

HW problem week 6

Your turned in assignment should be clearly written and easy to follow! Learning how to explain your work in a way that is as easy as possible to follow is an important part of your training as a physicist. An incoherent mess of equations with a correct final answer could receive less points than a solution which is clearly explained at every step but has an algebra mistake somewhere. Once you've solved the problem, you can rewrite it on a new piece of paper for clarity if you need to.

The normalized eigenfunctions of the infinite square well potential are $\psi_n(x) = \sqrt{\frac{2}{L}} \sin(n\pi x/L)$ inside the well.

1. Use these eigenstates to solve the time dependent schrodinger equation to find $\Psi_n(x, t)$. Using the full time dependent wave function, calculate $\langle x \rangle(t)$ and $\langle p \rangle(t)$ in the n 'th energy eigenstate.

Once we have an eigenstate ψ_n with energy ϵ_n , the time dependence is just an oscillating exponential. So $\Psi_n(x, t) = \psi_n(x)e^{-i\epsilon_n t/\hbar} \equiv \psi_n(x)e^{-i\omega_n t}$. To calculate the expectation value:

$$\langle x \rangle = \int dx \Psi_n^*(x, t)x\Psi_n(x, t) = \int dx \psi_n(x)^* e^{+i\omega_n t} x \psi_n(x) e^{-i\omega_n t} = \int dx \psi_n^*(x)x\psi_n(x)$$

$$\langle p_{op} \rangle = \int dx \psi_n(x)^* e^{+i\omega_n t} (-i\hbar\partial_x)\psi_n(x) e^{-i\omega_n t} = \int dx \psi_n^*(x)(-i\hbar\partial_x)\psi_n(x)$$

We've calculated both of the expressions at the right hand side in class/discussion/problem session: $\langle x \rangle = L/2$ and $\langle p \rangle = 0$. Neither of them have any time dependence. This is generally true whenever we calculate an expectation value in an energy eigenstate.

2. Calculate the average energy $\langle E \rangle(t)$ and the average square energy $\langle E^2 \rangle(t)$ to find the uncertainty in the energy $\Delta E = \sqrt{\langle E^2 \rangle - \langle E \rangle^2}$. Hint: use the fact that the ψ_n are eigenstates of \hat{H} .

Following the hint, $\hat{H}\psi_n = \epsilon_n\psi_n$ so $\hat{H}^2\psi_n = \hat{H}(\hat{H}\psi_n) = \hat{H}\epsilon_n\psi_n = \epsilon_n\hat{H}\psi_n = \epsilon_n^2\psi_n$. The average energy is

$$\langle E \rangle = \int dx \Psi^* \hat{H} \Psi = \epsilon_n \int dx \Psi^* \Psi = \epsilon_n$$

and the average square energy is

$$\langle E^2 \rangle = \int dx \Psi^* \hat{H}^2 \Psi = \epsilon_n^2 \int dx \Psi^* \Psi = \epsilon_n^2.$$

So the energy uncertainty is $\sqrt{\langle E^2 \rangle - \langle E \rangle^2} = 0$. There is no uncertainty in the energy, because we are in an energy eigenstate.

3. Now consider the superposition state

$$\Phi(x, t) = \frac{1}{\sqrt{2}} (\Psi_1(x, t) + \Psi_4(x, t)).$$

Verify that this state is normalized.

$$\begin{aligned} \int dx \Phi^* \Phi &= \frac{1}{2} \int dx (\psi_1(x)^* e^{i\omega_1 t} + \psi_4(x)^* e^{i\omega_4 t})(\psi_1(x) e^{-i\omega_1 t} + \psi_4(x) e^{-i\omega_4 t}) \\ &= \frac{1}{2} \left(\underbrace{\int dx |\psi_1|^2}_1 + \underbrace{\int dx |\psi_2|^2}_1 + e^{i(\omega_1 - \omega_4)t} \underbrace{\int dx \psi_1^* \psi_4}_{0, \text{check it}} + e^{-i(\omega_1 - \omega_4)t} \underbrace{\int dx \psi_1 \psi_4^*}_0 \right) = 1 \end{aligned}$$

4. For this new state Φ , calculate $\langle x \rangle(t)$ and $\langle p \rangle(t)$. Would you call the state $\Phi(x, t)$ a stationary state? Why or why not?

Look up at the solution for part 3, and imagine sandwiching an x or a p_{op} in between Φ^* and Φ . We can save ourselves a little bit of effort since we already know that $\langle x \rangle = L/2$ and $\langle p \rangle = 0$ if we evaluate their expectation value in a given eigenstate, so we only have to calculate the cross terms $\int \psi_1(x, \hat{p}) \psi_4$ and $\int \psi_4(x, \hat{p}) \psi_1$. The integrals for $\langle x \rangle$ turn out to be the same so I will just do one of them: they can be done by using trig product formulas to turn $\sin A \sin B$ into a combination of $\cos A \pm B$ and then using integration by parts. This is what is sometimes called ‘straightforward yet tedious’:

$$\int dx \psi_1 x \psi_4 = \frac{2}{L} \int dx x \sin\left(\frac{\pi x}{L}\right) \sin\left(\frac{4\pi x}{L}\right) = -\frac{32L}{225\pi^2}$$

We also need

$$\begin{aligned} \int dx \psi_1 (-i\hbar \partial_x) \psi_4 &= -i\hbar \frac{2}{L} \frac{4\pi}{L} \int dx \sin(\pi x/L) \cos(4\pi x/L) = -\frac{16\hbar}{15iL}, \\ \int dx \psi_4 (-i\hbar \partial_x) \psi_1 &= -i\hbar \frac{2}{L} \frac{\pi}{L} \int dx \cos(\pi x/L) \sin(4\pi x/L) = \frac{16\hbar}{15iL} \end{aligned}$$

Finally, putting it all together,

$$\langle x \rangle(t) = \frac{L}{2} - \frac{64L\hbar}{225\pi^2} \cos(\omega_{14}t)$$

where $\omega_{14} = \omega_1 - \omega_4$ and

$$\langle p \rangle(t) = \frac{32\hbar}{15L} \sin \omega_{41}t.$$

5. Repeat part 2 for the state Φ . Hint: be careful in how you apply the hint from part 2.

The equation $\hat{H}\psi_n = \epsilon_n\psi_n$ applies to each energy eigenstate individually (note that the function Φ is **not** an energy eigenstate). So when we calculate $\langle H \rangle$, the cross terms will drop out again. Explicitly:

$$\begin{aligned} & \frac{1}{2} \int dx (\psi_1(x)^* e^{i\omega_1 t} + \psi_4(x)^* e^{i\omega_4 t}) \hat{H} (\psi_1(x) e^{-i\omega_1 t} + \psi_4(x) e^{-i\omega_4 t}) \\ &= \frac{1}{2} \int dx (\psi_1(x)^* e^{i\omega_1 t} + \psi_4(x)^* e^{i\omega_4 t}) (\epsilon_1 \psi_1(x) e^{-i\omega_1 t} + \epsilon_4 \psi_4(x) e^{-i\omega_4 t}) \\ & \rightsquigarrow \frac{1}{2} (\epsilon_1 + \epsilon_4) \end{aligned}$$

(the squiggly arrow includes using the fact that the wave functions ψ_n are normalized, among other things). By a similar story (use $\hat{H}^2 = \hat{H}\hat{H}$ and apply them one after another to the functions on the right), find

$$\langle \hat{H}^2 \rangle = \frac{1}{2} (\epsilon_1^2 + \epsilon_4^2).$$

Now, $(\Delta E)^2 = \frac{1}{2} (\epsilon_1^2 + \epsilon_4^2) - \left(\frac{1}{2} (\epsilon_1 + \epsilon_4)\right)^2 = \frac{1}{4} (\epsilon_1 - \epsilon_4)^2$.

6. Using the time-energy uncertainty principle $\Delta E \Delta t > \hbar/2$, estimate approximately how much time the particle spends in a particular eigenstate state before flipping to the other one.

Setting $\Delta E \Delta t = \hbar/2$ we get

$$\Delta t = \frac{\hbar}{2} \frac{2}{|\epsilon_1 - \epsilon_4|} = \frac{\hbar}{|\Delta E_{14}|}.$$

There is a good lesson to learn here which is that there is a relationship in quantum mechanics between energy difference of a process, and the characteristic time scale over which that process occurs, as $\tau \sim \frac{\hbar}{\Delta E}$ or $\frac{1}{\omega}$. In this problem, we see that the energy difference between the two states **literally is** the frequency of the oscillations in the nonstationary state.