5 Matter Waves

5-2 The issue is: Can we use the simpler <u>classical</u> expression $p = (2mK)^{1/2}$ instead of the exact relativistic expression $p = \frac{K \left(1 + \frac{2mc^2}{K} \right)^{1/2}}{c}$? As the relativistic expression reduces to $p = (2mK)^{1/2}$ for $K \ll 2mc^2$, we can use the classical expression whenever $K \ll 1$ MeV because mc^2 for the electron is 0.511 MeV.

(a) Here 50 eV < 1 MeV, so
$$
p = (2mK)^{1/2}
$$

$$
\lambda = \frac{h}{p} = \frac{h}{\left[(2)\left(\frac{0.511 \text{ MeV}}{c^2}\right) (50 \text{ eV}) \right]^{1/2}} = \frac{hc}{\left[(2)(0.511 \text{ MeV})(50 \text{ eV}) \right]^{1/2}}
$$

$$
= \frac{1240 \text{ eV nm}}{\left[(2)(0.511 \times 10^6)(50)(\text{eV})^2 \right]^{1/2}} = 0.173 \text{ nm}
$$

(b) As $50 \text{ eV} \ll 1 \text{ MeV}$, $p = (2mK)^{1/2}$

$$
\lambda = \frac{hc}{\left[(2) \left(\frac{0.511 \text{ MeV}}{c^2} \right) \left(50 \times 10^3 \text{ eV} \right) \right]^{1/2}} = 5.49 \times 10^{-3} \text{ nm}.
$$

As this is clearly a worse approximation than in (a) to be on the safe side use the relativistic expression for p : $p = K \frac{\left(1 + \frac{2mc^2}{K}\right)^{1/2}}{E}$ $p = K \frac{(k + K) \cdot k}{c}$ so

$$
\lambda = \frac{h}{p} = \frac{hc}{\left(K^2 + 2Kmc^2\right)^{1/2}} = \frac{1240 \text{ eV nm}}{\left[\left(50 \times 10^3\right)^2 + (2)\left(50 \times 10^3\right)\left(0.511 \times 10^6 \text{ eV}\right)\right]^{1/2}}
$$

= 5.36 × 10⁻³ nm = 0.005 36 nm

5-7 A 10 MeV proton has $K = 10$ MeV $<< 2mc^2 = 1877$ MeV so we can use the classical expression $p = (2mK)^{1/2}$. (See Problem 5-2)

$$
\lambda = \frac{h}{p} = \frac{hc}{[(2)(938.3 \text{ MeV})(10 \text{ MeV})]^{1/2}} = \frac{1240 \text{ MeV fm}}{[(2)(938.3)(10)(\text{MeV})^2]^{1/2}} = 9.05 \text{ fm} = 9.05 \times 10^{-15} \text{ m}
$$

5-8
$$
\lambda = \frac{h}{p} = \frac{h}{(2mK)^{1/2}} = \frac{h}{(2meV)^{1/2}} = \left[\frac{h}{(2me)^{1/2}}\right]V^{-1/2}
$$

$$
\lambda = \left[\frac{6.626 \times 10^{-34} \text{ Js}}{(2 \times 9.105 \times 10^{-31} \text{ kg} \times 1.602 \times 10^{-19} \text{ C})^{1/2}}\right]V^{-1/2}
$$

$$
\lambda = \left[\frac{1.226 \times 10^{-9} \text{ kg}^{1/2} \text{m}^2}{sC^{1/2}}\right]V^{-1/2}
$$

5-10 As $\lambda = 2a_0 = 2(0.0529)$ nm = 0.105 8 nm the energy of the electron is nonrelativistic, so we can use

$$
p = \frac{h}{\lambda} \text{ with } K = \frac{p^2}{2m};
$$

\n
$$
K = \frac{h^2}{2m\lambda^2} = \frac{(6.626 \times 10^{-34} \text{ J} \cdot \text{s})^2}{2(9.11 \times 10^{-31} \text{ kg})(1.058 \times 10^{-10} \text{ m})^2} = 21.5 \times 10^{-18} \text{ J} = 134 \text{ eV}
$$

This is about ten times as large as the ground-state energy of hydrogen, which is 13.6 eV.

5-11 (a) In this problem, the electron must be treated relativistically because we must use relativity when $pc \approx mc^2$. (See problem 5-5). the momentum of the electron is

$$
p = \frac{h}{\lambda} = \frac{6.626 \times 10^{-34} \text{ J} \cdot \text{s}}{10^{-14} \text{ m}} = 6.626 \times 10^{-20} \text{ kg} \cdot \text{m/s}
$$

and
$$
pc = 124 \text{ MeV} \gg mc^2 = 0.511 \text{ MeV}
$$
. The energy of the electron is

$$
E = (p^2c^2 + m^2c^4)^{1/2}
$$

=
$$
\left[(6.626 \times 10^{-20} \text{ kg} \cdot \text{m/s})^2 (3 \times 10^8 \text{ m/s})^2 + (0.511 \times 10^6 \text{ eV})^2 (1.602 \times 10^{-19} \text{ J/eV})^2 \right]^{1/2}
$$

= 1.99×10^{-11} J = 1.24×10^8 eV

so that $K = E - mc^2 \approx 124 \text{ MeV}$.

(b) The kinetic energy is too large to expect that the electron could be confined to a region the size of the nucleus.

5-12 Using
$$
p = \frac{h}{\lambda} = mv
$$
, we find that $v = \frac{h}{m\lambda} = \frac{6.626 \times 10^{-34} \text{ J} \cdot \text{s}}{(9.11 \times 10^{-31} \text{ kg})(1 \times 10^{-10} \text{ m})} = 7.27 \times 10^6 \text{ m/s}$. From

the principle of conservation of energy, we get

$$
eV = \frac{mv^2}{2} = \frac{(9.11 \times 10^{-31} \text{ kg})(7.27 \times 10^6 \text{ m/s})^2}{2} = 2.41 \times 10^{-17} \text{ J} = 151 \text{ eV}.
$$

Therefore $V=151\,\mathrm{V}$.

5-15 For a free, non-relativistic electron $E = \frac{m_e c_0}{2}$ $e^{v_0^2} - p^2$ 2 $2 m_e$ $E = \frac{m_e v_0^2}{2} = \frac{p^2}{2m_e}$. As the wavenumber and angular frequency of the electron's de Broglie wave are given by $p = \hbar k$ and $E = \hbar \omega$, substituting these results gives the dispersion relation $\omega = \frac{\hbar k^2}{2}$ $2 m_{\rm e}$ *k* $rac{k^2}{m_e}$. So $v_g = \frac{d\omega}{dk} = \frac{\hbar k}{m_e} = \frac{p}{m_e} = v_0$.

5-17
$$
E^{2} = p^{2}c^{2} + (m_{e}c^{2})^{2}
$$

\n
$$
E = [p^{2}c^{2} + (m_{e}c^{2})^{2}]^{1/2}.
$$
 As $E = \hbar \omega$ and $p = \hbar k$
\n
$$
\hbar \omega = [\hbar^{2}k^{2}c^{2} + (m_{e}c^{2})^{2}]^{1/2} \text{ or}
$$

\n
$$
\omega(k) = \left[k^{2}c^{2} + \frac{(m_{e}c^{2})^{2}}{\hbar^{2}} \right]^{1/2}
$$

\n
$$
v_{p} = \frac{\omega}{k} = \frac{\left[k^{2}c^{2} + (m_{e}c^{2}/\hbar)^{2} \right]^{1/2}}{k} = \left[c^{2} + \left(\frac{m_{e}c^{2}}{\hbar k} \right)^{2} \right]^{1/2}
$$

\n
$$
v_{g} = \frac{d\omega}{dk} \Big|_{k_{0}} = \frac{1}{2} \left[k^{2}c^{2} + \left(\frac{m_{e}c^{2}}{\hbar} \right)^{2} \right]^{-1/2} 2kc^{2} = \frac{kc^{2}}{\left[k^{2}c^{2} + (m_{e}c^{2}/\hbar)^{2} \right]^{1/2}}
$$

\n
$$
v_{p}v_{g} = \begin{cases} \frac{\left[k^{2}c^{2} + (m_{e}c^{2}/\hbar)^{2} \right]^{1/2}}{k} \end{cases} \left\{ \left[k^{2}c^{2} + (m_{e}c^{2}/\hbar)^{2} \right]^{1/2} \right\} = c^{2}
$$

Therefore, $v_g < c$ if $v_p > c$.

5-23 (a)
$$
\Delta p \Delta x = m \Delta v \Delta x \ge \frac{\hbar}{2}
$$

$$
\Delta v \ge \frac{h}{4\pi m \Delta x} = \frac{2\pi \text{ J} \cdot \text{s}}{4\pi (2 \text{ kg})(1 \text{ m})} = 0.25 \text{ m/s}
$$

(b) The duck might move by $(0.25 \text{ m/s})(5 \text{ s}) = 1.25 \text{ m}$. With original position uncertainty of 1m, we can think of Δx growing to 1 m + 1.25 m = 2.25 m.

5-24 (a)
$$
\Delta x \Delta p = \hbar
$$
 so if $\Delta x = r$, $\Delta p \approx \frac{\hbar}{r}$

(b)
$$
K = \frac{p^2}{2m_e} \approx \frac{(\Delta p)^2}{2m_e} = \frac{\hbar^2}{2m_e r^2}
$$

$$
U = -\frac{ke^2}{r}
$$

$$
E = \frac{\hbar^2}{2m_e r^2} - \frac{ke^2}{r}
$$

(c) To minimize *E* take
$$
\frac{dE}{dr} = -\frac{\hbar^2}{m_e r^3} + \frac{ke^2}{r^2} = 0 \Rightarrow r = \frac{\hbar^2}{m_e ke^2} = \text{Bohr radius} = a_0
$$
. Then

$$
E = \left(\frac{\hbar}{2m_e}\right) \left(\frac{m_e ke^2}{\hbar^2}\right)^2 - ke^2 \left(\frac{m_e ke^2}{\hbar^2}\right) = \frac{m_e k^2 e^4}{2\hbar^2} = -13.6 \text{ eV}.
$$

5-26 The full width at half-maximum (FWHM) is 110 MeV. So ∆*E* = 55 MeV and using $\Delta E_{\text{min}} \Delta t_{\text{min}} = \frac{\hbar}{2}$ $E_{\min} \Delta t_{\min} = \frac{n}{2}$,

$$
\Delta t_{\text{min}} = \frac{\hbar}{2\Delta E} = \frac{6.58 \times 10^{-16} \text{ eV} \cdot \text{s}}{2(55 \times 10^6 \text{ eV})} \approx 6.0 \times 10^{-24} \text{ s}
$$

$$
\tau = \text{lifetime} \sim 2\Delta t_{\text{min}} = 1.2 \times 10^{-23} \text{ s}
$$

5-27 For a single slit with width a, minima are given by $\sin \theta = \frac{n\lambda}{a}$ where $n = 1, 2, 3, ...$ and $\sin \theta \approx \tan \theta = \frac{x}{L}$, $\frac{x_1}{L} = \frac{\lambda}{a}$ and $\frac{x_2}{L} = \frac{2\lambda}{a} \Rightarrow \frac{x_2 - x_1}{L} = \frac{\lambda}{a}$ $\frac{dz}{L} = \frac{2\pi}{a}$ \Rightarrow $\frac{x_2}{L}$ $\frac{x_1}{L} = \frac{\pi}{a}$ or

$$
\lambda = \frac{a\Delta x}{L} = \frac{5 \text{ Å} \times 2.1 \text{ cm}}{20 \text{ cm}} = 0.525 \text{ Å}
$$

$$
E = \frac{p^2}{2m} = \frac{h^2}{2m\lambda^2} = \frac{(hc)^2}{2mc^2\lambda^2} = \frac{(1.24 \times 10^4 \text{ eV} \cdot \text{Å})^2}{2(5.11 \times 10^5 \text{ eV})(0.525 \text{ Å})^2} = 546 \text{ eV}
$$

5-32 (a)
$$
f = \frac{E}{h} = \frac{(1.8)(1.6 \times 10^{-19} \text{ J})}{6.63 \times 10^{-34} \text{ J} \cdot \text{s}} = 4.34 \times 10^{14} \text{ Hz}
$$

$$
\lambda = \frac{c}{f} = 691 \text{ nm}
$$

(c)
$$
\Delta E \ge \frac{\hbar}{\Delta t} = \frac{6.63 \times 10^{-34} \text{ J} \cdot \text{s}}{2\pi (2 \times 10^{-6} \text{ s})}
$$

$$
\Delta E \ge 5.276 \times 10^{-29} \text{ J} = 3.30 \times 10^{-10} \text{ eV}
$$

5-34 (a)
$$
g(\omega) = (2\pi)^{-1/2} \int_{-\infty}^{\infty} V(t) (\cos \omega t - i \sin \omega t) dt
$$
, $V(t) \sin \omega t$ is an odd function so this integral vanishes leaving $g(\omega) = 2(2\pi)^{-1/2} \int_{0}^{t} V_0 \cos \omega t dt = \left(\frac{2}{\pi}\right)^{1/2} V_0 \frac{\sin \omega \tau}{\omega}$. A sketch of $g(\omega)$ is given below.

(b) As the major contribution to this pulse comes from ω 's between $-\frac{\pi}{\tau}$ and $\frac{\pi}{\tau}$, let

$$
\Delta \omega \approx \frac{\pi}{\tau}
$$
 and since $\Delta t = \tau$.

$$
\Delta \omega \Delta t = \left(\frac{\pi}{\tau}\right) \tau = \pi
$$

(c) Substituting $\Delta t = 0.5 \ \mu s$ in $\Delta \omega = \frac{\pi}{\Delta t}$ we find $\frac{\Delta 1}{2\Delta t} = \frac{1}{2(0.5 \times 10^{-6} \text{ s})} = 1 \times 10^6$ $\frac{1}{\Delta t} = \frac{1}{2(0.5 \times 10^{-6} \text{ s})} = 1 \times 10^6 \text{ Hz}$ $\frac{2\Delta t}{2\Delta t} = \frac{1}{2(0.5 \times 10^{-6} \text{ s})} = 1 \times 10^{6} \text{ Hz}$. As the range is 2∆*f*, the range is 2×10^6 Hz. For $\Delta t = 0.5$ ns, the range is

 $2\Delta f = 2 \times 10^9$ Hz.

5-35 (a)
$$
f(x) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{+\infty} a(k) e^{ikx} dk = \frac{A}{\sqrt{2\pi}} \int_{-\infty}^{+\infty} e^{-\alpha^2 (k-k_0)^2} e^{ikx} dk = \frac{A}{\sqrt{2\pi}} e^{-\alpha^2 k_0^2} \int_{-\infty}^{+\infty} e^{-\alpha^2 (k^2 - (2k_0 + ix/\alpha^2)k)} dk.
$$

Now complete the square in order to get the integral into the standard form +∞ [−] $\int_{-\infty}^{\infty} e^{-az^2} dz$:

$$
e^{-\alpha^2(k^2 - (2k_0 + ix/\alpha^2)k)} = e^{+\alpha^2(k_0 + ix/2\alpha^2)^2} e^{-\alpha^2(k - (k_0 + ix/2\alpha^2))^2}
$$

$$
f(x) = \frac{A}{\sqrt{2\pi}} e^{-\alpha^2 k_0^2} e^{\alpha^2(k_0 + ix/2\alpha^2)^2} \int_{k=-\infty}^{+\infty} e^{-\alpha^2(k - (k_0 + ix/2\alpha^2))^2} dk
$$

$$
= \frac{A}{\sqrt{2\pi}} e^{-x^2/4\alpha^2} e^{ik_0x} \int_{z=-\infty}^{+\infty} e^{-\alpha^2 z^2} dz
$$

where $z = k - \left(k_0 + \frac{ix}{2\alpha^2}\right)$. Since $\int_{z=-\infty}^{+\infty} e^{-\alpha^2 z^2} dz = \frac{\pi^4}{\alpha^4}$ +∞ [−] $\int_{z=-\infty}^{+\infty} e^{-\alpha^2 z^2} dz = \frac{\pi^{1/2}}{\alpha}$ $e^{-\alpha^2 z^2} dz = \frac{\pi^{\gamma - z}}{\alpha}, \ f(x) = \frac{A}{\alpha \sqrt{2}} e^{-x^2/4\alpha^2} e^{ik_0}$ 2 $f(x) = \frac{A}{\sqrt{2}} e^{-x^2/4\alpha^2} e^{ik_0x}$. The real part of *f*(*x*), Re *f*(*x*) is Re *f*(*x*) = $\frac{A}{\alpha\sqrt{2}}e^{-x^2 4\alpha^2} \cos k_0 x$ and is a gaussian envelope multiplying a harmonic wave with wave number k_0 . A plot of Re $f(x)$ is shown below:

Comparing $\frac{1}{\alpha\sqrt{2}}e^{-x^2/4\alpha}$ $-x^2 4\alpha^2$ 2 $\frac{A}{\sqrt{2}}e^{-x^2 4\alpha^2}$ to $Ae^{-(x/2\Delta x)^2}$ implies $\Delta x = \alpha$.

(c) By same reasoning because
$$
\alpha^2 = \frac{1}{4\Delta k^2}
$$
, $\Delta k = \frac{1}{2\alpha}$. Finally $\Delta x \Delta k = \alpha \left(\frac{1}{2\alpha}\right) = \frac{1}{2}$.

5-36 $E = K = \frac{1}{2} m u^2 = hf$ and $\lambda = \frac{h}{mu}$. $v_{\text{phase}} = f \lambda = \frac{m u^2}{2h} \frac{h}{m u} = \frac{u}{2} =$ $v_{\text{phase}} = f\lambda = \frac{mu^2}{2h} \frac{h}{mu} = \frac{u}{2} = v_{\text{phase}}$. This is different from the speed *u* at which the particle transports mass, energy, and momentum.