

Thermal Equilibrium Fluctuations

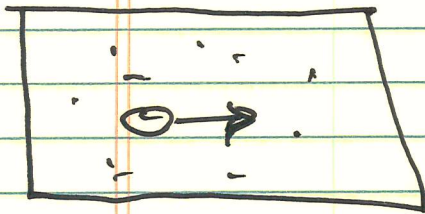
i.) Basic Ideas:

- dilute gas charged particles, quasi-neutral

$$T > e^2 n^{1/3} \Leftrightarrow 1/n \lambda_D^3 \ll 1$$

- ~ Thermal equilibrium fluctuations;

Balance: \rightarrow fluctuation vs. dissipation,
also Brownian Motion
 \rightarrow absorption vs. emission.



① emission

- discrete particle in
viscous fluid emits
waves

- Cerenkov emission, a/c
boat wake in water.

$$\underline{v} \cdot \underline{D} = \underline{v} \cdot \underline{E} = 4\pi q \rho(x - x(t))$$

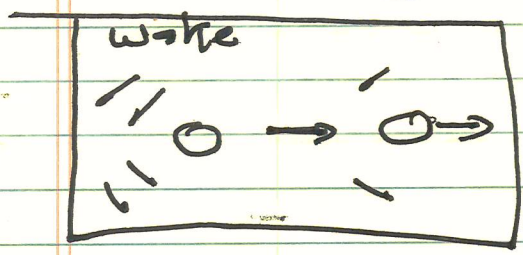
- emission \Rightarrow fluctuation

- kinetic energy of particle ultimately
coupled to collective response of
the medium.

- particle slows down due to
Cerenkov emission, \rightarrow drag

② Cerenkov emitted waves damped via Landau damping

⇒ absorption



emitted waves scatter other particles.

∞ Thermal equilibrium fluctuations result from detailed balance of:

- Cerenkov emission waves from discrete particle motion,
- absorption of wave energy via Landau damping of Ulesov fluid

Note: Particles are 'schizophrenic' = double agent

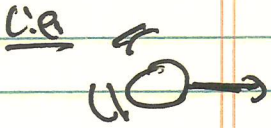
- act as discrete particle emitter
- act as element of Ulesov fluid absorber.

Here:

- assume confined system → no outgoing waves, radiative damping
- $\lambda_{damp} \sim \frac{c}{|\delta_{damp}|}$ ← L syst
- finite damping length

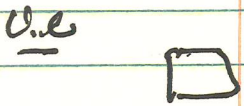
Double Agents:

- "emitter": a discrete particle moving along some specified (unperturbed) orbit



an identifiable 'pea' in a 'pea soup' composed of other peas

- "absorber": an element of the Vlasov fluid responding to, and damping (by Landau resonance) other discrete particles.



a 'crushed pea' element of the "pea soup" of the Vlasov fluid.

→ Thermal Equilibrium Plasma:

soup/gas of 'dressed' test particles
 ↓
 Vlasov fluid

⇒ Test Particle Model

Re: Pea soup model of Viscous Fluid:

"Every Pea in the soup acts like (fluid) soup for all the other peas"

Useful Analogy:

Here: Brownian Motion

dV/dt + \gamma V = F/M \quad \gamma = \delta/M

Brownian Motion | Plasma at Equilibrium

Table with 4 rows: Fluctuation, Excitation, Emission/Source, Absorption. Columns: Brownian Motion, Plasma at Equilibrium.

ii.) TPM For Fluctuation Spectrum

Basic ideas:

- dual identity as emitter and absorber
- Fluctuations weak \rightarrow unperturbed orbits valid (Contrast Turbulence),

Ala Debye length calculation, with stationary cons:

$$dF = F^C + \tilde{F}$$

\downarrow \downarrow
 coherent discrete
 Vlasov source
 response

$$dF = \frac{k \int_m E_{k,\omega} \langle \delta F \rangle dV}{-i(\omega - kv)} + |e| d(x - x(t)) d(v - v(t))$$

i.e.

$$\underline{D.P.} = \underline{D.} \underline{G} \underline{E} = 4\pi |e| \underline{d}(x - x(t))$$

equiv to:

$$\underline{D^2} \phi = 4\pi n_0 |e| \int dV dF^C + 4\pi n_0 |e| \int dV \tilde{F}$$

so

$$\hat{\phi}_{k,\omega} = \frac{4\pi n_0 k l}{k^2} \int dv \tilde{f} / \epsilon(k,\omega) \quad (\downarrow \uparrow)$$

$$\epsilon(k,\omega) = 1 + \frac{\omega_p^2}{\omega} \int dv \frac{\partial \langle f \rangle / \partial v}{\omega - kv}$$

check:

$$\int dv \tilde{f}_u = \int dv \int dx e^{-ikx} |e| d(x - x(t))$$

$$\underline{x} = \underline{x}_0 + \underline{v}t$$

$$\int dv \tilde{f}_u = \int |e| e^{-ikvt}$$

so

$$\hat{\phi}_k(t) = \underline{\underline{\epsilon}}^{-1}(k,t) \frac{4\pi n_0 k l}{k^2} \int dv e^{-ikvt}$$

strictly:

$$\underline{\underline{\epsilon}}(k,t) \hat{\phi}_u(t) = \frac{4\pi n_0 k l}{k^2} \int dv e^{-cvt}$$

so

$$\hat{\phi}_u(t) = \underline{\underline{\epsilon}}^{-1}(k,t) \frac{4\pi n_0 k l}{k^2} e^{-ikvt} + \phi_u^{homog.} e^{-i\omega_p t}$$

$$\omega_{ij} = \omega_r(k) + i \omega_i(k) \rightarrow \text{eigenmode frequency}$$

Now:

- time asymptotically ~~...~~

$\omega_i < 0 \rightarrow$ collective modes (free response) damp.

- only discretizers driven plots survive

- moving toward, but not to marginal stability \rightarrow Index $\rightarrow \infty$

- for sufficient source strength and/or weak enough damping, amplification to nonlinearity occurs

- if unstable modes require nonlinear damping to balance nonlinear noise (mode coupling)

i.e. 

so, for $\tilde{F} \equiv$ discretized fluctn.

$$\langle \hat{\phi}^2 \rangle_{k, \omega} = \left(\frac{4\pi n_0 k l}{k^2} \right)^2 \int d\underline{v}_1 \int d\underline{v}_2 \frac{\langle \tilde{F}(\underline{v}_1) \tilde{F}(\underline{v}_2) \rangle_{k, \omega}}{|\epsilon(k, \omega)|^2}$$

→ thermal equilibrium fluctuation spectrum

→ all content → $\langle \hat{\phi}^2 \rangle_{k, \omega}$
 $\epsilon(k, \omega)$

What of Discrete Noise?

$\langle \tilde{F}(\underline{v}_1) \tilde{F}(\underline{v}_2) \rangle \equiv$ discrete particle correlation function

will show:

$$\langle \tilde{F}(\underline{v}_1) \tilde{F}(\underline{v}_2) \rangle = \frac{\langle F \rangle}{N} \delta(\underline{x}_1 - \underline{x}_2) \delta(\underline{v}_1 - \underline{v}_2)$$

(iii) Discreteness Noise

For discreteness correlation function:

$$\tilde{F}(l) = \frac{1}{N} \sum_{i=1}^N \delta(x_i - x_i(t)) \delta(v_i - v_i(t))$$

\downarrow
 \hookrightarrow normalization

$l \equiv x_i, v_i \rightarrow$ phase space point

$$\tilde{F}(l) = \tilde{F}(x_i, v_i, t) \rightarrow \text{discrete particle dist.}$$

Can write $\tilde{F}(z)$, as well.

Now, for:

norm

$$\langle A \rangle = \int dx_i \int dv_i \rho \langle F \rangle A$$

i.e. av. over distribution, uncorrelated particles.
 \uparrow
dilute.

$$\langle \tilde{F}(1) \tilde{F}(2) \rangle$$

$$= \int dx_i \int dv_i \rho \langle F \rangle \left(\frac{1}{N} \sum_{i=1}^N \delta(x_i - x_i(t)) \delta(v_i - v_i(t)) \right)$$

$$\times \left(\frac{1}{N} \sum_{j=1}^N \delta(x_j - x_j(t)) \delta(v_j - v_j(t)) \right)$$

$$= \int dx_i \int dv_i \frac{\rho \langle F \rangle}{N} \sum_{\substack{i=1 \\ j=1}}^N \left[\delta(x_i - x_i(t)) \delta(x_j - x_j(t)) \delta(v_i - v_i(t)) \delta(v_j - v_j(t)) \right]$$

$\langle \tilde{f}(1) \tilde{f}(2) \rangle \neq 0$ only if arguments interchangeable

so

$$\langle \tilde{f}(1) \tilde{f}(2) \rangle = \int dx_1 \int dv_1 \left\langle \frac{f}{N} \right\rangle \left[\delta(\underline{x}_1 - \underline{x}_2) \delta(\underline{v}_1 - \underline{v}_2) + \delta(\underline{v}_1 - \underline{v}_2) \delta(\underline{x}_1 - \underline{x}_2) \right]$$

$$= \left\langle \frac{f}{N} \right\rangle \delta(\underline{x}_1 - \underline{x}_2) \delta(\underline{v}_1 - \underline{v}_2)$$

$$\langle \tilde{f}(1) \tilde{f}(2) \rangle = \left\langle \frac{f}{N} \right\rangle \delta(\underline{x}_1 - \underline{x}_2) \delta(\underline{v}_1 - \underline{v}_2)$$

- equal time discreteness correlation function.
- particles uncorrelated unless they are same particle, no width, no range.

Now, must obtain $\langle \tilde{f}(1) \tilde{f}(2) \rangle_{\eta, \omega}$
 \Rightarrow Transformed correlation function,

Now, need:

$$\langle \tilde{F}(1) \tilde{F}(2) \rangle_k = \int e^{-ik(x_2 - x_1)} \langle \tilde{F}(1) \tilde{F}(2) \rangle$$

and for time transform:

$$\langle \tilde{F}(1) \tilde{F}(2) \rangle_{k, \omega} = \int_0^{\infty} dt e^{i\omega t} U(2, T) \int dx e^{-i(x_2 - x_1)} \langle \tilde{F}(1) \tilde{F}(2) \rangle$$

$$+ \int_0^{\infty} dt e^{i\omega t} U(1, -T) \int dx e^{-ik(x_2 - x_1)} \langle \tilde{F}(1) \tilde{F}(2) \rangle dx$$

$$\langle \tilde{F}(1, t^1) \tilde{F}(2, t^2) \rangle = \langle \tilde{F}(1) \tilde{F}(2) [t - t^1] \rangle$$

so $T = t - t^1 \rightarrow$ can
push 2 forward T
push 1 back T

"push" $\equiv U \rightarrow$ operator evolving/pushing particle along unperturbed orbit

v.e. $U: x \rightarrow x + vT$

so (2) \rightarrow (3)

$$\textcircled{1} = \int_0^{\infty} dt \int dx e^{-ik(x_2 - x_1)} e^{i(\omega - kv_2)t} \langle \tilde{F}(1) \tilde{F}(2) \rangle dx$$

$$= \int_0^{\infty} dt e^{i(\omega - kv_2)t} \langle \tilde{F}(1) \tilde{F}(2) \rangle_k$$

0

$$\begin{aligned}
 \textcircled{a} &= \frac{-1}{i(\omega - kv_2)} \langle \tilde{f} \tilde{f} \rangle_{\omega} \\
 &= \frac{i}{(\omega - kv_2)} \langle \tilde{f} \tilde{f} \rangle_{\omega}
 \end{aligned}$$

real part

$$\textcircled{a} = \pi \delta(\omega - kv_2) \langle \tilde{f} \tilde{f} \rangle_{\omega}$$

Similarly,

$$\textcircled{b} = \pi \delta(\omega - kv_1) \langle \tilde{f} \tilde{f} \rangle_{\omega}$$

Now, ultimately seek:

$$\int dv_1 \int dv_2 \langle \tilde{f}(1) \tilde{f}(2) \rangle_{\omega}$$

$$\int dv_1 \int dv_2 = \int_{\text{com}} dv_+ \int_{\text{relative}} dv_-$$

$$\langle \tilde{f}(1) \tilde{f}(2) \rangle = \frac{\langle f \rangle}{n} \delta(x_-) \delta(v_-)$$

so

$$\int dV \langle \tilde{F}(u) \tilde{F}(z) \rangle = 2\pi \frac{\langle F \rangle}{N} \delta(\omega - kv_z)$$

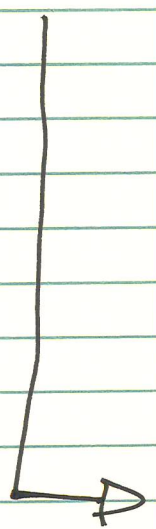
and,

$$\begin{aligned} \langle \frac{\tilde{n}}{n_0} \frac{\tilde{n}}{n_0} \rangle_{k, \omega} &= \int dV \frac{\langle F \rangle}{N} 2\pi \delta(\omega - kv_z) \\ &\equiv C(k, \omega) \\ &\downarrow \\ &\text{emission correlator} \end{aligned}$$

and so

\rightarrow v_{the} extracted.

$$\begin{aligned} C(k, \omega) &= \frac{2\pi}{N|k| v_{the}} \langle \bar{F}(\omega/k v_{the}) \rangle \\ &= \frac{2\pi}{N} T_{str} \langle \bar{F}(\omega/k v_{the}) \rangle \\ &\quad \downarrow \\ &\quad \text{streaming time} \\ &\quad \sim (2/|k| v_{the}) \end{aligned}$$



\rightarrow discreteness noise has Maxwellian Doppler spectrum (obvious).

Thermal Equilibrium Spectrum

$$\langle \hat{\phi}^2 \rangle_{k, \omega} = \left(\frac{4\pi N_0 e^2}{k^2} \right)^2 \frac{1}{N_0 k v_{th}} \frac{\langle \bar{F}(\omega/k v_{th}) \rangle}{|\epsilon(k, \omega)|^2}$$

Spectrum set by:

- particle emission distribution
 $\sim e^{-\omega^2/k^2 v_{th}^2}$

→ collective screening, i.e.

$$\omega \gg k v_{th}, \quad \epsilon \approx 1 - \frac{\omega_p^2}{\omega^2} + i\epsilon_{IM}$$

$$\omega \lesssim k v_{th}, \quad \epsilon \approx 1 + \frac{1}{k^2 \lambda_D^2} + i\epsilon_{IM}$$

- Coulomb Factor $\left(\frac{4\pi N_0 e^2}{k^2} \right)$

- $\mathcal{T}_{str} \rightarrow$ (time transform)

N.B:

- Collective response strongest at wave resonance ($\omega_p \sim \omega \sim kv$)

\Rightarrow expect peak.

- $\omega \gg \omega_p$, \Rightarrow high frequency limit

Noise decouples from collective dynamics
 $\epsilon \rightarrow 1$

$$\langle \phi^2 \rangle_{\omega} \underset{\omega \gg \omega_p}{\approx} n_0 \left(\frac{4\pi e}{k^2} \right)^2 \frac{2\pi}{|k| v_{th}} e^{-\omega^2 / k^2 v_{th}^2}$$

- $\omega < \omega_p \Rightarrow$ low frequency limit

\Rightarrow static plasma screening, expect

$$\langle \phi^2 \rangle_{\omega} \underset{\omega \ll \omega_p}{\approx} n_0 \left(\frac{4\pi e}{k^2} \right)^2 \frac{2\pi}{|k| v_{th}} \frac{e^{-\omega^2 / k^2 v_{th}^2}}{(k^2 + 1/\lambda_D^2)^2}$$

Now, can write:

$$\frac{\langle E^2 \rangle_{\omega}}{8\pi} = \frac{4\pi^2 \rho_0 c^2}{k^2 |k|} \frac{F(\omega/k v_{th})}{|E_r|^2 + |E_{IM}|^2}$$

$$F(\omega/k v_{th}) \equiv \langle \bar{F} \rangle / v_{th} \sqrt{2\pi}$$

re-absorb factor.

$$F = F(u)$$

$$F' = dF/du$$

$$u = \omega/k$$

$$E_{IM} = \frac{\pi \omega \rho_0}{|k| k} F'$$

Now useful to make contact with usual $k_B T/2$ per d-o-f.

$$W_k = \int \frac{d\omega}{2\pi} \frac{\langle E^2 \rangle_{\omega}}{8\pi}$$

How do frequency integral:

Pole Approximation!

$$\frac{1}{|E|^2} \approx \frac{1}{\left[(\omega - \omega_0)^2 \left(\frac{\partial \epsilon}{\partial \omega} \right)^2 + |E_{IM}|^2 \right]}$$

\downarrow real frequency sets location pole
 \rightarrow width

$$\begin{aligned} \frac{1}{L^2} &= \frac{1}{|G_{IM}|} \left\{ \frac{|G_{IM}|}{(\omega - \omega_k)^2 \left| \frac{\partial G}{\partial \omega} \right|^2 + |G_{IM}|^2} \right\} \\ &= \frac{1}{|G_{IM}|} \left| \frac{\partial G}{\partial \omega} \right|^{-1} \pi \delta(\omega - \omega_k) \end{aligned}$$

So integrating:

$$\omega_k = \frac{m_0 \omega_{p0}}{2|k|} F / |F'|$$

$$= \frac{m_0 \omega_{p0}}{2|k|} F / \frac{\omega_{p0} F}{|k| v_{th0}^2} = T/2$$

$$\text{so } \boxed{\omega_k = kT/2} \quad \text{for } k\lambda_D \ll 1$$

→ but for $k\lambda_D \gg 1$

$$\epsilon \rightarrow 1 + 1/k^2 \lambda_D^2$$

no collective resonance.

80

$$W_n \approx \frac{k_B T}{2} / k^2 \lambda_D^2$$

for $k \lambda_D > 1$

↳ strong cut-off beyond λ_D .

Now, for total energy density:
(3D)

$$\langle E^2 / 8\pi \rangle = \int d^3k W_n$$

$$\approx \left(\frac{k_B T}{2} \right) k_{\max}^3$$

$k_{\max} \sim 1/\lambda_D \rightarrow$ screening kills $k \lambda_D > 1$

$$\langle E^2 / 8\pi \rangle = n \left(\frac{k_B T}{2} \right) / n \lambda_D^3$$

$$= \underbrace{(k_B T)}_{\text{kinetic energy density}} / \underbrace{n \lambda_D^3}_{\substack{\rightarrow \text{diluteness} \\ \rightarrow \# \text{ in Debye sphere}}}$$

②

Field Energy Density \sim $\frac{\text{Kinetic Energy Density}}{n \lambda_D^3}$

↳ diluteness factor.

$$\text{so } \frac{FED}{kED} = \frac{2}{n} \frac{3}{\lambda^3}, \text{ so should be}$$

→ Connect to Fluctuation - Dissipation Theorem. (Notation reverts).

$$\text{Recall: } G_{IM} = \frac{-\omega_p^2 \pi}{|k| k} \frac{\partial \langle \bar{\phi} \rangle}{\partial V} \Big|_{\omega/k}$$

$$= \frac{2\pi\omega}{k^2 v_{tho}^2} \frac{\omega_p^2}{|k| v_{tho}} \langle \bar{\phi}(\omega/k) \rangle$$

Maxwellian.

$$\langle \bar{\phi}(\omega/k) \rangle = k^2 v_{tho}^2 |k| / v_{tho} G_{IM} / 2\pi \omega \omega_p^2$$

Now have:

$$\langle \phi^2 \rangle_{k,\omega} = \frac{2\pi n}{|k| v_{tho}} \left(\frac{4\pi |e|}{k^2} \right)^2 \frac{\langle \bar{\phi}(\omega/k) \rangle}{|G(k,\omega)|^2}$$

so plugging in:

$$\langle \phi^2 \rangle_{k,\omega} = \frac{8\pi T}{k^2 \omega} \frac{\text{Im } G}{|G|^2}$$

and, $\left\langle \frac{\hat{E}^2}{8\pi} \right\rangle_{\omega} = \frac{T}{\omega} \frac{\text{Im} \epsilon}{|\epsilon|^2}$

- Fluctu - Dissip Thm for plasma
- restates form of spectrum
- relates fluctuations to T and dissipation in collective modes (Im ϵ)
- obviously consistent, by construction.

Comments:

- key to TPM is causality, and use of linear fc or UPD
- assumed weak fluctuations
i.e. $\underline{x}(t) = \underline{x}(0) + vt + \cancel{\sigma x}$

How weak?

- $\tau_{\text{coll}} \ll \tau_T$ correlation

For stable plasma, $F' \rightarrow 0$ from
 $F' < 0$, so $\langle \phi^2 \rangle \uparrow$, even for
stable system.

- akin critical opalescence.

- TPM fails \rightarrow lack of damping.