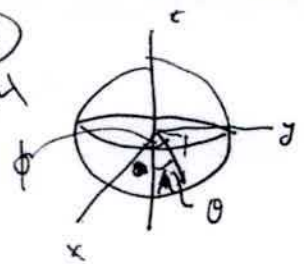
and $\theta = \frac{\pi}{2}$. Then $p^2 = p_r^2 + \frac{p_\phi^2}{r^2 \sin^2 \theta} + \frac{p_\theta^2}{r^2}$

$$= p_r^2 + \frac{p_\phi^2}{r^2}$$

$$\Rightarrow H = c \left(p_r^2 + \frac{p_\phi^2}{r^2} + m^2 c^2 \right)^{1/2} + V(r)$$

6.2

#4



$$L = \frac{1}{2} m (\dot{r}^2 + r^2 \dot{\theta}^2 + r^2 \sin^2 \theta \cdot \dot{\phi}^2) + mgl \cos \theta$$

$$= \frac{1}{2} m (l^2 \dot{\theta}^2 + l^2 \sin^2 \theta \cdot \dot{\phi}^2) + mgl \cos \theta$$

$$p_{\theta} = \frac{\partial L}{\partial \dot{\theta}} = ml^2 \dot{\theta} \rightarrow \dot{\theta} = p_{\theta} / ml^2$$

$$p_{\phi} = \frac{\partial L}{\partial \dot{\phi}} = ml^2 \sin^2 \theta \cdot \dot{\phi} \rightarrow \dot{\phi} = p_{\phi} / ml^2 \sin^2 \theta$$

$$H = \sum_{\sigma} p_{\sigma} \dot{q}_{\sigma} - L = \frac{p_{\theta}^2}{ml^2} + \frac{p_{\phi}^2}{ml^2 \sin^2 \theta} - \frac{1}{2} m \left(l^2 \cdot \frac{p_{\theta}^2}{(ml^2)^2} + l^2 \sin^2 \theta \cdot \frac{p_{\phi}^2}{(ml^2 \sin^2 \theta)^2} \right) - mgl \cos \theta$$

$$H = \frac{p_{\theta}^2}{2ml^2} + \frac{p_{\phi}^2}{2ml^2 \sin^2 \theta} - mgl \cos \theta$$

$$\frac{\partial H}{\partial p_{\phi}} = \dot{\phi} \Rightarrow \frac{p_{\phi}}{ml^2 \sin^2 \theta} = \dot{\phi}$$

$$\frac{\partial H}{\partial p_{\theta}} = \dot{\theta} \Rightarrow \frac{p_{\theta}}{ml^2} = \dot{\theta}$$

$$\Rightarrow \dot{p}_{\theta} = \ddot{\theta} ml^2$$

$$\frac{\partial H}{\partial \phi} = -\dot{p}_{\phi} \Rightarrow 0 = -\dot{p}_{\phi} \Rightarrow p_{\phi} = \text{const.}$$

$$\frac{\partial H}{\partial \theta} = -\dot{p}_{\theta} \Rightarrow mgl \sin \theta - \left(\frac{p_{\phi}^2}{ml^2} \right) \cdot \frac{\cos \theta}{\sin^3 \theta} = -\dot{p}_{\theta}$$

$$\Rightarrow mgl \sin \theta - ml^2 \dot{\phi}^2 \sin \theta \cos \theta = -\ddot{\theta} ml^2$$

$$\Rightarrow \ddot{\theta} + \frac{g}{l} \sin \theta - \dot{\phi}^2 \sin \theta \cos \theta = 0$$

(I'm interpreting exact Hamiltonian as exact Hamiltonian equations of motion...)

For uniform circular motion $\ddot{\theta} = \dot{\theta} = 0$.

$$\cos \theta_0 = \frac{g}{l \dot{\phi}^2}$$

Near Univ. Circular Motion: Neglect $O((\theta - \theta_0)^2)$

~~$$\ddot{\theta} + \frac{g}{l} (\sin \theta_0 + \cos \theta_0 (\theta - \theta_0) - \sin \theta_0 (\theta - \theta_0)^2) - \dot{\phi}^2 (\sin \theta_0 + \cos \theta_0 (\theta - \theta_0)) (\cos \theta_0 - \sin \theta_0 (\theta - \theta_0)) = 0$$

$$\ddot{\theta} + \frac{g}{l} (\cos \theta_0 (\theta - \theta_0) - \sin \theta_0 (\theta - \theta_0)^2) - \dot{\phi}^2 (-\sin^2 \theta_0 (\theta - \theta_0) + \cos^2 \theta_0 (\theta - \theta_0) - \sin \theta_0 \cos \theta_0 (\theta - \theta_0)^2) = 0$$

$$\ddot{\theta} + \left[\frac{g}{l} \cos \theta_0 + \dot{\phi}^2 (1 - 2 \cos^2 \theta_0) \right] (\theta - \theta_0) = 0$$~~

Next Page ▽.

(6.2) continued

This works out better if we use

$$\ddot{\theta} + \frac{g}{l} \sin \theta - \left(\frac{P\dot{\phi}}{mR^2}\right)^2 \frac{\cos \theta}{\sin^3 \theta} = 0$$

$$\text{Note: } \frac{\cos \theta}{(\sin \theta)^3} \approx \frac{\cos \theta_0 - \sin \theta_0 \cdot \epsilon}{(\sin \theta_0 + \epsilon \cdot \cos \theta_0)^3} = \frac{\cos \theta_0 - \epsilon \sin \theta_0}{\sin^3 \theta_0 (1 + \epsilon \cdot \frac{\cos \theta_0}{\sin \theta_0})^3} \approx \frac{(\cos \theta_0 - \epsilon \cdot \sin \theta_0) \left(1 - 3 \frac{\cos \theta_0}{\sin \theta_0} \epsilon\right)}{\sin^3 \theta_0}$$

$\theta = \theta_0 + \epsilon$

$$\Rightarrow \ddot{\epsilon} + \frac{g}{l} (\sin \theta_0 + \epsilon \cdot \cos \theta_0) - \left(\frac{P\dot{\phi}}{mR^2}\right)^2 \left(\frac{\cos \theta_0}{\sin^3 \theta_0} - \frac{3 \cos^2 \theta_0}{\sin^4 \theta_0} \epsilon - \epsilon \cdot \frac{1}{\sin^2 \theta_0} \right) = 0$$

$$\text{Recall } \cos \theta_0 = \frac{g}{l \dot{\phi}_0^2}, \quad \dot{\phi}_0^2 = \frac{P\dot{\phi}}{mR^2 \sin^2 \theta_0}, \quad \left(\frac{P\dot{\phi}}{mR^2}\right)^2 = \dot{\phi}_0^4 \sin^4 \theta_0$$

~~$$\ddot{\epsilon} + \frac{g}{l} (\sin \theta_0 + \epsilon \cos \theta_0) - \dot{\phi}_0^4 \sin^4 \theta_0 \left(\frac{\cos \theta_0}{\sin^3 \theta_0} - \frac{3 \cos^2 \theta_0}{\sin^4 \theta_0} \epsilon - \epsilon \frac{1}{\sin^2 \theta_0} \right) = 0$$~~

$$\Rightarrow \ddot{\epsilon} + \left(\frac{g}{l} \sin \theta_0 - \dot{\phi}_0^4 \sin \theta_0 \cos \theta_0 \right) + \epsilon \cdot \left(\frac{g}{l} \cos \theta_0 + 3 \dot{\phi}_0^4 \cos^2 \theta_0 + \dot{\phi}_0^4 \sin^2 \theta_0 \right) = 0$$

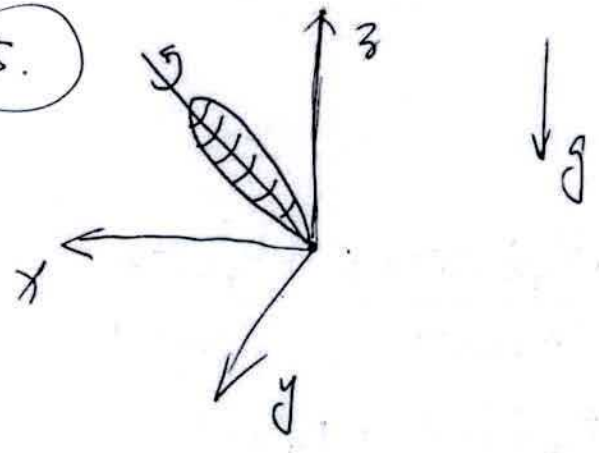
$= 0$, since $\dot{\phi}_0^2 \cos \theta_0 = g/l$ $\epsilon \cdot \dot{\phi}_0^2 (4 \cos^2 \theta_0 + 1 - \cos^2 \theta_0)$

$$\Rightarrow \ddot{\epsilon} = -\epsilon \cdot \dot{\phi}_0^2 (1 + 3 \cos^2 \theta_0)$$

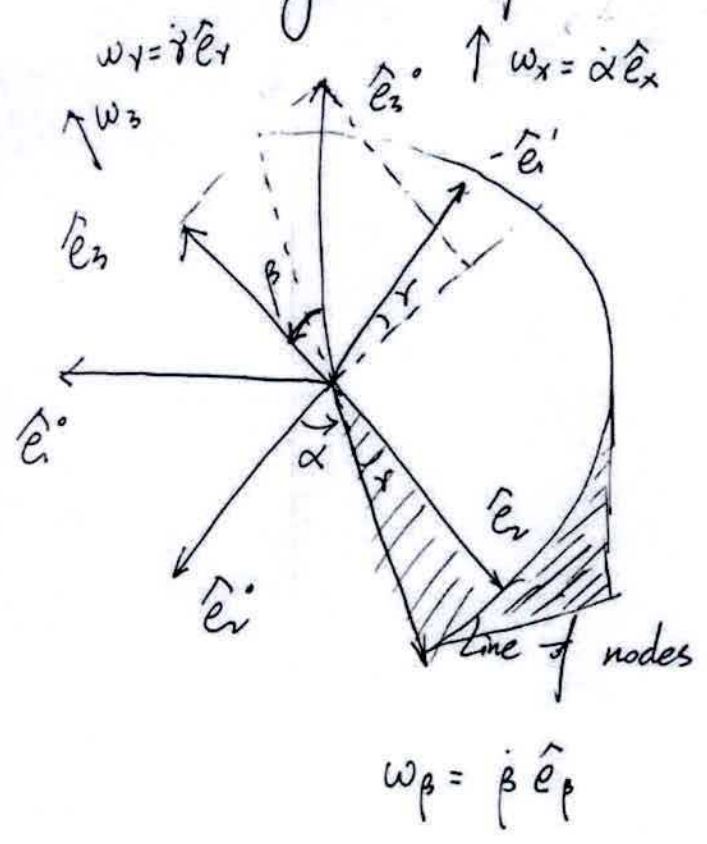
$$\ddot{\epsilon} = -\epsilon \cdot \frac{g}{l \cos \theta_0} \cdot (1 + 3 \cos^2 \theta_0)$$

Problem Set II.

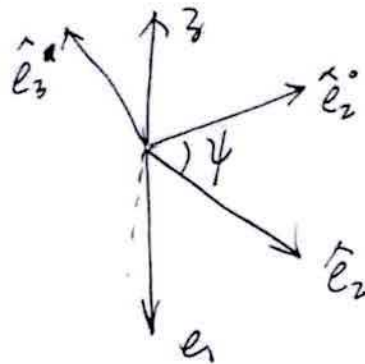
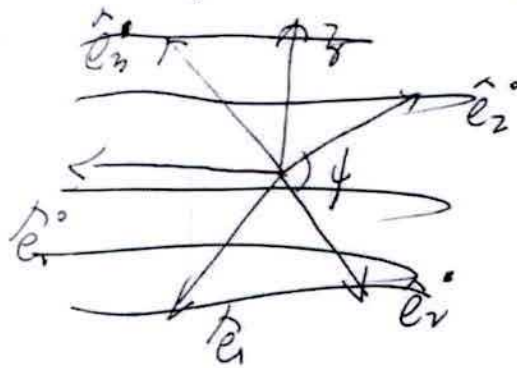
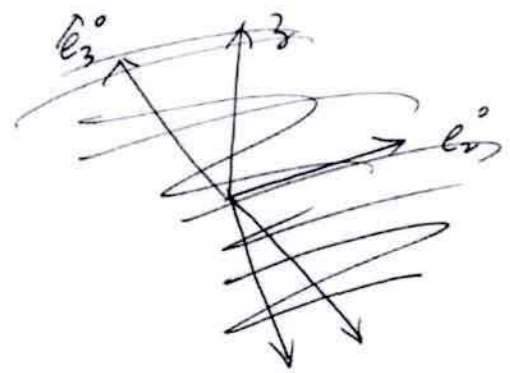
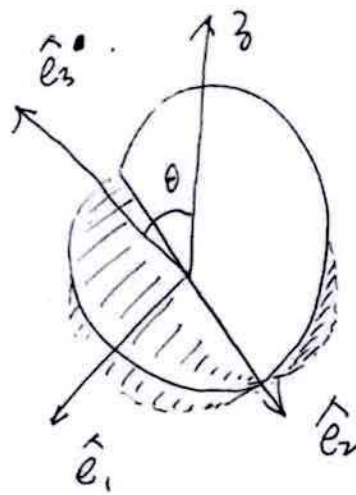
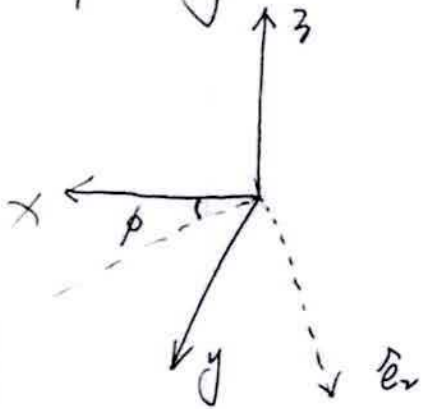
5.



Euler Angles Definition. α . β . γ P155



Separately



spin

$$\vec{\omega} = \dot{\phi} \hat{z} + \dot{\theta} \hat{e}_2 + \dot{\psi} \hat{e}_3$$

$$\hat{z} = \cos\theta \hat{e}_3^0 - \sin\theta \hat{e}_1 \Rightarrow \vec{\omega} = -\dot{\phi} \sin\theta \hat{e}_1 + \dot{\theta} \hat{e}_2 + (\dot{\psi} + \dot{\phi} \cos\theta) \hat{e}_3$$

Angular Momentum \vec{L} ($\lambda_1 \omega_1, \lambda_2 \omega_2, \lambda_3 \omega_3$)

$$\vec{L} = -\lambda_1 \dot{\phi} \sin\theta \hat{e}_1 + \lambda_2 \dot{\theta} \hat{e}_2 + \lambda_3 (\dot{\psi} + \dot{\phi} \cos\theta) \hat{e}_3$$

As a symmetric top, $\lambda_1 = \lambda_2$: $T = \frac{1}{2} I_{\perp} \omega_{\perp}^2 + \frac{1}{2} I_3 \omega_3^2$

$$T = \frac{1}{2} \lambda_1 (\dot{\phi}^2 \sin^2\theta + \dot{\theta}^2) + \frac{1}{2} \lambda_3 (\dot{\psi} + \dot{\phi} \cos\theta)^2 \left\{ \begin{array}{l} \mathcal{L} = \frac{1}{2} \lambda_1 (\dot{\phi}^2 \sin^2\theta + \dot{\theta}^2) + \frac{1}{2} \lambda_3 (\dot{\psi} + \dot{\phi} \cos\theta)^2 \\ - M g R \cos\theta \end{array} \right.$$

$$U = M g R \cos\theta$$

$$P_{\theta} = \frac{\partial \mathcal{L}}{\partial \dot{\theta}} = \lambda_1 \dot{\theta}$$

$$P_{\psi} = \frac{\partial \mathcal{L}}{\partial \dot{\psi}} = \lambda_3 (\dot{\psi} + \dot{\phi} \cos\theta)$$

$$P_{\phi} = \frac{\partial \mathcal{L}}{\partial \dot{\phi}} = \lambda_1 \dot{\phi} \sin^2\theta + P_{\psi} \cos\theta$$

$$\Rightarrow T = \frac{(\lambda_1 \dot{\phi} \sin^2\theta)^2}{2\lambda_1 \sin^2\theta} + \frac{\lambda_1}{2} \left(\frac{P_{\theta}}{\lambda_1}\right)^2 + \frac{\lambda_3}{2} \left(\frac{P_{\psi}}{\lambda_3}\right)^2 = \frac{(P_{\phi} - P_{\psi} \cos\theta)^2}{2\lambda_1 \sin^2\theta} + \frac{P_{\theta}^2}{2\lambda_1} + \frac{P_{\psi}^2}{2\lambda_3}$$

$$\therefore H = T + M g R \cos\theta = \sum p \dot{q} - \mathcal{L}$$

Show explicitly:

$$H = p_\theta \dot{\theta} + p_\psi \dot{\psi} + p_\phi \dot{\phi} - \mathcal{L}$$

$$\dot{\theta} = \frac{p_\theta}{\lambda_1} \quad ; \quad \dot{\psi} = \frac{p_\psi}{\lambda_3} - \dot{\phi} \cos \theta \quad ; \quad \dot{\phi} = \frac{p_\phi - p_\psi \cos \theta}{\lambda_1 \sin^2 \theta}$$

$$\Rightarrow p_\theta \dot{\theta} = \frac{p_\theta^2}{\lambda_1} \quad ; \quad p_\psi \dot{\psi} + p_\phi \dot{\phi} = \frac{p_\psi^2}{\lambda_3} - p_\psi \dot{\phi} \cos \theta + \frac{p_\phi^2}{\lambda_1 \sin^2 \theta} - \frac{p_\phi p_\psi \cos \theta}{\lambda_1 \sin^2 \theta}$$

$$p_\psi = \lambda_3 \dot{\psi} + \lambda_3 \dot{\phi} \cos \theta$$

$$p_\phi = \lambda_1 \dot{\phi} \sin^2 \theta + p_\psi \cos \theta$$

$$p_\phi \dot{\phi} = \frac{p_\phi}{\lambda_1} (p_\phi - p_\psi \cos \theta) \operatorname{cosec}^2 \theta$$

$$p_\psi \dot{\psi} = \frac{p_\psi^2}{\lambda_3} + \frac{p_\psi}{\lambda_1} (p_\psi \cos \theta - p_\phi) \cotan(\theta) \operatorname{cosec}(\theta)$$

$$\Rightarrow H = \frac{p_\theta^2}{\lambda_1} + \frac{p_\psi^2}{\lambda_3} + \frac{p_\phi^2}{\lambda_1 \sin^2 \theta} - \frac{p_\phi p_\psi \cos \theta}{\lambda_1 \sin^2 \theta} + \frac{p_\psi^2 \cos^2 \theta}{\lambda_1 \sin^2 \theta} - \frac{p_\psi p_\phi \cos \theta}{\lambda_1 \sin^2 \theta} - \mathcal{L}$$

$$\left[H = \frac{p_\theta^2}{\lambda_1} + \frac{(p_\phi - p_\psi \cos \theta)^2}{\lambda_1 \sin^2 \theta} + \frac{p_\psi^2}{\lambda_3} - \mathcal{L} \right]$$

$$\mathcal{L} = \frac{1}{2} \frac{(p_\phi - p_\psi \cos \theta)^2}{\lambda_1 \sin^2 \theta} + \frac{1}{2} \frac{p_\theta^2}{\lambda_1} + \frac{1}{2} \frac{p_\psi^2}{\lambda_3} - M g R \cos \theta$$

$$\Rightarrow \boxed{H = \frac{1}{2} \frac{p_\theta^2}{\lambda_1} + \frac{1}{2} \frac{p_\psi^2}{\lambda_3} + \frac{1}{2} \frac{(p_\phi - p_\psi \cos \theta)^2}{\lambda_1 \sin^2 \theta} + M g R \cos \theta}$$

Hamilton Equations:

$$\dot{\phi} = \frac{\partial H}{\partial P_{\phi}} = \frac{P_{\phi} - P_{\phi} \cos \theta}{\lambda_1 \sin^2 \theta} = 31.3a \text{ in FW}$$

$$\dot{\theta} = \frac{\partial H}{\partial P_{\theta}} = \frac{P_{\theta}}{\lambda_1}$$

$$\dot{\psi} = \frac{\partial H}{\partial P_{\psi}} = - \frac{P_{\psi} - P_{\psi} \cos \theta}{\lambda_1 \sin^2 \theta} \cos \theta + \frac{P_{\psi}}{\lambda_3} = 31.3b \text{ in FW}$$

$$\dot{P}_{\phi} = \dot{P}_{\psi} = 0 \quad \} = 31.2ab \text{ in FW}$$

$$\begin{aligned} \dot{P}_{\theta} &= - \frac{\partial H}{\partial \theta} = MgR \sin \theta + \frac{-2(P_{\phi} - P_{\psi} \cos \theta) \sin \theta \overset{P_{\psi}}{\cancel{\sin \theta}}}{2\lambda_1 \sin^2 \theta} + \frac{(P_{\phi} - P_{\psi} \cos \theta)^2 \sin \theta \cos \theta}{(\cancel{\lambda_1 \sin^2 \theta})^2 2\lambda_1} \\ &= \frac{(P_{\psi} \cos \theta - P_{\phi}) [(P_{\phi} \cos \theta - P_{\psi}) \cos \theta + \sin^2 \theta]}{\lambda_1 \sin^3 \theta} + MgR \sin \theta \end{aligned}$$

$$P_{\theta} \dot{\theta} = - \frac{1}{\lambda_1 \sin^3 \theta} \left(-(P_{\phi}^2 + P_{\psi}^2) \cos \theta + P_{\phi} P_{\psi} (1 + \cos^2 \theta) \right) + MgR \sin \theta$$

where P_{ϕ}, P_{ψ} are constants

so $\lambda_1 \ddot{\theta} = \dot{P}_{\theta}$ reduces to 31.5b in FW

Now: $\dot{\theta} d\theta = \frac{d\dot{\theta}}{dt} d\theta = \dot{\theta} d\dot{\theta}$

$$\Rightarrow \lambda_1 \int \dot{\theta} d\dot{\theta} = \int \underset{\substack{\downarrow \\ \text{function of } \theta}}{\dot{P}_{\theta}} d\theta \quad \} \text{ can integrate for solution } \checkmark$$

#6

a) Lagrangian density

$$\mathcal{L} = \frac{\hbar^2}{2m} \left(\frac{d\psi^*}{dx} \right) \left(\frac{d\psi}{dx} \right) - \psi^* (U - E) \psi$$

so equation for ψ^* is given by

$$\frac{d}{dx} \frac{\partial \mathcal{L}}{\partial \frac{d\psi^*}{dx}} - \frac{\partial \mathcal{L}}{\partial \psi^*} = 0$$

$$\Rightarrow \left[-\frac{\hbar^2}{2m} \frac{d^2 \psi}{dx^2} + U\psi = E\psi \right] \text{ which is the time-independent Schrödinger equation}$$

(b) the Lagrangian is invariant under gauge symmetry

$$\begin{aligned} \psi &\rightarrow e^{i\theta} \psi \\ \psi^* &\rightarrow e^{-i\theta} \psi^* \end{aligned}$$

phase rotation \rightarrow not gauge since $\theta \neq \theta(x)$

$$\theta \rightarrow \theta + \delta\theta \Rightarrow \psi \rightarrow \psi + \frac{i\delta\theta}{\hbar} \psi \sim \psi + (\delta\theta) Q\psi$$

$$\psi^* \rightarrow \psi^* - \frac{i\delta\theta}{\hbar} \psi^* \sim \psi^* - \frac{i}{\hbar} \psi^* (\delta\theta)$$

from Noether's theorem

$$\delta \int j_{\mu} dx = 0$$

$$\Rightarrow \cancel{j} \quad j_{\mu} = \frac{\partial \mathcal{L}}{\partial (\partial_{\mu} \phi_a)} \delta \phi_a$$

$$\text{where } \begin{cases} Q\psi = \frac{i}{\hbar} \psi \\ Q\psi^* = -\frac{i}{\hbar} \psi^* \end{cases}$$

$$\begin{aligned} \text{here } \Rightarrow j &= - \left(\frac{\partial \mathcal{L}}{\partial \frac{d\psi}{dx}} \right) Q\psi - \left(\frac{\partial \mathcal{L}}{\partial \frac{d\psi^*}{dx}} \right) Q\psi^* \\ &= \frac{\hbar}{2m} \left[\psi^* \frac{d\psi}{dx} - \frac{d\psi^*}{dx} \psi \right] \end{aligned}$$

which is the current we want \Rightarrow for time-independent \mathcal{L}

$$\frac{\partial j}{\partial x} = 0$$

for probability density $\rho = \psi^* \psi$

$$\frac{\partial \rho}{\partial t} = \frac{\partial (\psi^* \psi)}{\partial t} = \left(\psi^* \frac{\partial \psi}{\partial t} + \psi \frac{\partial \psi^*}{\partial t} \right)$$

$$\text{for } \hat{E} = i\hbar \frac{\partial}{\partial t}$$

$$\Rightarrow \frac{d\psi}{dt} = \frac{i}{\hbar} \left[-\frac{\hbar^2}{2m} \frac{d^2 \psi}{dx^2} + U\psi \right]$$

$$\Rightarrow \frac{d\psi^*}{dt} = \frac{1}{i\hbar} \left[-\frac{\hbar^2}{2m} \frac{d^2\psi^*}{dx^2} + U\psi^* \right]$$

$$\Rightarrow \frac{\partial P}{\partial t} = \frac{-i}{\hbar} \left[\frac{\hbar^2}{2m} \frac{d^2\psi^*}{dx^2} \psi - U\psi^*\psi - \frac{\hbar^2}{2m} \frac{d^2\psi}{dx^2} \psi^* + U\psi^*\psi \right]$$

$$= -\frac{dJ}{dx} = 0$$

~~or if we know the time dependence Schrödinger equation.~~
 \Rightarrow probability is conserved

And if we can get time-dependent Schrödinger equation we can use Noether's theorem in 4-vector form

⑦ (a) $L = L(q, \dot{q}, \ddot{q}, t)$

$$S = \int_{t_1}^{t_2} L(q, \dot{q}, \ddot{q}, t) dt, \quad \text{consider } q = q + \delta q$$

$$\begin{aligned} \text{Then } \delta S &= \int_{t_1}^{t_2} L(q, \dot{q}, \ddot{q}, t) dt - \int_{t_1}^{t_2} L(q + \delta q, \dot{q} + \delta \dot{q}, \ddot{q} + \delta \ddot{q}, t) dt \\ &= \int_{t_1}^{t_2} (L(q, \dot{q}, \ddot{q}, t) - L(q + \delta q, \dot{q} + \delta \dot{q}, \ddot{q} + \delta \ddot{q}, t)) dt \end{aligned}$$

First order Approx in δq

$$\delta S \approx \int_{t_1}^{t_2} \left(\frac{\partial L}{\partial q} \delta q + \frac{\partial L}{\partial \dot{q}} \delta \dot{q} + \frac{\partial L}{\partial \ddot{q}} \delta \ddot{q} \right) dt$$

We will use $\delta \dot{q} = \frac{d}{dt} \delta q$, $\delta \ddot{q} = \frac{d^2}{dt^2} \delta q$

$$\delta S = \int_{t_1}^{t_2} \left(\frac{\partial L}{\partial q} \delta q + \frac{\partial L}{\partial \dot{q}} \frac{d \delta q}{dt} + \frac{\partial L}{\partial \ddot{q}} \frac{d^2 \delta q}{dt^2} \right) dt$$

① ②

①: $\int_{t_1}^{t_2} \left(\frac{\partial L}{\partial \dot{q}} \frac{d \delta q}{dt} \right) dt$: let $u = \frac{\partial L}{\partial \dot{q}}$, $dv = \frac{d \delta q}{dt}$

$$= \underbrace{\frac{\partial L}{\partial \dot{q}} \delta q}_{\text{b.c. fixed endpoints}} \Big|_{t_1}^{t_2} - \int_{t_1}^{t_2} \left(\frac{d}{dt} \frac{\partial L}{\partial \dot{q}} \right) \delta q dt$$

o b.c. fixed endpoints

7) continued #1

$$\textcircled{2} \int_{t_1}^{t_2} \left(\frac{\partial L}{\partial \dot{q}} \frac{d^2}{dt^2} \delta q \right) dt = \left(\frac{\partial L}{\partial \dot{q}} \frac{d}{dt} \delta q \right) \Big|_{t_1}^{t_2} - \int_{t_1}^{t_2} \left[\frac{d}{dt} \delta q \left(\frac{d}{dt} \frac{\partial L}{\partial \dot{q}} \right) \right] dt$$

$$= \underbrace{\left(\frac{\partial L}{\partial \dot{q}} \frac{d}{dt} \delta q \right) \Big|_{t_1}^{t_2}}_{\text{fixed endpoints}} - \underbrace{\left(\frac{d}{dt} \frac{\partial L}{\partial \dot{q}} \cdot \delta q \right) \Big|_{t_1}^{t_2}}_{\text{fixed endpoints}} + \int_{t_1}^{t_2} \left(\frac{d^2}{dt^2} \frac{\partial L}{\partial \dot{q}} \delta q \right) dt$$

0 if we pick $\frac{d\delta q}{dt}(t_1, t_2) = 0$

which we will do.

Recap:
$$\delta S = \underbrace{\left(\frac{\partial L}{\partial \dot{q}} \delta q + \frac{\partial L}{\partial \ddot{q}} \frac{d}{dt} \delta q - \frac{d}{dt} \frac{\partial L}{\partial \ddot{q}} \delta q \right) \Big|_{t_1}^{t_2}}_{\text{All vanish}} + \int_{t_1}^{t_2} \left(\frac{\partial L}{\partial q} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}} + \frac{d^2}{dt^2} \frac{\partial L}{\partial \ddot{q}} \right) \delta q dt = 0$$

$$\Rightarrow \boxed{\frac{\partial L}{\partial q} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}} + \frac{d^2}{dt^2} \frac{\partial L}{\partial \ddot{q}} = 0}$$

⑦ continued #2

$$0 = \left(\frac{\partial L}{\partial q} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}} \right) + \frac{d^2}{dt^2} \frac{\partial L}{\partial \ddot{q}} \rightarrow \frac{\partial L}{\partial q} = \frac{d}{dt} \frac{\partial L}{\partial \dot{q}} - \frac{d^2}{dt^2} \frac{\partial L}{\partial \ddot{q}}$$

$$\frac{dL}{dt} = \frac{\partial L}{\partial q} \dot{q} + \frac{\partial L}{\partial \dot{q}} \ddot{q} + \frac{\partial L}{\partial \ddot{q}} \dddot{q}$$

b)
$$\frac{dL}{dt} = \dot{q} \left(\frac{d}{dt} \frac{\partial L}{\partial \dot{q}} - \frac{d^2}{dt^2} \frac{\partial L}{\partial \ddot{q}} \right) + \frac{\partial L}{\partial \dot{q}} \ddot{q} + \frac{\partial L}{\partial \ddot{q}} \dddot{q}$$

$$\frac{dL}{dt} = \frac{d}{dt} \left(\dot{q} \frac{\partial L}{\partial \dot{q}} - \dot{q} \frac{d}{dt} \frac{\partial L}{\partial \ddot{q}} + \ddot{q} \frac{\partial L}{\partial \ddot{q}} \right)$$

$$\Rightarrow \frac{d}{dt} \left[\underbrace{\dot{q} \left(\frac{\partial L}{\partial \dot{q}} - \frac{d}{dt} \frac{\partial L}{\partial \ddot{q}} \right) + \ddot{q} \frac{\partial L}{\partial \ddot{q}} - L}_{\text{Conserved Quantity (Analogy of Energy)}} \right] = 0$$

Conserved Quantity (Analogy of Energy)

For $L(q, \dot{q}, \ddot{q}, t)$ we found
$$\frac{d}{dt} \left[\underbrace{\dot{q} \frac{\partial L}{\partial \dot{q}} - L}_E \right] = 0$$

~~Can't~~ Can't Express
$$\frac{d}{dt} \left[\dot{q} \left(\frac{\partial L}{\partial \dot{q}} - \frac{d}{dt} \frac{\partial L}{\partial \ddot{q}} \right) + \ddot{q} \frac{\partial L}{\partial \ddot{q}} - L \right] = 0$$

or
$$\frac{d}{dt} \left[E - \underbrace{\frac{d}{dt} \frac{\partial L}{\partial \ddot{q}} + \ddot{q} \frac{\partial L}{\partial \ddot{q}}}_{\text{Corrections to Energy}} \right] = 0$$

Corrections to Energy.

(7) continued #1

(Alternative method (i) to I.B.P.)

$$\delta S = \int_{t_1}^{t_2} \left[\left(\frac{\partial L}{\partial q} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}} \right) \delta q + \frac{\partial L}{\partial \ddot{q}} \frac{d^2 \delta q}{dt^2} \right] dt = 0$$

Using i.b.p twice more gets us a term ~~of~~ of the

Form $\left(\frac{\partial L}{\partial \ddot{q}} \frac{d}{dt} \delta q \right) \Big|_{t_1}^{t_2}$, which ~~is~~ see no reason

for it to be zero. Instead, we are going to apply

a math theorem: Given continuous functions f_0, f_1, \dots, f_n on (a,b) satisfying

$$\int_a^b [f_0(x)h(x) + f_1(x)h'(x) + f_2(x)h''(x) + \dots + f_n(x)h^{(n)}(x)] dx = 0 \text{ for}$$

all smooth functions $h(x)$ on (a,b) , then there exist continuously

differentiable functions u_0, u_1, \dots, u_{n-1} on (a,b) such that

$$f_0 = u_0', f_1 = u_0 + u_1', \dots, f_{n-1} = u_{n-2} + u_{n-1}', f_n = u_{n-1}, \text{ everywhere in } (a,b).$$

$$\text{For us, } f_0 = \left(\frac{\partial L}{\partial q} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}} \right), f_1 = 0, f_2 = \frac{\partial L}{\partial \ddot{q}}, h(x) = \delta q(t).$$

This theorem tells us:

$$f_0 = u_0', f_1 = u_0 + u_1', f_2 = u_1$$

$$\text{Using } f_1 = 0, 0 = u_0' + u_1' \Rightarrow 0 = f_0' + f_2''$$

Equation of motion.

$$\Rightarrow \boxed{0 = \left(\frac{\partial L}{\partial q} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}} \right) + \frac{d^2}{dt^2} \frac{\partial L}{\partial \ddot{q}}}$$

Note: this is equivalent to setting $\left(\frac{\partial L}{\partial \ddot{q}} \frac{d}{dt} \delta q \right) \Big|_{t_1}^{t_2} = 0$.

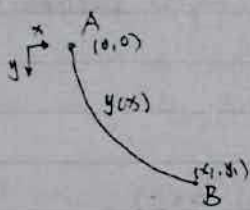
General
Case

... page on
Fundamental Lemma of
Calculus of Variations

problem of brachistochrone

Kanseng Zhang

8)



Suppose A is the origin $(0,0)$ and B has coordinates (x_1, y_1) . $y=y(x)$ is the curve along which we could descend in the least time from A to B.

So $ds = \sqrt{1 + (\frac{dy}{dx})^2} dx$ along this curve.

And according to energy conservation, we have

$$\frac{1}{2} m (\frac{ds}{dt})^2 = mgy$$

So $dt = \sqrt{\frac{1 + (\frac{dy}{dx})^2}{2gy}} dx$

We could get the total time taken from A to B.

$$T = \int_0^{x_1} \sqrt{\frac{1 + (\frac{dy}{dx})^2}{2gy}} dx$$

so we take $\delta T = 0$ for getting least time T.

It's kind of Lagrangian where

$$L' = \sqrt{\frac{1 + (\frac{dy}{dx})^2}{2gy}}$$

so we get equation:

$$\begin{aligned} L' &\Leftrightarrow L \\ x &\Leftrightarrow \delta \\ y &\Leftrightarrow X \end{aligned}$$

$$\frac{d}{dx} \left(\frac{\partial L'}{\partial (\frac{dy}{dx})} \right) = \frac{\partial L'}{\partial y}$$

However it's hard to solve it. We consider

$$H' = \frac{\partial L'}{\partial (\frac{dy}{dx})} \cdot \frac{dy}{dx} - L' = \frac{(\frac{dy}{dx})^2}{\sqrt{2gy} \sqrt{1 + (\frac{dy}{dx})^2}} - \sqrt{\frac{1 + (\frac{dy}{dx})^2}{2gy}} = \frac{-1}{\sqrt{2gy} \sqrt{1 + (\frac{dy}{dx})^2}}$$

We have $\frac{dH'}{dx} = \frac{\partial L'}{\partial (\frac{dy}{dx})} \frac{d^2y}{dx^2} + \frac{dy}{dx} \frac{d}{dx} \left(\frac{\partial L'}{\partial (\frac{dy}{dx})} \right) - \frac{\partial L'}{\partial y} \frac{dy}{dx} - \frac{\partial L'}{\partial (\frac{dy}{dx})} \frac{d^2y}{dx^2}$

= 0

So $H' = \frac{-1}{\sqrt{2gy} \sqrt{1 + (\frac{dy}{dx})^2}} = \text{const} \triangleq C$

So $y \left[1 + \left(\frac{dy}{dx} \right)^2 \right] = \frac{1}{2gc^2} \triangleq 2r$ Impose variable θ where $x = x_0$

Suppose $\frac{dy}{dx} = \cot \frac{\theta}{2}$

we have $y = 2r \sin^2 \frac{\theta}{2} = r(1 - \cos \theta)$

Then $dy = r \sin \theta d\theta$

we have $\frac{dx}{d\theta} = r \sin \theta \cdot \tan \frac{\theta}{2} = 2r \sin^2 \frac{\theta}{2} = r(1 + \cos \theta)$

So $x = r(\theta - \sin \theta) + x_0$

Plug $(0,0)$ into x, y we have

$$\begin{cases} x = r(\theta - \sin \theta) \\ y = r(1 + \cos \theta) \end{cases}$$

where r could be given by

$B(x_1, y_1)$, θ is a real parameter, corresponding to the angle through which the rolling circle has rotated.

Cycloid: A curve traced by a point on the rim of a

circular wheel as it rolls along a straight line without slipping.

$$x = r \cos^{-1} \left(1 - \frac{y}{r} \right) - \sqrt{y(2r-y)}$$