

Lecture 12 (Nov. 11)

- Inhomogeneous Sturm - Liouville equation (Eq. 9.7):

$$\mu(x) \frac{\partial^2 y}{\partial t^2} - \frac{\partial}{\partial x} \left[\tau(x) \frac{\partial y}{\partial x} \right] + v(x) y = \mu(x) \operatorname{Re} \left[f(x) e^{-i\omega t} \right]$$

Here the string is forced at frequency ω .

We write the solution as

$$y(x,t) = \operatorname{Re} \left[y(x) e^{-i\omega t} \right]$$

where

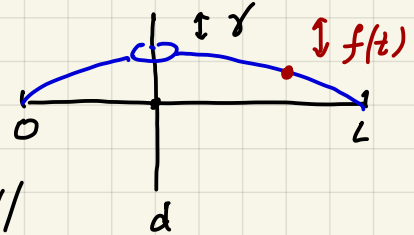
$$\left[\hat{K} - \omega^2 \mu(x) \right] y(x) = \mu(x) f(x)$$

could redefine as $\tilde{f}(x)$ but it is convenient if we include $\mu(x)$

with

$$\hat{K} = -\frac{d}{dx} \tau(x) \frac{d}{dx} + v(x)$$

$$\rightarrow y(x,t) = y_{\text{hom}}(x,t) + y_{\text{inh}}(x,t)$$



the Sturm - Liouville operator. Recall

$$\hat{K} \psi_n(x) = \omega_n^2 \mu(x) \psi_n(x)$$

$$\langle \psi_m | \psi_n \rangle = \int_{x_L}^{x_R} dx \mu(x) \psi_m^*(x) \psi_n(x) = \delta_{mn}$$

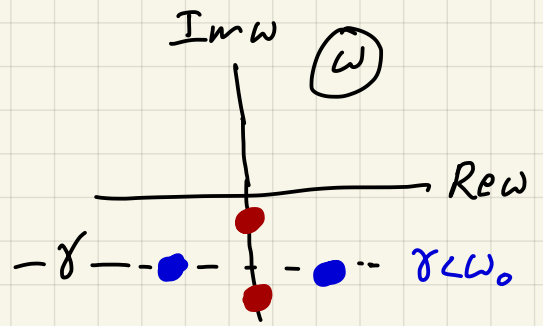
$$\mu(x) \sum_n \psi_n(x) \psi_n^*(x') = \delta(x-x')$$

Taking the inverse of $\hat{K} - \omega^2 \mu(x)$, we have that the inhomogeneous solution is

Scratch

Unforced, damped SHO:

$$\ddot{x} + 2\gamma\dot{x} + \omega_0^2 x = 0$$



Solⁿ: $x = A e^{-i\omega t} \Rightarrow -\omega^2 - 2i\gamma\omega + \omega_0^2 = 0$

$$\omega^2 + 2i\gamma\omega - \omega_0^2 = 0 \Rightarrow \omega = -i\gamma \pm \sqrt{\omega_0^2 - \gamma^2}$$

$e^{-i\omega_{\pm} t} \rightarrow 0$ as $t \rightarrow \infty$ due to $\gamma > 0$

$\gamma^2 < \omega_0^2 \Rightarrow$ underdamped, $\gamma^2 > \omega_0^2 \Rightarrow$ overdamped

Harmonic forcing:

$$f(t) = \int \frac{d\Omega}{2\pi} \hat{f}(\Omega) e^{-i\Omega t}$$

$$\ddot{x} + 2\gamma\dot{x} + \omega_0^2 x = \hat{f}(\Omega) e^{-i\Omega t}$$

$\hat{x}(\Omega) e^{-i\Omega t}$

Solⁿ: $x(t) = x_{\text{hom}}(t) + x_{\text{inh}}(t)$

$$A_+ e^{-i\omega_+ t} + A_- e^{-i\omega_- t} \rightarrow 0$$

$$(\omega_0^2 - 2i\gamma\Omega - \Omega^2) \hat{x}(\Omega) = \hat{f}(\Omega)$$

Single frequency: $x_{\text{inh}}(t) = A(\Omega) \cos[\Omega t + \delta(\Omega)]$

amplitude: $A(\Omega) = \left[(\omega_0^2 - \Omega^2)^2 + 4\gamma^2 \Omega^2 \right]^{-1/2}$

phase shift: $\delta(\Omega) = \tan^{-1} \left(\frac{2\gamma\Omega}{\Omega^2 - \omega_0^2} \right)$

$$y_{\text{inh}}(x) = \int_{x_L}^{x_R} dx' \mu(x') G_\omega(x, x') f(x')$$

where $G_\omega(x, x')$ is the Green's function, satisfying

$$[\hat{K} - \omega^2 \mu(x)] G_\omega(x, x') = \delta(x - x')$$

I.e. $G_\omega(x, x') = [\hat{K} - \omega^2 \mu]_{x, x'}^{-1}$. We may write

$$G_\omega(x, x') = \sum_n \frac{\psi_n(x) \psi_n^*(x')}{\omega_n^2 - \omega^2}, \quad [G_\omega] = \frac{T^2}{M}$$

You can read about how to obtain $G_\omega(x, x')$ without having to do the infinite sum over all the eigenfunctions in §9.7.1. For now, I just quote the result for the case where $\mu(x) = \mu$, $\tau(x) = \tau$, $v(x) = 0$, and $[x_L, x_R] = [0, L]$. Then

$$G_\omega(x, x') = \frac{\sin(\omega x_</c>/c) \sin(\omega(L - x_>/c)}{(\omega \tau/c) \sin(\omega L/c)}$$

where $x_< = \min(x, x')$ and $x_> = \max(x, x')$, $c = \sqrt{\frac{\tau}{\mu}}$

Example: Let $f(x) = f_0 \delta(x - x_0)$. Then

$$y_{\text{inh}}(x) = \mu f_0 G_\omega(x, x_0)$$

Note that there are no constants of integration.

The full solⁿ is then

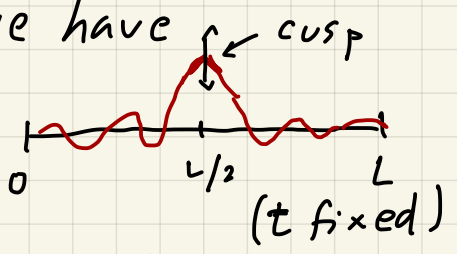
homogeneous solⁿ (e.g. Bernoulli)

inhomogeneous solⁿ

$$y(x,t) = y_{\text{hom}}(x,t) + y_{\text{inh}}(x,t)$$

The initial conditions enter in $y_{\text{hom}}(x,t)$ as we have learned from the Bernoulli solution. If there is some small damping, then at long times we have

$$y(x, t \gg \gamma^{-1}) = y_{\text{inh}}(x, t)$$

$$= \mu \int_0 G_{\omega}(x, x_0) \cos(\omega t)$$


where γ is the damping rate (i.e. rate of energy loss for unforced system). If $x_0 = \frac{1}{2}L$, then

$$G_{\omega}(x, \frac{1}{2}L) = \frac{c}{2\omega L \cos(\omega L/2c)} \times \begin{cases} \sin(\omega x/c) & \text{if } x < L/2 \\ \sin(\omega(L-x)/c) & \text{if } x > L/2 \end{cases}$$

Note that $y_{\text{inh}}(x,t)$ is continuous at $x = \frac{1}{2}L$ but its spatial derivative $y'_{\text{inh}}(x,t)$ is discontinuous at $x = \frac{1}{2}L$.

- Continua in higher dimensions : $h(\vec{x}, t)$ displacement
Generalization of wave operator : e.g. drum head :

$$\hat{K} = - \frac{\partial}{\partial x^{\alpha}} T_{\text{op}}(\vec{x}) \frac{\partial}{\partial x^{\beta}} + v(\vec{x})$$



This arises from

$$\mathcal{L} = \frac{1}{2} \mu(\vec{x}) \left(\frac{\partial h}{\partial t} \right)^2 - \frac{1}{2} \tau_{\alpha\beta}(\vec{x}) \frac{\partial h}{\partial x^\alpha} \frac{\partial h}{\partial x^\beta} - \frac{1}{2} v(\vec{x}) h^2$$

The wave equation is

$$\hat{K} h(\vec{x}, t) = -\mu(\vec{x}) \frac{\partial^2}{\partial t^2} h(\vec{x}, t)$$

Since $[\hat{K}, \partial_t] = 0$, solutions may be written as

$$h(\vec{x}, t) = \text{Re} \left[h(\vec{x}) e^{-i\omega t} \right]$$

where

$$\left[\hat{K} - \omega^2 \mu(\vec{x}) \right] h(\vec{x}) = 0$$

This is again an eigenvalue equation, with solutions

$$\psi_n(\vec{x}) \Rightarrow \hat{K} \psi_n(\vec{x}) = \omega_n^2 \mu(\vec{x}) \psi_n(\vec{x})$$

The eigenfunctions and eigenvalues satisfy

$$\langle \psi_m | \psi_n \rangle = \int_{\Omega} d^d x \mu(\vec{x}) \psi_m^*(\vec{x}) \psi_n(\vec{x}) = \delta_{mn}$$

$$\mu(\vec{x}) \sum_n \psi_n(\vec{x}) \psi_n^*(\vec{x}') = \delta(\vec{x} - \vec{x}')$$

where the medium is confined to a region $\Omega \subset \mathbb{R}^d$.

We must also apply boundary conditions of the form

$$(i) h(\vec{x})|_{\partial\Omega} = 0, \text{ where } \partial\Omega = \text{boundary of } \Omega$$

$$(ii) \tau(\vec{x}) \hat{n} \cdot \vec{\nabla} h|_{\partial\Omega} = 0, \text{ where } \hat{n} \text{ is normal to } \partial\Omega$$

$$(iii) \text{PBCs, e.g. in a box of dim}^{\text{ns}} L_1 \times L_2 \times \dots \times L_d$$

$$(iv) [\alpha \psi(\vec{x}) + \beta \hat{n} \cdot \vec{\nabla} \psi(\vec{x})]_{\partial\Omega} = 0$$

The Green's function is

$$G_\omega(\vec{x}, \vec{x}') = \sum_n \frac{\psi_n(\vec{x}) \psi_n^*(\vec{x}')}{\omega_n^2 - \omega^2}$$

with

$$[\hat{K} - \omega^2 \mu(\vec{x})] G_\omega(\vec{x}, \vec{x}') = \delta(\vec{x} - \vec{x}')$$

The variational approach generalizes as well, with

$$\omega^2[\psi(\vec{x})] \equiv \frac{N[\psi(\vec{x})]}{D[\psi(\vec{x})]}$$

and

$$N[\psi(\vec{x})] = \int_{\Omega} d^d x \psi^*(\vec{x}) \left\{ \overbrace{-\frac{\partial}{\partial x^\alpha} \tau_{\alpha\beta}(\vec{x}) \frac{\partial}{\partial x^\beta} + v(\vec{x})}^{\hat{K}} \right\} \psi(\vec{x})$$

$$D[\psi(\vec{x})] = \int_{\Omega} d^d x \mu(\vec{x}) \psi^2(\vec{x})$$

Demanding $\delta\omega^2 = 0$ yields the wave equation

$$\hat{K} \psi(\vec{x}) = \omega^2 \mu(\vec{x}) \psi(\vec{x})$$

- Membranes : $z = h(x, y)$

The equation of a surface is $F(x, y, z) = z - h(x, y) = 0$.

Let the differential surface area be dS . The projection onto the (x, y) plane is then

$$dA = dx dy = \hat{n} \cdot \hat{z} dS = n^z dS$$

The unit normal is

$$\hat{n} = \frac{\vec{\nabla} F}{|\vec{\nabla} F|} = \frac{\hat{z} - \vec{\nabla} h}{\sqrt{1 + |\vec{\nabla} h|^2}} \quad (\text{note } \hat{z} \cdot \vec{\nabla} h = 0)$$

Thus,

$$dS = \frac{dx dy}{\hat{n} \cdot \hat{z}} = \sqrt{1 + |\vec{\nabla} h|^2} dx dy$$

We consider a model where before: $ds = \sqrt{1 + h'^2} dx$

$$U[h(x, y, t)] = \int dS \sigma = U_0 + \frac{1}{2} \int d^2x \sigma(\vec{x}) |\vec{\nabla} h|^2 + \dots$$

with σ the surface tension. Other energy functions are possible. The kinetic energy is

$$T[h(x, y, t)] = \frac{1}{2} \int d^2x \mu(\vec{x}) \left(\frac{\partial h}{\partial t} \right)^2$$

Thus

$$S = \int dt \int d^2x \mathcal{L}(h, \partial_t h, \vec{\nabla} h, t, \vec{x})$$

$$\mathcal{L} = \frac{1}{2} \mu(\vec{x}) (\partial_t h)^2 - \frac{1}{2} \sigma(\vec{x}) |\vec{\nabla} h|^2$$

The equations of motion are then

$$\frac{\partial \mathcal{L}}{\partial h} - \frac{\partial}{\partial t} \frac{\partial \mathcal{L}}{\partial \dot{h}} - \vec{\nabla} \cdot \frac{\partial \mathcal{L}}{\partial \vec{\nabla} h} = 0$$
$$\begin{array}{c} \parallel \\ 0 \end{array} - \begin{array}{c} \parallel \\ \left(\mu(\vec{x}) \frac{\partial^2 h}{\partial t^2} \right) \end{array} - \begin{array}{c} \parallel \\ \left(-\vec{\nabla} \cdot [\sigma(\vec{x}) \vec{\nabla} h] \right) \end{array} = 0$$

Thus

$$\vec{\nabla} \cdot [\sigma(\vec{x}) \vec{\nabla} h(\vec{x}, t)] = \frac{\partial^2 h(\vec{x}, t)}{\partial t^2}$$

which is a generalization of the Helmholtz equation. When μ and σ are constants, we get Helmholtz:

$$\left(\vec{\nabla}^2 - \frac{1}{c^2} \frac{\partial^2}{\partial t^2} \right) h(\vec{x}, t) = 0$$

Note $[\mu] = M L^{-2}$ and $[\sigma] = E L^{-2} = M T^{-2}$, thus with $c \equiv (\sigma/\mu)^{1/2}$ we have $[c] = L T^{-1}$ as before.

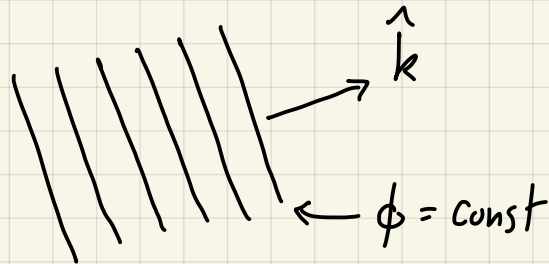
d'Alembert solution:

$$h(\vec{x}, t) = f(\hat{k} \cdot \vec{x} - ct)$$

where \hat{k} is a fixed direction in space. These are plane waves (really "line waves"). The locus of points of constant $h(\vec{x}, t)$ satisfies

$$\phi(\vec{x}, t) = \hat{k} \cdot \vec{x} - ct = \text{constant}$$

and setting $d\phi = 0$ then yields $\hat{k} \cdot \frac{d\vec{x}}{dt} = c$, i.e. the velocity along \hat{k} is c . The component of \vec{x} lying perpendicular to \hat{k} is arbitrary, so constant $\phi(\vec{x}, t)$ corresponds to lines orthogonal to \hat{k} .



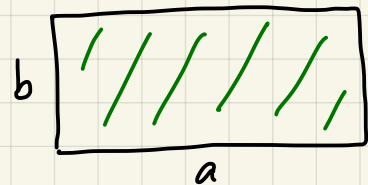
Due to linearity of the wave eqn, we can superpose plane wave solutions to arrive at the general solution,

$$h(\vec{x}, t) = \int \frac{d^2k}{(2\pi)^2} \left[A(\vec{k}) e^{i(\vec{k} \cdot \vec{x} - ckt)} + B(\vec{k}) e^{i(\vec{k} \cdot \vec{x} + ckt)} \right]$$

\uparrow \uparrow
 $+\hat{k}$ mover $k=|\vec{k}|$ $-\hat{k}$ mover

- Rectangles: $\Omega = [0, a] \times [0, b]$

Separation of variables solves PDE:



$$h(x, y, t) = X(x) Y(y) T(t)$$

Helmholtz eqn $\frac{1}{h} \left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} - \frac{1}{c^2} \frac{\partial^2}{\partial t^2} \right) h = 0$ yields

$$\frac{1}{X} \frac{\partial^2 X}{\partial x^2} + \frac{1}{Y} \frac{\partial^2 Y}{\partial y^2} = \frac{1}{c^2} \cdot \frac{1}{T} \frac{\partial^2 T}{\partial t^2}$$

depends only on x

depends only on y

depends only on t

So we conclude

$$\frac{1}{X} \frac{\partial^2 X}{\partial x^2} = -k_x^2, \quad \frac{1}{Y} \frac{\partial^2 Y}{\partial y^2} = -k_y^2, \quad \frac{1}{T} \frac{\partial^2 T}{\partial t^2} = -\omega^2$$

with

$$k_x^2 + k_y^2 = \frac{\omega^2}{c^2}$$

Thus, $\omega = c|k|$. Most general solⁿ:

$$X(x) = A \sin(k_x x + \alpha)$$

$$Y(y) = B \sin(k_y y + \beta), \quad h(x, y, t) = X(x)Y(y)T(t)$$

$$T(t) = C \sin(\omega t + \gamma)$$

but imposing boundary conditions $h(\vec{x}, t)|_{\partial\Omega} = 0$ then requires

$$\alpha = \beta = 0, \quad \sin(k_x a) = \sin(k_y b) = 0 \Rightarrow \begin{cases} k_x = m\pi/a \\ k_y = n\pi/b \end{cases}$$

The most general solⁿ consistent with the BCs is then

$$h(x, y, t) = \sum_{m=1}^{\infty} \sum_{n=1}^{\infty} A_{mn} \sin\left(\frac{m\pi x}{a}\right) \sin\left(\frac{n\pi y}{b}\right) \sin(\omega_{mn} t + \gamma_{mn})$$

where

$$\omega_{mn} = \sqrt{\left(\frac{m\pi c}{a}\right)^2 + \left(\frac{n\pi c}{b}\right)^2}$$

and the constants $\{A_{mn}, \gamma_{mn}\}$ are determined by the initial conditions.

- Circles : $\Omega = \{ (x, y) \mid x^2 + y^2 \leq a^2 \}$

It is convenient to work in 2d polar coordinates (r, φ) .

The Helmholtz equation takes the form

$$\nabla^2 h = \frac{1}{r} \frac{\partial}{\partial r} \left(r \frac{\partial h}{\partial r} \right) + \frac{1}{r^2} \frac{\partial^2 h}{\partial \varphi^2} = \frac{1}{c^2} \frac{\partial^2 h}{\partial t^2}$$

Separation of variables:

$$h(r, \varphi, t) = R(r) \Phi(\varphi) T(t)$$

Again we have

$$\frac{1}{R} \cdot \frac{1}{r} \frac{\partial}{\partial r} \left(r \frac{\partial R}{\partial r} \right) + \frac{1}{\Phi} \cdot \frac{1}{r^2} \frac{\partial^2 \Phi}{\partial \varphi^2} = \frac{1}{c^2} \frac{1}{T} \frac{\partial^2 T}{\partial t^2}$$

with

$$\Phi(\varphi) = \cos(m\varphi + \beta)$$

$$T(t) = \cos(\omega t + \gamma)$$

and

$$\frac{d^2 R}{dr^2} + \frac{1}{r} \frac{dR}{dr} + \left(\frac{m^2}{r^2} - \frac{\omega^2}{c^2} \right) R = 0$$

Since $h(r, \varphi + 2\pi, t) = h(r, \varphi, t)$, we must have $m \in \mathbb{Z}$.

This is Bessel's equation, with solutions

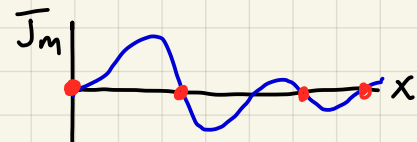
$$R(r) = A J_m \left(\frac{\omega r}{c} \right) + B N_m \left(\frac{\omega r}{c} \right)$$

with $J_m(z)$ and $N_m(z)$ the Bessel and Neumann functions

of order m , respectively. Since $N_m(z)$ diverges as $z \rightarrow 0$ for all m , we must have $B=0$. (For an annulus, we may have $B \neq 0$.) The boundary condition at $r=a$ yields

$$J_m\left(\frac{W a}{c}\right) = 0 \Rightarrow W = W_{ml} = x_{ml} \cdot \frac{c}{a}$$

where $J_m(x_{ml}) = 0$, i.e. x_{ml} is the l^{th} zero ($l=1, 2, \dots, \infty$) of $J_m(x)$. Thus,



$$h(r, \varphi, t) = \sum_{m=0}^{\infty} \sum_{l=1}^{\infty} A_{ml} J_m(x_{ml} r/a) \cos(m\varphi + \beta_{ml}) \cos(W_{ml} t + \gamma_{ml})$$

The constants A_{ml} , β_{ml} , and γ_{ml} are set by the initial conditions. Note $h(r=a, \varphi, t) = 0$ for all φ and for all t .

ω

- Read § 9.3.6 (sound in fluids) and § 9.4 (dispersion)

- Classical Field Theory

Independent variables: $\{x^1, \dots, x^n\} \in \Omega \subset \mathbb{R}^n$

Real fields: $\{\phi_1, \dots, \phi_k\}$

or $\{x^0, x^1, \dots, x^d\}$

Lagrangian density: $\mathcal{L} = \mathcal{L}(\phi_a, \partial_\mu \phi_a, x^\mu)$

$n=d+1$

Action: $S = \int d^n x \mathcal{L}$

Let's compute the variation of S :

$$\begin{aligned} \delta S &= \int_{\Omega} d^n x \left\{ \frac{\partial \mathcal{L}}{\partial \phi_a} \delta \phi_a + \frac{\partial \mathcal{L}}{\partial (\partial_{\mu} \phi_a)} \frac{\partial \delta \phi_a}{\partial x^{\mu}} \right\} \\ &= \int_{\Omega} d^n x \left\{ \frac{\partial \mathcal{L}}{\partial \phi_a} - \frac{\partial}{\partial x^{\mu}} \left(\frac{\partial \mathcal{L}}{\partial (\partial_{\mu} \phi_a)} \right) \right\} \delta \phi_a \quad \text{differential surface area} \\ &\quad + \oint_{\partial \Omega} d\Sigma n^{\mu} \frac{\partial \mathcal{L}}{\partial (\partial_{\mu} \phi_a)} \delta \phi_a \end{aligned}$$

The surface term vanishes if we demand

$$\delta \phi_a(\vec{x}) \Big|_{\partial \Omega} = 0 \quad \text{or} \quad n^{\mu} \frac{\partial \mathcal{L}}{\partial (\partial_{\mu} \phi_a)} \Big|_{\partial \Omega} = 0$$

Then we have

$$\frac{\delta S}{\delta \phi_a(\vec{x})} = \left[\frac{\partial \mathcal{L}}{\partial \phi_a} - \frac{\partial}{\partial x^{\mu}} \left(\frac{\partial \mathcal{L}}{\partial (\partial_{\mu} \phi_a)} \right) \right]_{\vec{x}} \quad \leftarrow \text{evaluate at } \vec{x}$$

Thus $\delta S = 0$ entails the Euler-Lagrange equations,

$$\frac{\partial \mathcal{L}}{\partial \phi_a} - \frac{\partial}{\partial x^{\mu}} \left(\frac{\partial \mathcal{L}}{\partial (\partial_{\mu} \phi_a)} \right) = 0$$

When \mathcal{L} is independent of the independent variables x^{μ} , the stress-energy tensor is conserved:

$$\partial_{\mu} T^{\mu}_{\nu} = 0 \quad \text{with} \quad T^{\mu}_{\nu} = \sum_a \frac{\partial \mathcal{L}}{\partial (\partial_{\mu} \phi_a)} \partial_{\nu} \phi_a - \delta^{\mu}_{\nu} \mathcal{L}$$

This is analogous to $\frac{dH}{dt} = 0$ in particle mechanics.

Maxwell theory

The Lagrangian density, with sources, is

$$\mathcal{L}(A^\nu, \partial_\mu A^\nu) = -\frac{1}{16\pi} F_{\mu\nu} F^{\mu\nu} - J_\mu A^\mu$$

where $\partial_\mu = \frac{\partial}{\partial x^\mu}$ with $x^\mu = (ct, x, y, z) = (x^0, x^1, x^2, x^3)$
and

$$F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu ; A_\nu = g_{\nu\lambda} A^\lambda , g = \text{diag}(+, -, -, -)$$

$$F^{\mu\nu} = \partial^\mu A^\nu - \partial^\nu A^\mu = g^{\mu\alpha} g^{\nu\beta} A_{\alpha\beta} ; g_{\mu\nu} = g^{\mu\nu}$$

The EL equations are

$$\frac{\partial \mathcal{L}}{\partial A_\nu} - \frac{\partial}{\partial x^\mu} \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu A_\nu)} \right) = 0 \Rightarrow \partial_\mu F^{\mu\nu} = 4\pi J^\nu$$

Conserved currents in field theory

In particle mechanics, a one-parameter family of transformations $\tilde{q}_\sigma(q, \tilde{s})$ which leaves $L(q, \dot{q}, t)$ invariant results in a conserved "charge"

$$\Lambda = \sum_\sigma \frac{\partial L}{\partial \dot{q}_\sigma} \frac{\partial \tilde{q}_\sigma}{\partial \tilde{s}} \Big|_{\tilde{s}=0} ; \tilde{q}_\sigma(q, \tilde{s}=0) = q_\sigma$$

with $d\Lambda/dt = 0$. We generalize to field theory

by taking $q_0(t) \rightarrow \phi_a(x^\mu)$. Then

$$\begin{aligned} \frac{d}{d\zeta} \left| \mathcal{L}(\tilde{\phi}_a, \partial_\mu \tilde{\phi}_a, x^\mu) \right|_{\zeta=0} &= \left. \frac{\partial \mathcal{L}}{\partial \phi_a} \frac{\partial \tilde{\phi}_a}{\partial \zeta} \right|_{\zeta=0} + \left. \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_a)} \frac{\partial}{\partial x^\mu} \frac{\partial \tilde{\phi}_a}{\partial \zeta} \right|_{\zeta=0} \\ &= \left. \frac{\partial}{\partial x^\mu} \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_a)} \frac{\partial \tilde{\phi}_a}{\partial \zeta} \right) \right|_{\zeta=0}, \end{aligned}$$

where we have invoked the EL eqns,

$$\frac{\partial \mathcal{L}}{\partial \phi_a} = \frac{\partial}{\partial x^\mu} \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_a)} \right)$$

Thus we have

$$\partial_\mu J^\mu = 0 \quad \text{with} \quad J^\mu = \sum_a \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_a)} \frac{\partial \tilde{\phi}_a}{\partial \zeta} \Big|_{\zeta=0}$$

Let us write $x^\mu = \{x^0, x^1, \dots, x^d\}$ with $n = d+1$. Then with $x^0 \equiv ct$ and $Q_\Omega \equiv c^{-1} \int d^d x J^0$, we have

$$\frac{dQ_\Omega}{dt} = \int_\Omega d^d x \partial_0 J^0 = - \int_\Omega d^3 x \vec{\nabla} \cdot \vec{J} = - \oint_{\partial\Omega} d\Sigma \hat{n} \cdot \vec{J} = 0$$

provided $\hat{n} \cdot \vec{J} \Big|_{\partial\Omega} = 0$. Thus, the rate of change of Q_Ω is minus the integrated flux exiting the region Ω .

Example:

$$\mathcal{L}(\psi, \psi^*, \partial_\mu \psi, \partial_\mu \psi^*) = \frac{1}{2} k (\partial_\mu \psi^*) (\partial^\mu \psi) - U(\psi^* \psi)$$

The Lagrangian density is invariant under

$$\psi \rightarrow \tilde{\psi} = e^{i\zeta} \psi, \quad \psi^* \rightarrow \tilde{\psi}^* = e^{-i\zeta} \psi^*$$

We regard ψ and ψ^* as independent fields. Thus,

$$\frac{\partial \tilde{\psi}}{\partial \zeta} = i e^{i\zeta} \psi, \quad \frac{\partial \tilde{\psi}^*}{\partial \zeta} = -i e^{-i\zeta} \tilde{\psi}^*$$

and thus

$$\begin{aligned} J^\mu &= \frac{\partial \mathcal{L}}{\partial (\partial_\mu \psi)} \cdot (i\psi) + \frac{\partial \mathcal{L}}{\partial (\partial_\mu \psi^*)} (-i\psi^*) \\ &= \frac{\kappa}{2i} (\psi^* \partial^\mu \psi - \psi \partial^\mu \psi^*) = \kappa \operatorname{Im}(\psi^* \partial^\mu \psi) \end{aligned}$$

Note that $U(\tilde{\psi}^* \tilde{\psi}) = U(\psi^* \psi)$ is independent of ζ .

- Gross - Pitaevskii model

This is a model of nonrelativistic interacting bosons, with

$$\mathcal{L} = i\hbar \psi^* \frac{\partial \psi}{\partial t} - \frac{\hbar^2}{2m} \vec{\nabla} \psi^* \cdot \vec{\nabla} \psi - g(\psi^* \psi - n_0)^2$$

Details in §9.5.3 of the notes. The EL equations are

$$i\hbar \frac{\partial \psi}{\partial t} = -\frac{\hbar^2}{2m} \nabla^2 \psi + 2g(|\psi|^2 - n_0) \psi$$

and its complex conjugate. This is called the nonlinear Schrödinger equation (NLSE). The one-parameter invariance of \mathcal{L} is again

$$\psi(\vec{x}, t) \rightarrow \tilde{\psi}(\vec{x}, t) \equiv e^{-i\zeta} \psi(\vec{x}, t)$$

$$\psi^*(\vec{x}, t) \rightarrow \tilde{\psi}^*(\vec{x}, t) \equiv e^{+i\zeta} \psi^*(\vec{x}, t)$$

The conserved current is

$$J^\mu = \frac{\partial \mathcal{L}}{\partial(\partial_\mu \psi)} \frac{\partial \tilde{\psi}}{\partial \zeta} \Big|_{\zeta=0} + \frac{\partial \mathcal{L}}{\partial(\partial_\mu \psi^*)} \frac{\partial \tilde{\psi}^*}{\partial \zeta} \Big|_{\zeta=0}$$

with components

$$J^0 = \hbar |\psi|^2 \equiv \hbar \rho$$

$$\vec{j} = \frac{\hbar^2}{2im} (\psi^* \vec{\nabla} \psi - \psi \vec{\nabla} \psi^*) \equiv \hbar \vec{j}$$

Thus,

$$\frac{\partial \rho}{\partial t} + \vec{\nabla} \cdot \vec{j} = 0 \quad (\text{continuity eqn.})$$

In this example, $x^\mu = x_\mu$ and there is no difference between raised and lowered indices.