

$$\hat{M} = \frac{1}{2} \hat{N} = \# \text{ Cooper pairs} \Rightarrow \hat{M} \leftrightarrow \frac{1}{i} \frac{\partial}{\partial \alpha}$$

Project onto state of definite particle number $N = 2M$:

$$|M\rangle = \int_{-\pi}^{\pi} \frac{d\alpha}{2\pi} e^{-iM\alpha} |\alpha\rangle$$

Number fluctuations:

$$\frac{\langle \alpha | \hat{N}^2 | \alpha \rangle - \langle \alpha | \hat{N} | \alpha \rangle^2}{\langle \alpha | \hat{N} | \alpha \rangle} = \frac{2 \int d^3k \sin^2 \vartheta_k \cos^2 \vartheta_k}{\int d^3k \sin^2 \vartheta_k}$$

$$\text{Thus, } \Delta N_{\text{RMS}} \propto \sqrt{\langle N \rangle}$$

Lecture 18 (March 2)

Finite temperature: The finite temperature gap equation is

$$\Delta_k = - \sum_{k'} V_{k,k'} \frac{\Delta_{k'}}{2E_{k'}} \tanh\left(\frac{E_{k'}}{2k_B T}\right)$$

As $T \rightarrow \infty$, we see that $\Delta_k = 0$ is the only solution.

At what temperature does the gap collapse? We again take

$$V_{k,k'} = -\frac{v}{V} \Theta(\hbar\omega_D - |\xi_k|) \Theta(\hbar\omega_D - |\xi_{k'}|)$$

$$\Delta_k = \Delta \Theta(\hbar\omega_D - |\xi_k|)$$

and we obtain

$$1 = \frac{1}{2} g(\varepsilon_F) v \int_0^{\hbar\omega_D} d\xi (\xi^2 + \Delta^2)^{-1/2} \tanh\left(\frac{(\xi^2 + \Delta^2)^{1/2}}{k_B T}\right)$$

Setting $\Delta(T_c) = 0$, we have

$$\int_0^{\hbar\omega_D/2k_B T_c} ds s^{-1} \tanh s = \frac{2}{g(\epsilon_F) v}$$

This is an implicit equation for T_c . Assuming $k_B T_c \ll \hbar\omega_D$,

$$\int_0^{\hbar\omega_D/2k_B T_c} ds s^{-1} \tanh s = \ln\left(\frac{2e^C}{\pi} \cdot \frac{\hbar\omega_D}{k_B T_c}\right) + \mathcal{O}(e^{-\hbar\omega_D/2k_B T_c})$$

with $C = 0.57721566\dots$ the Euler-Mascheroni constant. So

$$k_B T_c = \underbrace{\frac{2e^C}{\pi}}_{1.134} \hbar\omega_D e^{-2/g(\epsilon_F) v}$$

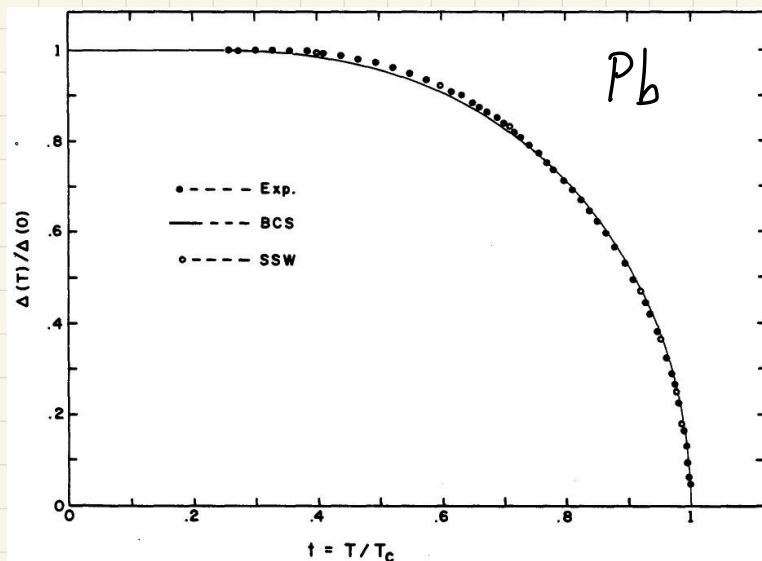
Recall how in this limit we also have

$$\Delta(T=0) = 2\hbar\omega_D e^{-2/g(\epsilon_F) v}$$

so combining these results we find the famous relation

$$2\Delta(0) = 2\pi e^{-C} k_B T_c = 3.52 k_B T_c$$

which relates the $T=0$ gap in the electronic spectrum to T_c .



Isotope effect: The logarithm of the T_c equation is

$$\ln T_c = \ln W_D - \frac{2}{g(\epsilon_F)v} + \text{const.}$$

Suppose we vary the mass of the ions via isotopic substitution. Then $W_D \propto M^{-1/2}$ changes, but $g(\epsilon_F)$ and v are largely insensitive. Then

$$\delta \ln T_c = \delta \ln W_D = -\frac{1}{2} \delta \ln M$$

where M is the ionic mass. This relation is fairly well-established among low- T_c materials.

Landau free energy of a superconductor: Let's derive an expression for the free energy of the superconductor as an expansion in the gap Δ . We start with

$$\hat{K}_{BCS} = \sum_{\mathbf{k}, \sigma} E_{\mathbf{k}} \gamma_{\mathbf{k}\sigma}^+ \gamma_{\mathbf{k}\sigma} + \sum_{\mathbf{k}} (\tilde{\xi}_{\mathbf{k}} - E_{\mathbf{k}}) - \sum_{\mathbf{k}, \mathbf{k}'} V_{\mathbf{k}, \mathbf{k}'} \langle C_{\mathbf{k}\uparrow}^+ C_{-\mathbf{k}\downarrow}^+ \rangle \langle C_{-\mathbf{k}\downarrow} C_{\mathbf{k}\uparrow} \rangle$$

and the relation

$$\langle C_{-\mathbf{k}-\sigma} C_{\mathbf{k}\sigma} \rangle = -\frac{\sigma \Delta_{\mathbf{k}}}{2E_{\mathbf{k}}} \tanh\left(\frac{1}{2}\beta E_{\mathbf{k}}\right)$$

which we derived previously. After invoking the gap equation,

$$\hat{K}_{BCS} = \sum_{\mathbf{k}, \sigma} E_{\mathbf{k}, \sigma} \gamma_{\mathbf{k}\sigma}^+ \gamma_{\mathbf{k}\sigma} + \sum_{\mathbf{k}} \left\{ \tilde{\xi}_{\mathbf{k}} - E_{\mathbf{k}} + \frac{|\Delta_{\mathbf{k}}|^2}{2E_{\mathbf{k}}} \tanh\left(\frac{E_{\mathbf{k}}}{2k_B T}\right) \right\}$$

Now compute $\Omega_S = -k_B T \ln \text{Tr} \exp(-\hat{K}_{BCS}/k_B T)$. We find

$$\Omega_S = -2k_B T \sum_{\mathbf{k}} \ln(1 + e^{-E_{\mathbf{k}}/k_B T}) + \sum_{\mathbf{k}} \left\{ \tilde{\xi}_{\mathbf{k}} - E_{\mathbf{k}} + \frac{|\Delta_{\mathbf{k}}|^2}{2E_{\mathbf{k}}} \tanh\left(\frac{E_{\mathbf{k}}}{2k_B T}\right) \right\}$$

The normal state free energy is obtained by setting $\Delta_{\mathbf{k}} \rightarrow 0$:

$$\Omega_n = -2k_B T \sum_{\mathbf{k}} \ln(1 + e^{-|\xi_{\mathbf{k}}|/k_B T}) + \sum_{\mathbf{k}} (\xi_{\mathbf{k}} - |\xi_{\mathbf{k}}|)$$

With $\Delta_{\mathbf{k}}(T) = \Delta(T) \Theta(\hbar\omega_D - |\xi_{\mathbf{k}}|)$, and assuming $\Delta(T) \ll \hbar\omega_D$, we obtain

$$\frac{\Omega_S - \Omega_n}{V} = -\frac{1}{4} g(\epsilon_F) \Delta^2 \left\{ 1 + 2 \ln\left(\frac{\Delta_0}{\Delta}\right) - \left(\frac{\Delta}{2\hbar\omega_D}\right)^2 + \mathcal{O}(\Delta^4) \right\} - 2g(\epsilon_F) k_B T \Delta \int_0^{\infty} ds \ln(1 + e^{-\sqrt{1+s^2}\Delta/k_B T}) + \frac{\pi^2}{6} g(\epsilon_F) (k_B T)^2 + \dots$$

where $\Delta_0 = \Delta(T=0) = \pi e^{-C} k_B T_c$.

① The limit $T \rightarrow 0^+$: We have

$$\frac{\Omega_S - \Omega_n}{V} = -\frac{1}{4} g(\epsilon_F) \Delta^2 \left\{ 1 + 2 \ln\left(\frac{\Delta_0}{\Delta}\right) + \mathcal{O}(\Delta^2) \right\} - g(\epsilon_F) \sqrt{2\pi} (k_B T)^3 \Delta e^{-\Delta/k_B T} + \frac{\pi^2}{6} g(\epsilon_F) (k_B T)^2 + \dots$$

Extremize wrt Δ to obtain an equation for $\Delta(T)$ at low temperatures:

$$\Delta(T) = \Delta_0 - \sqrt{2\pi} k_B T \Delta_0 e^{-\Delta_0/k_B T} + \dots$$

Substituting this result into the formula for the free energy difference, we find

$$\frac{\Omega_S - \Omega_n}{V} = -\frac{1}{4} g(\epsilon_F) \Delta_0^2 + \frac{\pi^2}{6} g(\epsilon_F) (k_B T)^2 - g(\epsilon_F) \sqrt{2\pi} (k_B T)^3 \Delta_0 e^{-\Delta_0/k_B T} + \dots$$

With $\Delta_0 = \pi e^{-C} k_B T_c$, setting the above to the condensation energy $-H_c^2(T)/8\pi$ gives

$$H_c(T) = H_c(0) \left\{ \underbrace{1 - \frac{1}{3} e^{2C} \left(\frac{T}{T_c}\right)^2}_{1.057} + \dots \right\}$$

where $H_c(0) = \sqrt{2\pi g(\epsilon_F)} \Delta_0$.

② The limit $T \rightarrow T_c^-$: The analysis here is tricky; see § 12.14 of the lecture notes. One finds $\zeta(3) = 1.20205690\dots$

$$\begin{aligned} \frac{\Omega_s - \Omega_n}{V} &= \frac{1}{2} g(\epsilon_F) \ln\left(\frac{T}{T_c}\right) \Delta^2 + \frac{7\zeta(3)}{32\pi^2} \frac{g(\epsilon_F)}{(k_B T)^2} \Delta^4 + \mathcal{O}(\Delta^6) \\ &\equiv \tilde{a}(T) \Delta^2 + \frac{1}{2} \tilde{b}(T) \Delta^4 + \mathcal{O}(\Delta^6) \end{aligned}$$

with, working to lowest order in $T - T_c$,

$$\tilde{a}(T) = \frac{1}{2} g(\epsilon_F) \left(\frac{T}{T_c} - 1\right), \quad \tilde{b}(T_c) = \frac{7\zeta(3)}{32\pi^2} \frac{g(\epsilon_F)}{(k_B T_c)^2}$$

The heat capacity jump is then

$$C_s(T_c^-) - C_n(T_c^+) = \frac{T_c [\tilde{a}'(T_c)]^2}{\tilde{b}(T_c)} = \frac{4\pi^2}{7\zeta(3)} g(\epsilon_F) k_B^2 T_c$$

We then have

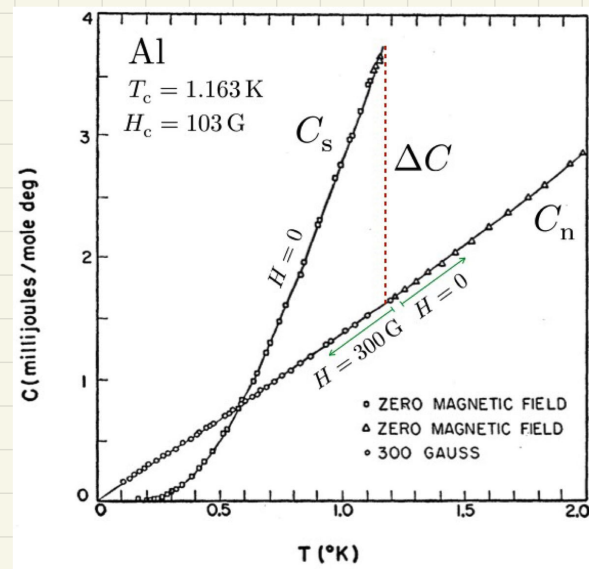
$$\frac{C_s(T_c^-) - C_n(T_c^+)}{C_n(T_c^+)} = \frac{12}{7\zeta(3)} = 1.43$$

The order parameter is accordingly given by

$$\Delta^2(T) = -\frac{\tilde{a}(T)}{\tilde{b}(T)} = \frac{8e^{2c}}{73(3)} \left(1 - \frac{T}{T_c}\right) \Delta_0^2$$

in which case

$$\frac{\Delta(T)}{\Delta(0)} = \underbrace{\left(\frac{8e^{2c}}{73(3)}\right)^{1/2}}_{1.734} \left(1 - \frac{T}{T_c}\right)^{1/2}$$



Just below T_c , the thermodynamic critical field H_c is given by the expression $H_c^2 = 4\pi \tilde{a}^2(T)/\tilde{b}(T_c)$, hence

$$\frac{H_c(T)}{H_c(0)} = 1.734 \left(1 - \frac{T}{T_c}\right)$$

Paramagnetic susceptibility: Add a weak magnetic field:

$$\hat{H}_1 = -\mu_B H \sum_{\mathbf{k}, \sigma} \sigma c_{\mathbf{k}\sigma}^\dagger c_{\mathbf{k}\sigma} = -\mu_B H \sum_{\mathbf{k}, \sigma} \sigma \gamma_{\mathbf{k}\sigma}^\dagger \gamma_{\mathbf{k}\sigma}$$

We compute the free energy shift,

$$\Delta\Omega_s(T, V, \mu, H) \equiv \Omega_s(T, V, \mu, H) - \Omega_s(T, V, \mu, 0)$$

$$= -k_B T \sum_{\mathbf{k}, \sigma} \ln \left(\frac{1 + e^{-(E_{\mathbf{k}} + \sigma \mu_B H)/k_B T}}{1 + e^{-E_{\mathbf{k}}/k_B T}} \right)$$

$$= -\frac{(\mu_B H)^2}{k_B T} \sum_{\mathbf{k}, \sigma} \frac{e^{E_{\mathbf{k}}/k_B T}}{(e^{E_{\mathbf{k}}/k_B T} + 1)^2} + \mathcal{O}(H^4)$$

The magnetic susceptibility is then

$$\chi_s = -\frac{1}{V} \left. \frac{\partial^2 \Delta \Omega_s}{\partial H^2} \right|_{H=0} = g(\epsilon_F) \mu_B^2 Y(T)$$

Yoshida function

$$Y(T) = 2 \int_0^{\infty} d\xi \left(-\frac{\partial f}{\partial E} \right) = \frac{1}{2k_B T} \int_0^{\infty} d\xi \operatorname{sech}^2 \left(\frac{\sqrt{\xi^2 + \Delta^2}}{2k_B T} \right)$$

Limits: $Y(T \rightarrow 0) = (2\pi\Delta/k_B T) e^{-\Delta/k_B T}$ and $Y(T_c) = 1$.

Since $\chi_n(T) = g(\epsilon_F) \mu_B^2 =$ Pauli susceptibility,

$$\frac{\chi_s(T)}{\chi_n(T)} = Y(T)$$

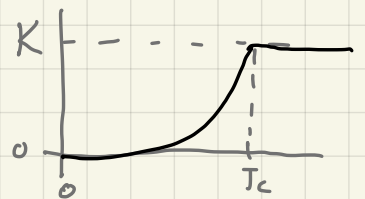
As $T \rightarrow 0$, $Y(T)$ is exponentially suppressed as $e^{-\Delta_0/k_B T}$.

The susceptibility vanishes exponentially because it requires a finite energy Δ_0 to create a Bogoliubov quasiparticle out of the spin singlet BCS ground state. In metals, nuclear spins

experience a shift in their resonance energies due to hyperfine coupling to the conduction electrons, called the **Knight shift**.

The formula for the Knight shift is

$$K = \frac{\omega - \omega_0}{\omega_0} = \frac{8\pi}{3} |\psi(0)|^2 \chi_{el}$$



where $\psi(0)$ is the amplitude of the electron wavefunction at the nucleus. In a superconductor, $K(T \rightarrow 0) \rightarrow 0$ exponentially.

Electrons remain unpolarized in a weak external field due to the finite energy gap.

Supercurrent: As we saw within Ginzburg-Landau theory, a spatially varying order parameter $\Psi(\vec{x}) = \Psi_0 e^{i\vec{q}\cdot\vec{x}}$, in the absence of external fields, corresponds to finite current density. The free energy density is $\leftarrow K |\nabla\Psi|^2$

$$f = a|\Psi_0|^2 + \frac{1}{2}b|\Psi_0|^4 + K\vec{q}^2|\Psi_0|^2$$

Extremizing wrt $|\Psi_0|^2$ gives $|\Psi_0|^2 = -(a + K\vec{q}^2)/b$ if $a + K\vec{q}^2 < 0$ and zero otherwise. If $a(T) = \alpha(T - T_c^0)$, we conclude $T_c(\vec{q}) = T_c(\vec{q}=0) - \alpha^{-1}K\vec{q}^2$. The current density is

$$\vec{j} = -\frac{2ke^*}{\hbar^2} \cdot \frac{\hbar}{2i} (\Psi^* \vec{\nabla}\Psi - \Psi \vec{\nabla}\Psi^*) = -\frac{2ke^*}{\hbar} \vec{q}$$

To describe a moving condensate within BCS theory, we must give finite momentum to the Cooper pairs. This means

$$\langle C_{-\vec{k}+\frac{1}{2}\vec{p}\downarrow} C_{\vec{k}+\frac{1}{2}\vec{p}\uparrow} \rangle = \Psi_{\vec{k},\vec{q}} \delta_{\vec{p},\vec{q}}$$

Then

$$\hat{K}_{BCS} = \sum_{\vec{k}} \begin{pmatrix} C_{\vec{k}+\frac{1}{2}\vec{q}\uparrow}^+ & C_{-\vec{k}+\frac{1}{2}\vec{q}\downarrow} \end{pmatrix} \begin{pmatrix} \xi_{\vec{k}+\frac{1}{2}\vec{q}} & \Delta_{\vec{k},\vec{q}} \\ \Delta_{\vec{k},\vec{q}}^* & \xi_{-\vec{k}+\frac{1}{2}\vec{q}} \end{pmatrix} \begin{pmatrix} C_{\vec{k}+\frac{1}{2}\vec{q}\uparrow} \\ C_{-\vec{k}+\frac{1}{2}\vec{q}\downarrow}^+ \end{pmatrix} + \sum_{\vec{k}} (\xi_{\vec{k}} - \Delta_{\vec{k},\vec{q}} \langle C_{\vec{k}+\frac{1}{2}\vec{q}\uparrow}^+ C_{-\vec{k}+\frac{1}{2}\vec{q}\downarrow}^+ \rangle)$$

The technical details are discussed in §12.11 of the lecture notes.

One finds

$$\Delta_{\vec{k},\vec{q}} = \Delta_{0,\vec{q}} \oplus (\hbar\omega_0 - |\xi_{\vec{k}}|)$$

The gap equation is then

$$\sinh^{-1}\left(\frac{\hbar\omega_D + \eta_{\vec{q}}}{\Delta_{0,\vec{q}}}\right) = \frac{2}{g(\epsilon_F)V} + \sinh^{-1}\left(\frac{\eta_{\vec{q}}}{\Delta_{0,\vec{q}}}\right)$$

with $\eta_{\vec{q}} \equiv \hbar^2 \vec{q}^2 / 8m^*$ in the case $\epsilon_{\vec{q}} = \hbar^2 \vec{q}^2 / 2m^*$. We determine the critical wavevector q_c where the gap collapses by taking $\Delta_{0,\vec{q}} \rightarrow 0$, resulting in

$$\frac{2}{g(\epsilon_F)V} = \ln\left(1 + \frac{\hbar\omega_D}{\eta_{q_c}}\right) \Rightarrow \eta_{q_c} \approx \hbar\omega_D e^{-2/g(\epsilon_F)V} = \frac{1}{2}\Delta_0$$

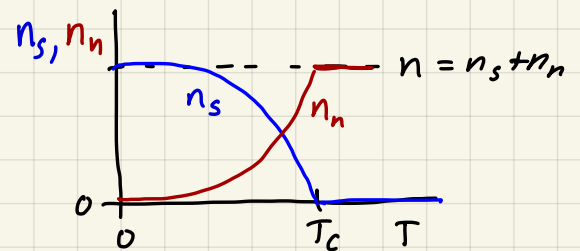
Assuming $\eta_{\vec{q}} \ll \Delta_0$, we have the gap $\Delta_{0,\vec{q}} \approx \Delta_0 - \frac{\hbar^2 \vec{q}^2}{8m^*}$.

Now for the super-current. We have

$$\vec{j} = \frac{n\epsilon\hbar}{2m^*} \vec{q} + \frac{2e\hbar}{m^*V} \sum_{\vec{k}} \vec{k} \langle C_{\vec{k} + \frac{1}{2}\vec{q}\uparrow}^\dagger C_{\vec{k} + \frac{1}{2}\vec{q}\uparrow} \rangle$$

where $n = N/V$. One obtains $\vec{j} = n_s(T) e \hbar \vec{q} / 2m^*$ where the superfluid density $n_s(T)$ is given by

$$\begin{aligned} n_s(T) &= n \left\{ 1 + 2 \int_0^\infty d\zeta \frac{\partial f}{\partial \epsilon} \right\} \\ &= n \{ 1 - Y(T) \} \end{aligned}$$



where $Y(T)$ is the Yosida function. Note then $n_n(T) = n Y(T)$.

Recall $Y(T_c) = 1$. Finally, the GL free energy density is

$$\frac{\Omega_s - \Omega_n}{V} = \tilde{a}(T) |\Delta|^2 + \frac{1}{2} \tilde{b}(T_c) |\Delta|^4 + \tilde{K} |\Delta|^2 \vec{q}^2$$

with

$$\tilde{K} = \frac{\hbar^2}{2m^*} \frac{n \tilde{b}(T_c)}{g(\epsilon_F)}$$

Effect of repulsive interactions: Let's now take

$$V_{\mathbf{k},\mathbf{k}'} = \begin{cases} (v_c - v_p)/V & \text{if } |\mathbf{k}| < \hbar\omega_D \text{ and } |\mathbf{k}'| < \hbar\omega_D \\ v_c/V & \text{otherwise} \end{cases}$$

Here v_p is due to phonon-mediated attraction and v_c due to Coulomb repulsion ($v_p, v_c > 0$). We posit a solution

$$\Delta_{\mathbf{k}} = \begin{cases} \Delta_0 & \text{if } |\mathbf{k}| < \hbar\omega_D \\ \Delta_1 & \text{if } |\mathbf{k}| \geq \hbar\omega_D \end{cases}$$

with both $\Delta_{0,1} \in \mathbb{R}$. We presume $v_p > v_c > 0$ so that attraction wins close to the Fermi surface, but as we shall see below, this is not absolutely necessary! The gap equation then yields

$$\Delta_0 = \frac{1}{2} g(\epsilon_F) (v_p - v_c) \int_0^{\hbar\omega_D} d\tilde{\zeta} \frac{\Delta_0}{\sqrt{\tilde{\zeta}^2 + \Delta_0^2}} - \frac{1}{2} g(\epsilon_F) v_c \int_0^B d\tilde{\zeta} \frac{\Delta_1}{\sqrt{\tilde{\zeta}^2 + \Delta_1^2}}$$

$$\Delta_1 = -\frac{1}{2} g(\epsilon_F) v_c \int_0^{\hbar\omega_D} d\tilde{\zeta} \frac{\Delta_0}{\sqrt{\tilde{\zeta}^2 + \Delta_0^2}} - \frac{1}{2} g(\epsilon_F) v_c \int_{\hbar\omega_D}^B d\tilde{\zeta} \frac{\Delta_1}{\sqrt{\tilde{\zeta}^2 + \Delta_1^2}}$$

where B is the electron bandwidth. Assuming $|\Delta_{0,1}| \ll \hbar\omega_D \ll B$ we obtain

$$\Delta_0 = \frac{1}{2} g(\epsilon_F) (v_p - v_c) \Delta_0 \ln\left(\frac{2\hbar\omega_D}{\Delta_0}\right) - \frac{1}{2} g(\epsilon_F) v_c \Delta_1 \ln\left(\frac{B}{\hbar\omega_D}\right)$$

$$\Delta_1 = -\frac{1}{2} g(\epsilon_F) v_c \Delta_0 \ln\left(\frac{2\hbar\omega_D}{\Delta_0}\right) - \frac{1}{2} g(\epsilon_F) v_c \Delta_1 \ln\left(\frac{B}{\hbar\omega_D}\right)$$

The second of these equations gives

$$\Delta_1 = - \frac{\frac{1}{2} g(\epsilon_F) v_c \ln(2\hbar\omega_D / \Delta_0)}{1 + \frac{1}{2} g(\epsilon_F) v_c \ln(B / \hbar\omega_D)} \Delta_0 < 0$$

Insert this into the first equation to find

$$\frac{2}{g(\epsilon_F) v_p} = \left\{ 1 - \frac{v_c}{v_p} \frac{1}{1 + \frac{1}{2} g(\epsilon_F) v_c \ln(B / \hbar\omega_D)} \right\} \cdot \ln \left(\frac{2\hbar\omega_D}{\Delta_0} \right)$$

which has a solution provided

$$v_p > \frac{v_c}{1 + \frac{1}{2} g(\epsilon_F) v_c \ln(B / \hbar\omega_D)}$$

The RHS reflects a renormalized value of the bare Coulomb repulsion v_c . Thus, we can have a superconducting solution even when

$$v_c > v_p > \frac{v_c}{1 + \frac{1}{2} g(\epsilon_F) v_c \ln(B / \hbar\omega_D)}$$

and the interactions are always repulsive!! Very surprising!

To find T_c , set $\Delta_{0,1} \rightarrow 0$ with $r \equiv \Delta_1 / \Delta_0$ finite:

$$\frac{2}{g(\epsilon_F)} = (v_p - v_c) \int_0^{\tilde{\Omega}} ds s^{-1} \tanh s - r v_c \int_{\tilde{\Omega}}^{\tilde{B}} ds s^{-1} \tanh s$$

$$\frac{2}{g(\epsilon_F)} = -r^{-1} v_c \int_0^{\tilde{\Omega}} ds s^{-1} \tanh s - v_c \int_{\tilde{\Omega}}^{\tilde{B}} ds s^{-1} \tanh s$$

where $\tilde{\Omega} \equiv \hbar\omega_D / 2k_B T_c$ and $\tilde{B} \equiv B / 2k_B T_c$. We obtain

$$\frac{2}{g(\epsilon_F)v_p} = \left\{ 1 - \frac{v_c}{v_p} \frac{1}{1 + \frac{1}{2}g(\epsilon_F)v_c \ln(B/\hbar\omega_D)} \right\} \cdot \ln\left(\frac{1.134\hbar\omega_D}{k_B T_c}\right)$$

Once again, $2\Delta_0(T=0) = 3.52 k_B T_c$, but with

$$k_B T_c = 1.134\hbar\omega_D e^{-2/g(\epsilon_F)v_{\text{eff}}}$$

and

$$v_{\text{eff}} = v_p - \frac{v_c}{1 + \frac{1}{2}g(\epsilon_F)v_c \ln(B/\hbar\omega_D)}$$

Standard notation:

$$\lambda \equiv \frac{1}{2}g(\epsilon_F)v_p, \quad \mu \equiv \frac{1}{2}g(\epsilon_F)v_c, \quad \mu^* \equiv \frac{\mu}{1 + \mu \ln(B/\hbar\omega_D)}$$

so that

$$k_B T_c = 1.134\hbar\omega_D e^{-1/(\lambda - \mu^*)}$$

and

$$\Delta_0 = 2\hbar\omega_D e^{-1/(\lambda - \mu^*)}, \quad \Delta_1 = -\frac{\mu^* \Delta_0}{\lambda - \mu^*}$$

Since μ^* depends on ω_D , the isotope effect is affected.