

Mechanics

Lecture 2 (Oct 7)

Hamilton's principle: $\delta S = 0$ where

$$S[q(t)] = \int_{t_1}^{t_2} dt L(q, \dot{q}, t) = \text{action functional}$$

with $q = \{q_1, \dots, q_n\}$ = set of generalized coordinates

The function $L(q, \dot{q}, t)$ is the Lagrangian, and is given by $L = T - U$, where T = kinetic energy and U = potential energy. Typically $T = T(q, \dot{q})$ is a quadratic form in the generalized velocities $\{\dot{q}_\sigma\}$, i.e. $T(q, \dot{q}) = T_{\sigma\sigma'}(q) \dot{q}_\sigma \dot{q}_{\sigma'}$. For example

$$T = \frac{1}{2} m \dot{\vec{x}}^2 = \frac{1}{2} m (\dot{x}^2 + \dot{y}^2 + \dot{z}^2) \quad \text{Cartesian } (x, y, z)$$
$$= \frac{1}{2} m (\dot{r}^2 + r^2 \dot{\theta}^2 + r^2 \sin^2 \theta \dot{\phi}^2) \quad \text{polar } (r, \theta, \phi)$$

The potential energy U is most often a function of q , but $U = U(q, \dot{q})$ applies, e.g., for charged particles in a magnetic field, where

$$U(\vec{x}, \dot{\vec{x}}) = q \phi(\vec{x}) - \frac{q}{c} \vec{A}(\vec{x}) \cdot \frac{d\vec{x}}{dt}$$

← charge ← scalar potential ← vector potential

Free particle $\Rightarrow L = \frac{1}{2} m \dot{\vec{v}}^2$ (§ 3.6.3)

• NB: In general $L = \frac{1}{2} T_{\sigma\sigma'}(q, t) \dot{q}_\sigma \dot{q}_{\sigma'} - U(q, \dot{q}, t)$

Equations of motion: $F_\sigma \equiv \frac{\partial L}{\partial q_\sigma} = \text{generalized force}$

$$\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}_\sigma} \right) = \frac{\partial L}{\partial q_\sigma}, \quad \sigma \in \{1, \dots, n\}$$

$y \rightarrow q_\sigma$
 $y' \rightarrow \dot{q}_\sigma$

$$P_\sigma \equiv \frac{\partial L}{\partial \dot{q}_\sigma} = \text{generalized momentum}$$

Thus, $\dot{P}_\sigma = F_\sigma$, i.e. Newton's second law.

• Conservation laws:

Most general setting: to be discussed (Noether's theorem)

For now, recall results from COV:

① $\frac{\partial L}{\partial q_\sigma} = 0 \Rightarrow P_\sigma = \frac{\partial L}{\partial \dot{q}_\sigma} = \text{constant} \quad (\dot{P}_\sigma = 0)$

Momentum P_σ is conserved because the force $F_\sigma = 0$

Example: $T = \frac{1}{2} m (\dot{x}^2 + \dot{y}^2 + \dot{z}^2)$, $U = mgz$

$$\Rightarrow F_x = \frac{\partial L}{\partial x} = 0, \quad F_y = \frac{\partial L}{\partial y} = 0$$

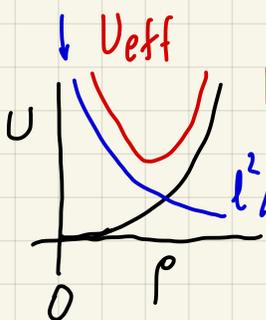
$$P_x = m\dot{x} \Rightarrow x(t) = x(0) + \frac{P_x}{m} t$$

$$P_y = m\dot{y} \Rightarrow y(t) = y(0) + \frac{P_y}{m} t$$

$$P_z = m\dot{z}, \quad F_z = -\frac{\partial U}{\partial z} = -mg$$

$$m\ddot{z} = -mg \Rightarrow z(t) = z(0) + \dot{z}(0)t - \frac{1}{2}gt^2$$

angular momentum barrier



Example': $L = \frac{1}{2} m (\dot{\rho}^2 + \rho^2 \dot{\phi}^2) - U(\rho)$ (2D polar)

$$P_\phi = \frac{\partial L}{\partial \dot{\phi}} = m\rho^2 \dot{\phi} = \text{angular momentum} \equiv l$$

ρ eqn: $m\ddot{\rho} = m\rho \dot{\phi}^2 - U'(\rho) = \frac{l^2}{m\rho^3} - U'(\rho) = -U'_{\text{eff}}(\rho)$

$$U_{\text{eff}}(\rho) = \frac{l^2}{2m\rho^2} + U(\rho)$$

Scratch

$$L = \frac{1}{2}m(\dot{\rho}^2 + \rho^2\dot{\phi}^2) - U(\rho)$$

$$p_{\phi} = \frac{\partial L}{\partial \dot{\phi}} = m\rho^2\dot{\phi} = l$$

IMPORTANT: Can substitute $\dot{\phi} = \frac{l}{m\rho^2}$ in eqns of motion but not in Lagrangian itself!

WRONG:

$$\begin{aligned} L &= \frac{1}{2}m\dot{\rho}^2 + \frac{1}{2}m\rho^2\dot{\phi}^2 - U(\rho) \\ &= \frac{1}{2}m\dot{\rho}^2 + \frac{1}{2}m\rho^2\left(\frac{l}{m\rho^2}\right)^2 - U(\rho) \\ &= \frac{1}{2}m\dot{\rho}^2 + \frac{l^2}{2m\rho^2} - U(\rho) \end{aligned}$$

$$\frac{d}{dt}\left(\frac{\partial L}{\partial \dot{\rho}}\right) = m\ddot{\rho} = \frac{\partial L}{\partial \rho} = -\frac{l^2}{m\rho^3} - U'(\rho)$$

wrong sign!

RIGHT:

$$L = \frac{1}{2}m\dot{\rho}^2 + \frac{1}{2}m\rho^2\dot{\phi}^2 - U(\rho)$$
$$p_{\phi} = \frac{\partial L}{\partial \dot{\phi}} = m\rho^2\dot{\phi} = l \text{ constant } (\dot{p}_{\phi} = 0)$$

$$\begin{aligned} \frac{d}{dt}\left(\frac{\partial L}{\partial \dot{\rho}}\right) &= m\ddot{\rho} = \frac{\partial L}{\partial \rho} = m\rho\dot{\phi}^2 - U'(\rho) \\ &= m\rho\left(\frac{l}{m\rho^2}\right)^2 - U'(\rho) \\ &= +\frac{l^2}{m\rho^3} - U'(\rho) = -U'_{\text{eff}}(\rho) \end{aligned}$$

right sign!

$$\textcircled{2} \quad \frac{\partial L}{\partial t} = 0 \Rightarrow H = \dot{q}_\sigma \frac{\partial L}{\partial \dot{q}_\sigma} - L = \dot{q}_\sigma p_\sigma - L \quad \text{conserved}$$

← implied summation on repeated indices

See it again:

$$\frac{dH}{dt} = \cancel{\ddot{q}_\sigma p_\sigma} + \cancel{\dot{q}_\sigma \dot{p}_\sigma} - \underbrace{\frac{\partial L}{\partial \dot{q}_\sigma}}_{F_\sigma} \dot{q}_\sigma - \cancel{\frac{\partial L}{\partial \dot{q}_\sigma} \ddot{q}_\sigma} - \frac{\partial L}{\partial t}$$

Thus, $\frac{dH}{dt} = -\frac{\partial L}{\partial t}$, and for

$$\frac{dL}{dt} = \frac{\partial L}{\partial q_\sigma} \frac{dq_\sigma}{dt} + \frac{\partial L}{\partial \dot{q}_\sigma} \frac{d\dot{q}_\sigma}{dt} + \frac{\partial L}{\partial t}$$

$$L = \sum_{j=1}^N \frac{1}{2} m_j \dot{\vec{x}}_j^2 - U(\vec{x}_1, \dots, \vec{x}_N)$$

$$\text{we have that } H = \sum_{j=1}^N \frac{1}{2} m_j \dot{\vec{x}}_j^2 + U(\vec{x}_1, \dots, \vec{x}_N)$$

is a constant of the motion.

• In general, $H = \dot{q}_\sigma p_\sigma - L(q, \dot{q}, t)$ is a Legendre transform of L :

$$dH = \cancel{p_\sigma d\dot{q}_\sigma} + \dot{q}_\sigma dp_\sigma - \frac{\partial L}{\partial q_\sigma} dq_\sigma - \cancel{\frac{\partial L}{\partial \dot{q}_\sigma} d\dot{q}_\sigma} - \frac{\partial L}{\partial t} dt$$

and hence $H = H(q, p, t)$ with

$$\frac{\partial H}{\partial q_\sigma} = -\frac{\partial L}{\partial q_\sigma} = -F_\sigma, \quad \frac{\partial H}{\partial p_\sigma} = \dot{q}_\sigma, \quad \frac{dH}{dt} = \frac{\partial H}{\partial t} = -\frac{\partial L}{\partial t}$$

We then have Hamilton's equations of motion:

$$\dot{q}_\sigma = \frac{\partial H}{\partial p_\sigma}, \quad \dot{p}_\sigma = -\frac{\partial H}{\partial q_\sigma} \Rightarrow \dot{\xi}_\alpha = J_{\alpha\beta} \frac{\partial H}{\partial \xi_\beta}$$

$$J = \begin{pmatrix} 0 & \mathbb{1} \\ -\mathbb{1} & 0 \end{pmatrix}, \quad \xi = \begin{pmatrix} q \\ p \end{pmatrix}$$

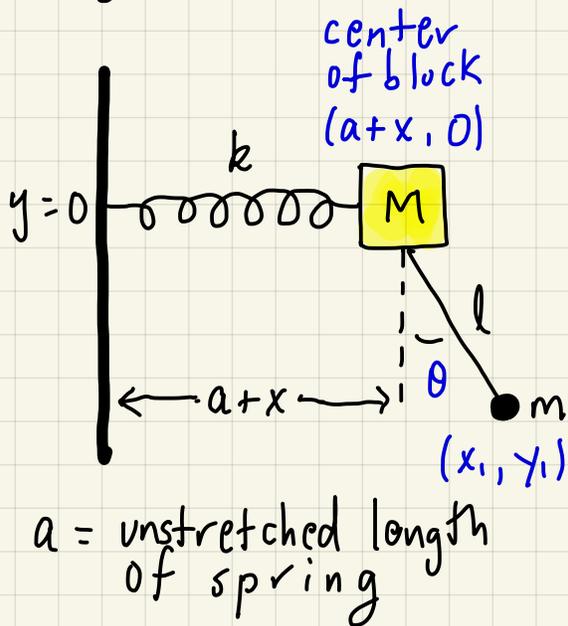
Procedure

- (i) Choose a set of generalized coordinates
- (ii) Find KE $T(q, \dot{q}, t)$ and PE $U(q, t)$ or $U(q, \dot{q}, t)$ and thus the Lagrangian $L(q, \dot{q}, t) = T - U$.
- (iii) Find the canonical momenta $p_\sigma = \frac{\partial L}{\partial \dot{q}_\sigma}$ and the generalized forces $F_\sigma = \frac{\partial L}{\partial q_\sigma}$.
 \searrow $p_\sigma = p_\sigma(q, \dot{q}, t)$
- (iv) Identify any conserved quantities (later: Noether's thm)
- (v) Evaluate \dot{p}_σ (carefully!) and write $\dot{p}_\sigma = F_\sigma$
- (vi) Integrate the equations of motion to get $\{q_\sigma(t)\}$, the motion of the system. $2n$ constants of integration $\{q_\sigma(0), \dot{q}_\sigma(0)\}$

§ 3.8 : Cartesian, cylindrical, and polar coordinates

§ 3.10 : Examples

§ 3.10.4 : Pendulum attached to mass on a spring



coordinates of mass m : (x_1, y_1)

$$x_1 = a + x + l \sin \theta, \quad y_1 = -l \cos \theta$$

$$\begin{aligned} T &= \frac{1}{2} M \dot{x}^2 + \frac{1}{2} m (\dot{x}_1^2 + \dot{y}_1^2) \\ &= \frac{1}{2} (M+m) \dot{x}^2 + \frac{1}{2} m l^2 \dot{\theta}^2 + m l \cos \theta \dot{x} \dot{\theta} \end{aligned}$$

$$\begin{aligned} U &= \frac{1}{2} k x^2 + m g y_1 \\ &= \frac{1}{2} k x^2 - m g l \cos \theta \end{aligned}$$

Lagrangian:

$$L = T - U$$

$$= \frac{1}{2}(M+m)\dot{x}^2 + \frac{1}{2}ml^2\dot{\theta}^2 + ml\cos\theta\dot{x}\dot{\theta} - \frac{1}{2}kx^2 + mgl\cos\theta$$

Generalized momenta:

$$\bullet p_x = \frac{\partial L}{\partial \dot{x}} = (M+m)\dot{x} + ml\cos\theta\dot{\theta}$$

$$\bullet p_\theta = \frac{\partial L}{\partial \dot{\theta}} = ml\cos\theta\dot{x} + ml^2\dot{\theta}$$

Generalized forces:

$$\bullet F_x = \frac{\partial L}{\partial x} = -kx$$

$$\bullet F_\theta = \frac{\partial L}{\partial \theta} = -ml\sin\theta\dot{x}\dot{\theta} - mgl\sin\theta$$

Equations of motion:

$$\bullet \dot{p}_x = F_x \Rightarrow (M+m)\ddot{x} + ml\cos\theta\ddot{\theta} - ml\sin\theta\dot{\theta}^2 = -kx$$

$$\bullet \dot{p}_\theta = F_\theta \Rightarrow ml\cos\theta\ddot{x} + ml^2\ddot{\theta} - \cancel{ml\sin\theta\dot{x}\dot{\theta}} = -\cancel{ml\sin\theta\dot{x}\dot{\theta}} - mgl\sin\theta$$

Conserved quantities:

$$\text{Only } H = \dot{x}p_x + \dot{\theta}p_\theta - L$$

$$\begin{aligned} &= \left[(M+m)\dot{x}^2 + ml\cos\theta\dot{x}\dot{\theta} \right] + \left[ml\cos\theta\dot{x}\dot{\theta} + ml^2\dot{\theta}^2 \right] \\ &\quad - \frac{1}{2}(M+m)\dot{x}^2 - \frac{1}{2}ml^2\dot{\theta}^2 - ml\cos\theta\dot{x}\dot{\theta} + \frac{1}{2}kx^2 - mgl\cos\theta \\ &= \frac{1}{2}(M+m)\dot{x}^2 + \frac{1}{2}ml^2\dot{\theta}^2 + ml\cos\theta\dot{x}\dot{\theta} + \frac{1}{2}kx^2 - mgl\cos\theta \\ &= T + U \equiv E \end{aligned}$$

Small oscillations: linearize the equations of motion

$$\bullet (M+m)\ddot{x} + m\ell\cos\theta\ddot{\theta} - m\ell\sin\theta\dot{\theta}^2 = -kx$$

$$\bullet m\ell\cos\theta\ddot{x} + m\ell^2\ddot{\theta} = -mgl\sin\theta$$

$$\Rightarrow \begin{cases} \bullet (M+m)\ddot{x} + m\ell\ddot{\theta} = -kx \\ \bullet \ddot{x} + \ell\ddot{\theta} = -g\theta \end{cases} \quad \left(\begin{array}{l} \text{expand about } x=\theta=0 \\ \text{assume } x, \theta, \dot{x}, \dot{\theta} \text{ small} \end{array} \right)$$

The five parameters (M, m, ℓ, k, g) may be reduced to three:

$$u \equiv \frac{x}{\ell}, \quad \alpha \equiv \frac{m}{M}, \quad \omega_0^2 \equiv \frac{k}{M}, \quad \omega_1^2 \equiv \frac{g}{\ell}$$

Then we have

$$\bullet (1+\alpha)\ddot{u} + \alpha\ddot{\theta} + \omega_0^2 u = 0$$

$$\bullet \ddot{u} + \ddot{\theta} + \omega_1^2 \theta = 0$$

This linear system may be solved by writing

$$\begin{pmatrix} u(t) \\ \theta(t) \end{pmatrix} = \begin{pmatrix} u_0 \\ \theta_0 \end{pmatrix} e^{-i\omega t} \quad \frac{d^2}{dt^2} \rightarrow -\omega^2$$

$$\Rightarrow \begin{pmatrix} \omega_0^2 - (1+\alpha)\omega^2 & -\alpha\omega^2 \\ -\omega^2 & \omega_1^2 - \omega^2 \end{pmatrix} \begin{pmatrix} u_0 \\ \theta_0 \end{pmatrix} = 0$$

A nontrivial solⁿ requires that the determinant vanish:

$$\omega^4 - [\omega_0^2 + (1+\alpha)\omega_1^2]\omega^2 + \omega_0^2\omega_1^2 = 0$$

$$\omega_{\pm}^2 = \frac{1}{2}[\omega_0^2 + (1+\alpha)\omega_1^2] \pm \frac{1}{2}\sqrt{[\omega_0^2 - (1+\alpha)\omega_1^2]^2 + 4\alpha\omega_0^2\omega_1^2}$$

There are two eigenvalues for ω^2 , given by

$$\omega_{\pm}^2 = \frac{1}{2}[\omega_0^2 + (1+\alpha)\omega_1^2] \pm \frac{1}{2}\sqrt{[\omega_0^2 - (1+\alpha)\omega_1^2]^2 + 4\alpha\omega_0^2\omega_1^2}$$

The general solⁿ is then

$$\begin{pmatrix} u(t) \\ \theta(t) \end{pmatrix} = \text{Re} \left[\begin{pmatrix} u_0^+ \\ \theta_0^+ \end{pmatrix} e^{-i\omega_+ t} + \begin{pmatrix} u_0^- \\ \theta_0^- \end{pmatrix} e^{-i\omega_- t} \right]$$

where

↑ solution must be real

normal modes

$$\begin{pmatrix} \omega_0^2 - (1+\alpha)\omega_{\pm}^2 & -\alpha\omega_{\pm}^2 \\ -\omega_{\pm}^2 & \omega_1^2 - \omega_{\pm}^2 \end{pmatrix} \begin{pmatrix} u_0^{\pm} \\ \theta_0^{\pm} \end{pmatrix} = 0$$

$$\begin{pmatrix} u_0^{\pm} \\ \theta_0^{\pm} \end{pmatrix} e^{-i\omega_{\pm} t}$$

This fixes the ratios $\frac{u_0^{\pm}}{\theta_0^{\pm}} = \left(\frac{\omega_1^2}{\omega_{\pm}^2} - 1 \right) \in \mathbb{R}$

$$\begin{pmatrix} u_1 \\ u_2 \\ \vdots \\ u_n \end{pmatrix} \quad \begin{matrix} u_2/u_1 \\ \vdots \\ u_n/u_1 \end{matrix}$$

Thus, we are free to choose θ_0^{\pm} , which are two complex constants \Rightarrow four real parameters.

We fix them via the initial conditions,

$$\begin{pmatrix} u(0) \\ \theta(0) \end{pmatrix} \quad \text{and} \quad \begin{pmatrix} \dot{u}(0) \\ \dot{\theta}(0) \end{pmatrix} \Rightarrow \text{four real pieces of initial data}$$

Here we have used the fact that if $\begin{pmatrix} u_0 \\ \theta_0 \end{pmatrix} e^{-i\omega t}$ is a solⁿ, then so is $\begin{pmatrix} u_0 \\ \theta_0 \end{pmatrix} e^{+i\omega t}$. In this sense, we might speak of four eigenfrequencies $\{\omega_+, \omega_-, -\omega_+, -\omega_-\}$ of which two are positive and two are negative.

Scratch

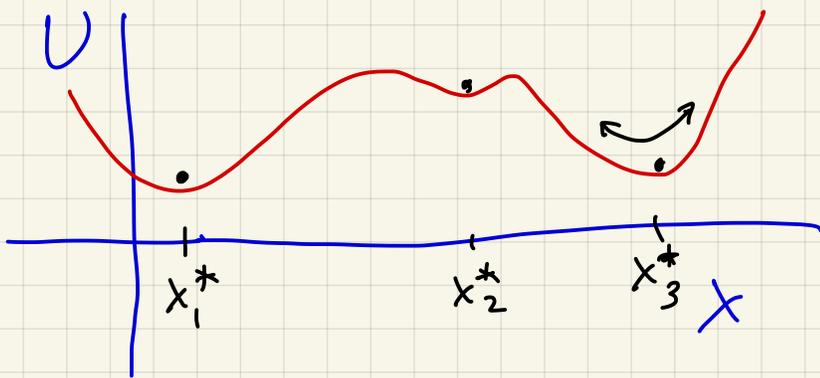
$$U(x) = \frac{1}{2} k x^2 + \frac{c}{4} b x^4$$

$$T(\dot{x}) = \frac{1}{2} m \dot{x}^2$$

$$L = \frac{1}{2} m \dot{x}^2 - \frac{1}{2} k x^2 - \frac{1}{4} b x^4$$

$$m \ddot{x} = -kx - \cancel{bx^3}$$

Eqbm @ $x=0, \dot{x}=0$



expand about $x = x_j^*$ (sol^{ns} to $U'(x^*) = 0$)

$$\Rightarrow x = x_j^* + \delta x$$

$$\Rightarrow m \delta \ddot{x} = -U''(x_j^*) \delta x$$

$$\omega_j^2 = \sqrt{U''(x_j^*)/m}$$



U'' small
 ω_j small



U'' big
 ω big

Virial Theorem

- formula describing time-averaged motion of a mechanical system

Define the **virial** $G(q,p) = \sum_{\sigma} q_{\sigma} p_{\sigma}$, for which

$$\frac{dG}{dt} = \sum_{\sigma} (\dot{q}_{\sigma} p_{\sigma} + \dot{p}_{\sigma} q_{\sigma}) = \sum_{\sigma} \left\{ \dot{q}_{\sigma} \frac{\partial L}{\partial \dot{q}_{\sigma}} + q_{\sigma} \frac{\partial L}{\partial q_{\sigma}} \right\}$$

Suppose $T = \frac{1}{2} T_{00}$, $(q) \dot{q}_{\sigma} \dot{q}_{\sigma}$ is homogeneous of degree $k=2$ in the generalized velocities, and that $\partial U / \partial \dot{q}_{\sigma} = 0$. Then

$$\sum_{\sigma} \dot{q}_{\sigma} \frac{\partial L}{\partial \dot{q}_{\sigma}} = \sum_{\sigma} \dot{q}_{\sigma} \frac{\partial T}{\partial \dot{q}_{\sigma}} = 2T$$

Now consider the time average of \dot{G} over $[0, \tau]$:

$$\left\langle \frac{dG}{dt} \right\rangle_{\tau} = \frac{1}{\tau} \int_0^{\tau} dt \frac{dG}{dt} = \frac{G(\tau) - G(0)}{\tau}$$

If G is bounded, then we have $\langle \dot{G} \rangle_{\tau} \rightarrow 0$ as $\tau \rightarrow \infty$.

This is the case for any bounded motion, such as planetary orbits. In such cases,

$$2\langle T \rangle = - \left\langle \sum_{\sigma=1}^n q_{\sigma} F_{\sigma} \right\rangle$$

\dim^n of space

$$n = d \cdot N$$

$$= \left\langle \sum_{j=1}^N \vec{x}_j \cdot \frac{\partial}{\partial \vec{x}_j} U(\vec{x}_1, \dots, \vec{x}_N) \right\rangle = k \langle U \rangle$$

if $U(\vec{x}_1, \dots, \vec{x}_N)$ homogeneous of degree k in $\{x_j^{\alpha}\}$.

Scratch

Euler's thm for homogeneous functions:

$f(x_1, \dots, x_n)$ homogeneous of degree k if

$$f(\lambda x_1, \dots, \lambda x_n) = \lambda^k f(x_1, \dots, x_n)$$

examples

$$f(x, y) = x^5 + ax^4y + b \frac{y^6}{x} \quad k=5$$

$$T(\dot{q}_1, \dots, \dot{q}_n) = \frac{1}{2} T_{\sigma\sigma'}(q) \dot{q}_\sigma \dot{q}_{\sigma'} \quad k=2$$

$$\begin{aligned} \left. \frac{\partial}{\partial \lambda} \right|_{\lambda=1} f(\lambda x_1, \dots, \lambda x_n) &= x_1 \frac{\partial f}{\partial x_1} + \dots + x_n \frac{\partial f}{\partial x_n} \\ &= \left. \frac{\partial}{\partial \lambda} \right|_{\lambda=1} \lambda^k f(x_1, \dots, x_n) \\ &= k \lambda^{k-1} f(x_1, \dots, x_n) \Big|_{\lambda=1} \end{aligned}$$

$$\therefore \sum_{j=1}^n x_j \frac{\partial f}{\partial x_j} = k f$$

$$\text{Check: } \left(x \frac{\partial}{\partial x} + y \frac{\partial}{\partial y} \right) \left(x^5 + ax^4y + b \frac{y^6}{x} \right) =$$

$$\begin{aligned} &x \cdot 5x^4 + x \cdot 4ax^3y - x \cdot b \frac{y^6}{x^2} + y \cdot 0 + y \cdot ax^4 + y \cdot 6b \frac{y^5}{x} \\ &= 5x^5 + 5ax^4y + 5b \frac{y^6}{x} = 5 \left(x^5 + ax^4y + b \frac{y^6}{x} \right) \end{aligned}$$

Since $T+U = E$ is conserved, we have

$$\langle T \rangle = \frac{kE}{k+2}, \quad \langle U \rangle = \frac{2E}{k+2}$$

Application: Keplerian orbits, $k = -1$

$$\langle T \rangle = -E, \quad \langle U \rangle = 2E; \quad E < 0$$

Note then that a satellite losing energy due to frictional losses as it enters the atmosphere must increase its kinetic energy, i.e. it moves faster! (Think also about angular momentum conservation.)

Noether's Theorem

Lecture 3 (Oct. 12)

"To each independent, continuous one-parameter family of coordinate transformations which leave L invariant there corresponds an associated conserved charge."

(In fact, we only need require S is invariant. See § 3.14 of the notes.)

Proof: Let $q_\sigma \rightarrow \bar{q}_\sigma(q, \zeta)$ be our one-parameter family of transformations with continuous parameter ζ , and with $\bar{q}_\sigma(q, \zeta=0) = q_\sigma \forall \sigma$. Invariance of $L \Rightarrow$

$$\begin{aligned} \frac{d}{d\zeta} \left. L(\bar{q}, \dot{\bar{q}}, t) \right|_{\zeta=0} &= \left. \frac{\partial L}{\partial \bar{q}_\sigma} \frac{\partial \bar{q}_\sigma}{\partial \zeta} \right|_{\zeta=0} + \left. \frac{\partial L}{\partial \dot{\bar{q}}_\sigma} \frac{\partial \dot{\bar{q}}_\sigma}{\partial \zeta} \right|_{\zeta=0} \quad \wedge \text{ (conserved charge)} \\ &= \left. \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}_\sigma} \right) \frac{\partial \bar{q}_\sigma}{\partial \zeta} \right|_{\zeta=0} + \left. \frac{\partial L}{\partial \dot{q}_\sigma} \frac{d}{dt} \left(\frac{\partial \bar{q}_\sigma}{\partial \zeta} \right) \right|_{\zeta=0} = \frac{d}{dt} \left. \left(\frac{\partial L}{\partial \dot{q}_\sigma} \frac{\partial \bar{q}_\sigma}{\partial \zeta} \right) \right|_{\zeta=0} = 0 \end{aligned}$$

← evaluate along motion of system