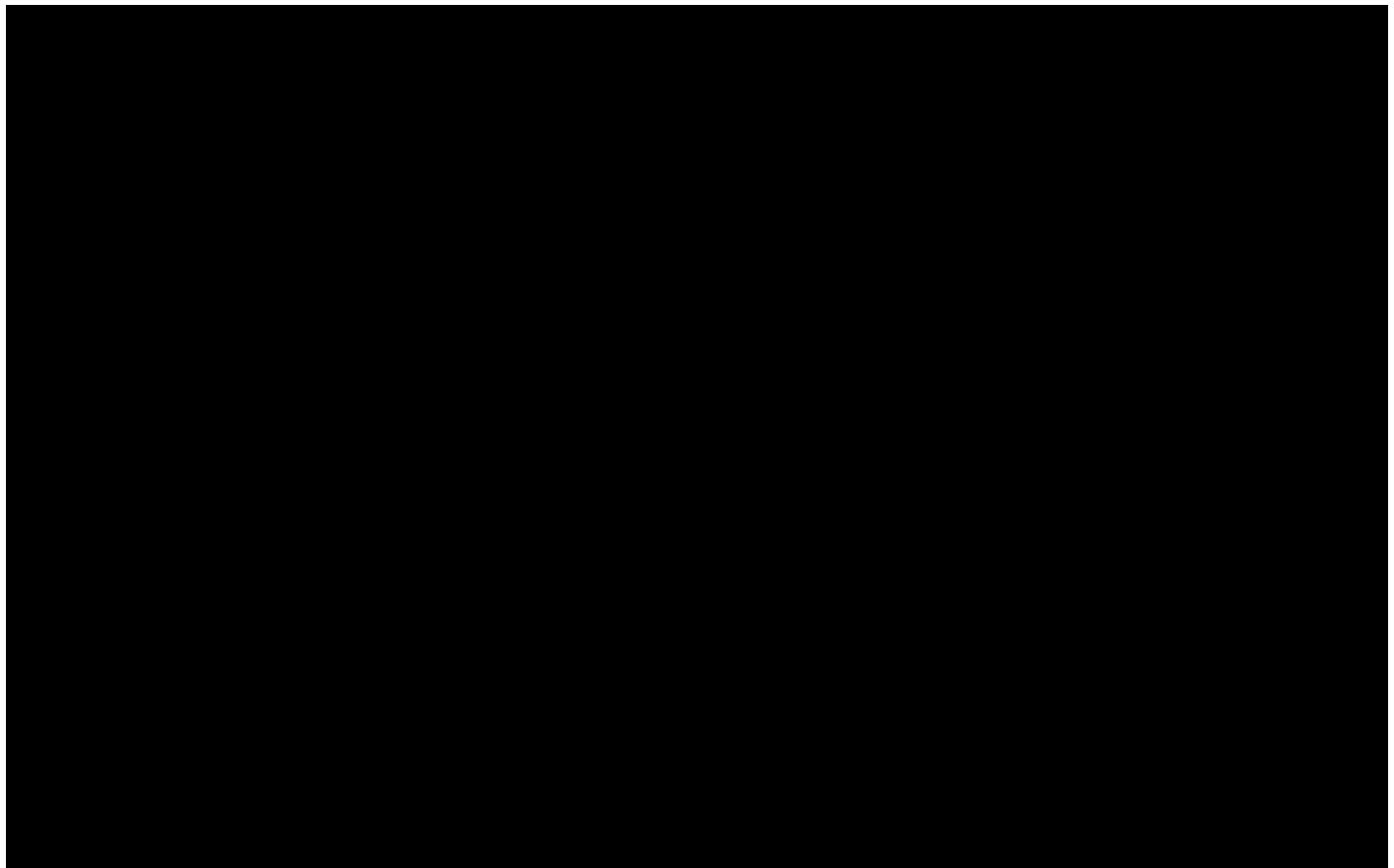


4

The Particle Nature of Matter



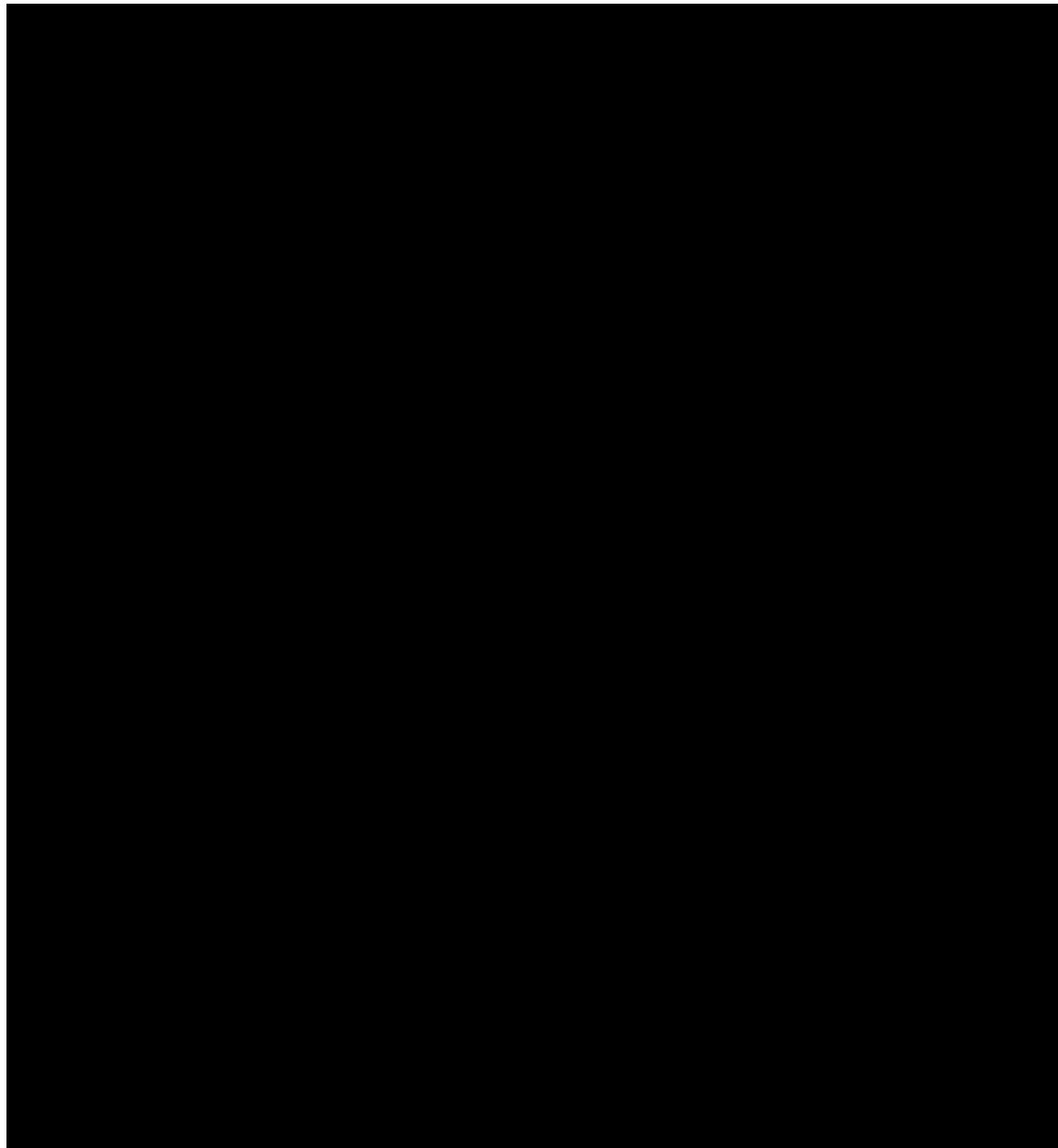
4-3 Thomson's device will work for positive and negative particles, so we may apply $\frac{q}{m} \approx \frac{V\theta}{B^2 ld}$.

(a)
$$\frac{q}{m} \approx \frac{V\theta}{B^2 ld} = (2\ 000 \text{ V}) \frac{0.20 \text{ radians}}{(4.57 \times 10^{-2} \text{ T})^2} (0.10 \text{ m})(0.02 \text{ m}) = 9.58 \times 10^7 \text{ C/kg}$$

- (b) As the particle is attracted by the negative plate, it carries a positive charge and is a proton.
$$\left(\frac{q}{m_p} = \frac{1.60 \times 10^{-19} \text{ C}}{1.67 \times 10^{-27} \text{ kg}} = 9.58 \times 10^7 \text{ C/kg} \right)$$

$$(c) \quad v_x = \frac{E}{B} = \frac{V}{dB} = \frac{2\ 000 \text{ V}}{0.02 \text{ m}} (4.57 \times 10^{-2} \text{ T}) = 2.19 \times 10^6 \text{ m/s}$$

(d) As $v_x \sim 0.01c$ there is no need for relativistic mechanics.



- 4-8 (a) From Equation 4.16 we have $\Delta n \propto \left(\frac{\sin \phi}{2}\right)^{-4}$ or $\Delta n_2 = \Delta n_1 \frac{\left(\frac{\sin \phi_1}{2}\right)^4}{\left(\frac{\sin \phi_2}{2}\right)^4}$. Thus the number of α 's scattered at 40 degrees is given by

$$\Delta n_2 = (100 \text{ cpm}) \frac{\left(\frac{\sin 20}{2}\right)^4}{\left(\frac{\sin 40}{2}\right)^4} = (100 \text{ cpm}) \left(\frac{\sin 10}{\sin 20}\right)^4 = 6.64 \text{ cpm}.$$

Similarly

$$\Delta n \text{ at 60 degrees} = 1.45 \text{ cpm}$$

$$\Delta n \text{ at 80 degrees} = 0.533 \text{ cpm}$$

$$\Delta n \text{ at 100 degrees} = 0.264 \text{ cpm}$$

- (b) From 4.16 doubling $\left(\frac{1}{2}\right)m_\alpha v_\alpha^2$ reduces Δn by a factor of 4. Thus Δn at 20 degrees $= \left(\frac{1}{4}\right)(100 \text{ cpm}) = 25 \text{ cpm}$.

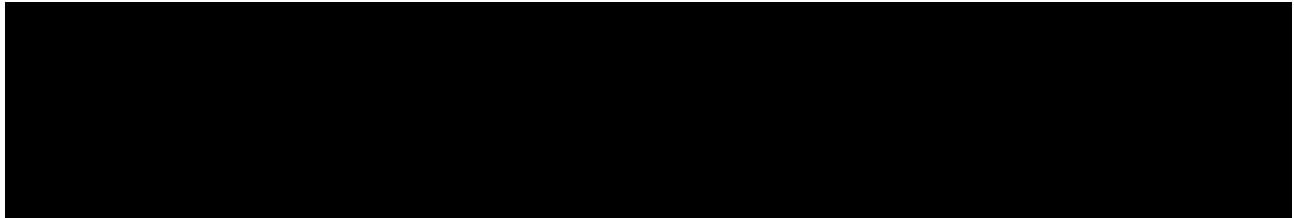
- (c) From 4.16 we find $\frac{\Delta n_{\text{Cu}}}{\Delta n_{\text{Au}}} = \frac{Z_{\text{Cu}}^2 N_{\text{Cu}}}{Z_{\text{Au}}^2 N_{\text{Au}}}$, $Z_{\text{Cu}} = 29$, $Z_{\text{Au}} = 79$.

$$\begin{aligned} N_{\text{Cu}} &= \text{number of Cu nuclei per unit area} \\ &= \text{number of Cu nuclei per unit volume} * \text{foil thickness} \\ &= \left[(8.9 \text{ g/cm}^3) \left(\frac{6.02 \times 10^{23} \text{ nuclei}}{63.54 \text{ g}} \right) \right] t = 8.43 \times 10^{22} t \\ N_{\text{Au}} &= \left[(19.3 \text{ g/cm}^3) \left(\frac{6.02 \times 10^{23} \text{ nuclei}}{197.0 \text{ g}} \right) \right] t = 5.90 \times 10^{22} t \end{aligned}$$

$$\text{So } \Delta n_{\text{Cu}} = \Delta n_{\text{Au}} (29)^2 \frac{8.43 \times 10^{22}}{(79)^2} (5.90 \times 10^2) = (100) \left(\frac{29}{79} \right)^2 \left(\frac{8.43}{5.90} \right) = 19.3 \text{ cpm}.$$

- 4-9 The initial energy of the system of α plus copper nucleus is 13.9 MeV and is just the kinetic energy of the α when the α is far from the nucleus. The final energy of the system may be evaluated at the point of closest approach when the kinetic energy is zero and the potential energy is $k(2e)\frac{Ze}{r}$ where r is approximately equal to the nuclear radius of copper. Invoking conservation of energy $E_i = E_f$, $K_\alpha = (k)\frac{2Ze^2}{r}$ or

$$r = (k) \frac{2Ze^2}{K_\alpha} = \frac{(2)(29)(1.60 \times 10^{-19})^2 (8.99 \times 10^9)}{(13.9 \times 10^6 \text{ eV})(1.60 \times 10^{-19} \text{ J/eV})} = 6.00 \times 10^{-15} \text{ m}.$$



- 4-11 $\frac{1}{\lambda} = R \left(\frac{1}{n_f^2} - \frac{1}{n_i^2} \right)$. For the Balmer series, $n_f = 2$; $n_i = 3, 4, 5, \dots$. The first three lines in the series have wavelengths given by $\frac{1}{\lambda} = R \left(\frac{1}{2^2} - \frac{1}{n^2} \right)$ where $R = 1.09737 \times 10^7 \text{ m}^{-1}$.

$$\text{1st line: } \frac{1}{\lambda} = R \left(\frac{1}{4} - \frac{1}{9} \right) = \left(\frac{5}{36} \right) R; \quad \lambda = \frac{36}{5R} = 656.112 \text{ nm}$$

$$\text{2nd line: } \frac{1}{\lambda} = R \left(\frac{1}{4} - \frac{1}{16} \right) = \left(\frac{3}{16} \right) R; \quad \lambda = \frac{16}{3R} = 486.009 \text{ nm}$$

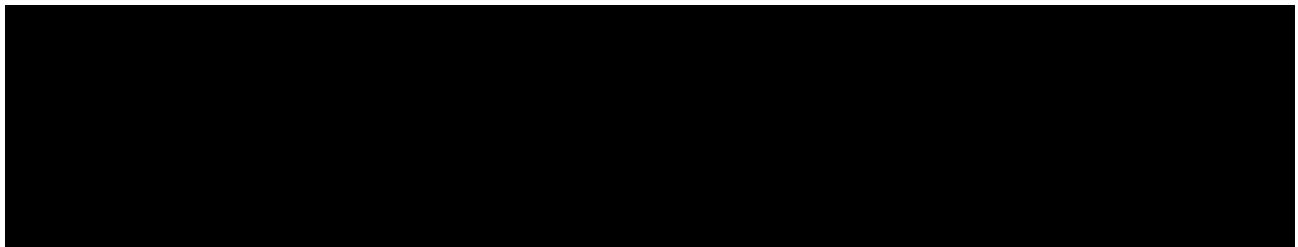
$$\text{3rd line: } \frac{1}{\lambda} = R \left(\frac{1}{4} - \frac{1}{25} \right) = \left(\frac{21}{100} \right) R; \quad \lambda = \frac{100}{21R} = 433.937 \text{ nm}$$

- 4-12 $\frac{1}{\lambda} = R \left(\frac{1-1}{n^2} \right)$ where $n = 2, 3, 4, \dots$ and $R = 1.0973732 \times 10^7 \text{ m}^{-1}$;

$$\text{For } n = 2: \lambda = R^{-1} \left(1 - \frac{1}{2^2} \right)^{-1} = 1.21502 \times 10^{-7} \text{ m} = 121.502 \text{ nm (UV)}$$

$$\text{For } n = 3: \lambda = R^{-1} \left(1 - \frac{1}{3^2} \right)^{-1} = 1.02517 \times 10^{-7} \text{ m} = 102.517 \text{ nm (UV)}$$

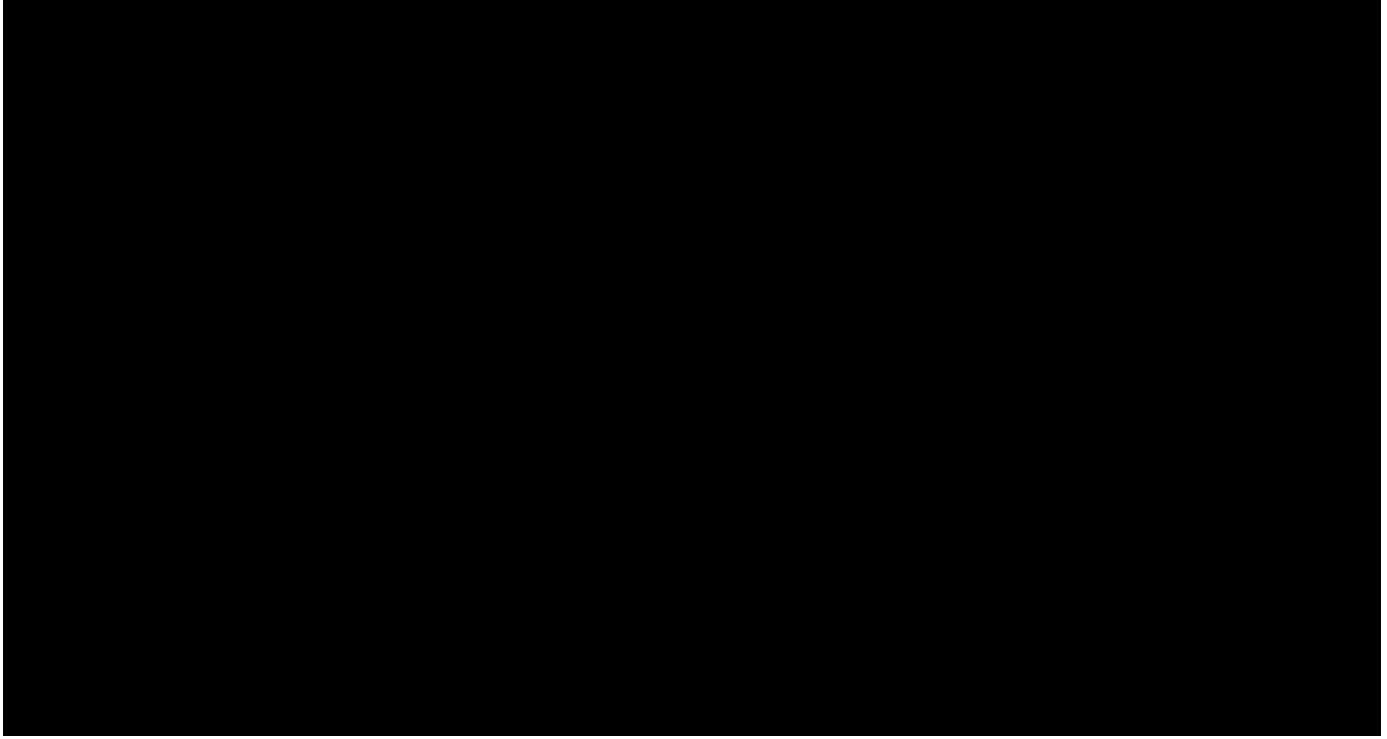
$$\text{For } n = 4: \lambda = R^{-1} \left(1 - \frac{1}{4^2} \right)^{-1} = 1.972018 \times 10^{-7} \text{ m} = 97.2018 \text{ nm (UV)}$$



- 4-14 (a) $r_n = \frac{n^2 \hbar^2}{m_e k e^2}$; where $n = 1, 2, 3, \dots$
- $$r_n = n^2 \frac{(1.055 \times 10^{-34} \text{ Js})^2}{(9.11 \times 10^{-31} \text{ kg})(9.0 \times 10^9 \text{ Nm}^2/\text{C}^2)(1.6 \times 10^{-19} \text{ C})^2} = (0.0529 \text{ nm})n^2$$
- For $n = 1$: $r_n = 0.0529 \text{ nm}$
 For $n = 2$: $r_n = 0.2121 \text{ nm}$
 For $n = 3$: $r_n = 0.4772 \text{ nm}$

- (b) From Equation 4.26, $v = \left(\frac{k e^2}{m_e r} \right)^{1/2}$
- $$v_1 = \left[\frac{(8.99 \times 10^9 \text{ Nm}^2/\text{C}^2)(1.60 \times 10^{-19} \text{ C})^2}{(9.11 \times 10^{-31} \text{ kg})(0.0529 \times 10^{-9} \text{ m})} \right]^{1/2} = 2.19 \times 10^6 \text{ m/s}$$
- $$v_2 = 1.09 \times 10^6 \text{ m/s}$$
- $$v_3 = 7.28 \times 10^5 \text{ m/s}$$

(c) As $c = 3.0 \times 10^8$ m/s, $v \ll c$ and no relativistic correction is necessary.



4-16 For Li²⁺, $Z = 3$ from Equation 4.36

$$\begin{array}{c} 0 \\ \hline \hline E_3 = -13.6 \text{ eV} \end{array}$$

$$E_n = -\frac{13.6Z^2}{n^2 \text{ eV}} = -\frac{122.4}{n^2 \text{ eV}}$$

$$\begin{array}{c} \hline \hline E_2 = -30.6 \text{ eV} \end{array}$$

So $E_1 = -122.4$ eV

$E_2 = -30.6$ eV

$E_3 = -13.6$ eV, etc.

$$\begin{array}{c} \hline \hline E_1 = -122.4 \text{ eV} \end{array}$$

4-17 $r = \frac{n^2 \hbar^2}{Z m_e k e^2} = \left(\frac{n^2}{Z} \right) \left(\frac{\hbar^2}{m_e k e^2} \right); n = 1$

$$r = \frac{1}{Z} \left[\frac{(1.055 \times 10^{-34} \text{ Js})^2}{(9.11 \times 10^{-31} \text{ kg})(9 \times 10^9 \text{ Nm}^2/\text{C}^2)(1.6 \times 10^{-19} \text{ C})^2} \right] = \frac{5.30 \times 10^{-11} \text{ m}}{Z}$$

(a) For He^+ , $Z = 2$, $r = \frac{5.30 \times 10^{-11} \text{ m}}{2} = 2.65 \times 10^{-11} \text{ m} = 0.0265 \text{ nm}$

(b) For Li^{2+} , $Z = 3$, $r = \frac{5.30 \times 10^{-11} \text{ m}}{3} = 1.77 \times 10^{-11} \text{ m} = 0.0177 \text{ nm}$

(c) For Be^{3+} , $Z = 4$, $r = \frac{5.30 \times 10^{-11} \text{ m}}{4} = 1.32 \times 10^{-11} \text{ m} = 0.0132 \text{ nm}$

4-18 (a) $\Delta E = (13.6 \text{ eV}) \left(\frac{1}{1^2} - \frac{1}{3^2} \right) = 12.1 \text{ eV}$

(b) Either $\Delta E = 12.1 \text{ eV}$ or $\Delta E = (13.6 \text{ eV}) \left(\frac{1}{1} - \frac{1}{2^2} \right) = 10.2 \text{ eV}$ and
 $\Delta E = (13.6 \text{ eV}) \left(\frac{1}{2^2} - \frac{1}{3^2} \right) = 1.89 \text{ eV}.$

4-19 (a) $\Delta E = (-13.6 \text{ eV}) \left(\frac{1}{n_i^2} - \frac{1}{n_f^2} \right) = (-13.6 \text{ eV}) \left(\frac{1}{9} - \frac{1}{4} \right) = 1.89 \text{ eV}$

(b) $E = \frac{hc}{\lambda} \Rightarrow \lambda = \frac{hc}{E} = (4.14 \times 10^{-15} \text{ eV s}) \frac{3 \times 10^8 \text{ m/s}}{1.89 \text{ eV}} = 658 \text{ nm}$

(c) $c = \lambda f \Rightarrow f = \frac{c}{\lambda} = \frac{3 \times 10^8 \text{ m/s}}{657 \times 10^{-9} \text{ m}} = 4.56 \times 10^{14} \text{ Hz}$

4-22 $E = K + U = \frac{mv^2}{2} - \frac{ke^2}{r}$. But $\frac{mv^2}{2} = \left(\frac{1}{2}\right) \frac{ke^2}{r}$. Thus $E = \left(\frac{1}{2}\right) \left(\frac{-ke^2}{r}\right) = \frac{U}{2}$, so $U = 2E = 2(-13.6 \text{ eV}) = -27.2 \text{ eV}$ and $K = E - U = -13.6 \text{ eV} - (-27.2 \text{ eV}) = 13.6 \text{ eV}$.

4-23 (a) $r_1 = (0.0529 \text{ nm})n^2 = 0.0529 \text{ nm}$ (when $n = 1$)

(b) $m_e v = m_e \left(\frac{ke^2}{m_e r} \right)^{1/2}$
 $m_e = \left[\frac{(9.1 \times 10^{-31} \text{ kg})(9 \times 10^9 \text{ Nm}^2/\text{C}^2)}{5.29 \times 10^{-11} \text{ m}} \right]^{1/2} \times (1.6 \times 10^{-19} \text{ C})$
 $M_e v = 1.99 \times 10^{-24} \text{ kg m/s}$

(c) $L = m_e v r = (1.99 \times 10^{-24} \text{ kg m/s})(5.29 \times 10^{-11} \text{ m})$, $L = 1.05 \times 10^{-34} (\text{kg m}^2/\text{s}) = \hbar$

(d) $K = |E| = 13.6 \text{ eV}$

(e) $U = -2K = -27.2 \text{ eV}$

(f) $E = K + U = -13.6 \text{ eV}$

4-25 (a) $\Delta E = hf = (13.6 \text{ eV}) \left(\frac{1}{n_f^2} - \frac{1}{n_i^2} \right)$ or $f = (13.6 \text{ eV}) \left(\frac{\frac{1}{9} - \frac{1}{16}}{4.14 \times 10^{-15} \text{ eV s}} \right) = 1.60 \times 10^{14} \text{ Hz}$

(b) $T = \frac{2\pi r_n}{v}$ so $f_{\text{rev}} = \frac{1}{T} = \frac{v}{2\pi r_n}$. Using $v = \left(\frac{ke^2}{m_e r_n} \right)^{1/2}$, $f_{\text{rev}} = \left(\frac{ke^2}{mr_n} \right)^{1/2}$. For $n = 3$, $r_3 = (3)^2 a_0$ and

$$f_{\text{rev}} = \frac{(8.99 \times 10^9 \text{ Nm}^2/\text{C}^2)(1.60 \times 10^{-19} \text{ C})^2}{[(9.11 \times 10^{-31} \text{ kg})(9)(5.29 \times 10^{-11} \text{ m})]^{1/2}}$$

$$f_{\text{rev}} = 2.44 \times 10^{14} \text{ Hz } (n = 3)$$

$$f_{\text{rev}} = 1.03 \times 10^{14} \text{ Hz } (n = 4)$$

Thus the photon frequency is about halfway between the two frequencies of the revolution.

4-32 (a)

$$\begin{aligned}\mu_{1H} &= \frac{m_e M}{m_e + M} = \frac{(9.109\ 390 \times 10^{-31} \text{ kg})(1.672\ 63 \times 10^{-27} \text{ kg})}{(9.109\ 390 \times 10^{-31} \text{ kg}) + (1.672\ 63 \times 10^{-27} \text{ kg})} \\ &= \frac{(9.109\ 390 \times 10^{-31} \text{ kg})(1.672\ 63 \times 10^{-27} \text{ kg})}{(0.000\ 910\ 939\ 0 \times 10^{-27} \text{ kg}) + (1.672\ 63 \times 10^{-27} \text{ kg})} = 9.104\ 431 \times 10^{-31} \text{ kg} \\ \frac{1}{\lambda} &= \frac{\mu}{m_e}(k) \left(\frac{1}{n_f^2} - \frac{1}{n_i^2} \right) = \left(\frac{9.104\ 431\ 6 \times 10^{-31} \text{ kg}}{9.109\ 390 \times 10^{-31} \text{ kg}} \right) (1.097\ 315\ 3 \times 10^7 \text{ m}^{-1}) \left(\frac{1}{2^2} - \frac{1}{3^2} \right) \\ \lambda_{1H} &= 656.469\ 1 \text{ nm}\end{aligned}$$

(b) Similarly, we find $\lambda_{2H} = 656.292\ 5 \text{ nm}$.

(c) $\lambda_{3H} = 656.232\ 5 \text{ nm}$

$$4-36 \quad (a) \quad f_{\text{revolution}} = \frac{v}{2\pi r} = \frac{\frac{n\hbar}{m_e r}}{2\pi r} = \frac{n\hbar}{2\pi m_e r^2} \text{ as } r^2 = n^4 (a_0)^2,$$

$$f_{\text{revolution}} = \frac{\hbar}{2\pi m_e} (a_0)^2 n^3 = \frac{6.58 \times 10^{15}}{n^3 \text{ Hz}}. \text{ Thus:}$$

$$\text{for } n = 100, f_{\text{revolution}} = 6.62 \times 10^9 \text{ Hz}$$

$$\text{for } n = 1000, f_{\text{revolution}} = 6.62 \times 10^6 \text{ Hz}$$

$$\text{for } n = 10000, f_{\text{revolution}} = 6.62 \times 10^3 \text{ Hz}$$

$$\text{using } r_n = n^2 a_0 = n^2 (0.529 \times 10^{-10} \text{ m}),$$

$$r_{100} = 0.529 \times 10^{-6} \text{ m}$$

$$r_{1000} = 0.529 \times 10^{-4} \text{ m}$$

$$r_{10000} = 0.529 \times 10^{-2} \text{ m} \equiv \frac{1}{2} \text{ cm}$$

$$(b) \quad f_{\text{photon}} = \frac{\Delta E}{h} = \frac{13.6Z^2}{h} \left(\frac{1}{(n-1)^2} - \frac{1}{n^2} \right) = \left(\frac{(13.6)(1)^2 \text{ eV}}{4.14 \times 10^{-15} \text{ eV s}} \right) \left(\frac{1}{(n-1)^2} - \frac{1}{n^2} \right)$$

$$= 3.285 \times 10^{15} \left(\frac{1}{(n-1)^2} - \frac{1}{n^2} \right) \text{ Hz}$$

for $n = 100$, $f_{\text{photon}} = 6.5798 \times 10^9 \text{ Hz}$

for $n = 1000$, $f_{\text{photon}} = 6.5798 \times 10^6 \text{ Hz}$

for $n = 10000$, $f_{\text{photon}} = 6.5798 \times 10^3 \text{ Hz}$

Thus, the difference in frequency, $\Delta f = f_{\text{revolution}} - f_{\text{photon}}$, are

$$n = 100, \Delta f = -0.05 \times 10^9 \text{ Hz}$$

$$n = 1000, \Delta f = -0.04 \times 10^6 \text{ Hz}$$

$$n = 10000, \Delta f = -0.05 \times 10^3 \text{ Hz}$$

- (c) These results show that f_{photon} tends to $f_{\text{revolution}}$ in the limit of large quantum numbers ($n = 10000$) and macroscopic sizes ($r \sim \frac{1}{2} \text{ cm}$).

$$4-37 \quad hf = \Delta E = \frac{4\pi^2 m_e k^2 e^4}{2h^2} \left(\frac{1}{(n-1)^2} - \frac{1}{n^2} \right), \quad f = \left(\frac{2\pi^2 m_e k^2 e^4}{h^3} \right) \left(\frac{2n-1}{(n-1)^2(n^2)} \right) \text{ as } n \rightarrow \infty,$$

$$f \rightarrow \left(\frac{2\pi^2 m_e k^2 e^4}{h^3} \right) \left(\frac{2}{n^3} \right). \text{ The revolution frequency is } f = \frac{v}{2\pi r} = \left(\frac{1}{2\pi} \right) \left(\frac{ke^2}{m_e} \right)^{1/2} \left(\frac{1}{r^{3/2}} \right) \text{ where}$$

$$r = \frac{n^2 h^2}{4\pi^2 m_e k e^2} \text{ substituting for } r, \quad f = \left(\frac{1}{2\pi} \right) \left(\frac{ke^2}{m_e} \right)^{1/2} \left(\frac{8\pi^3 m_e k e^3 (m_e k)^{1/2}}{n^3 h^3} \right) = \frac{4\pi^2 m_e k^2 e^4}{h^3 n^3}.$$